### New Approaches to Lunar Ice Detection and Mapping

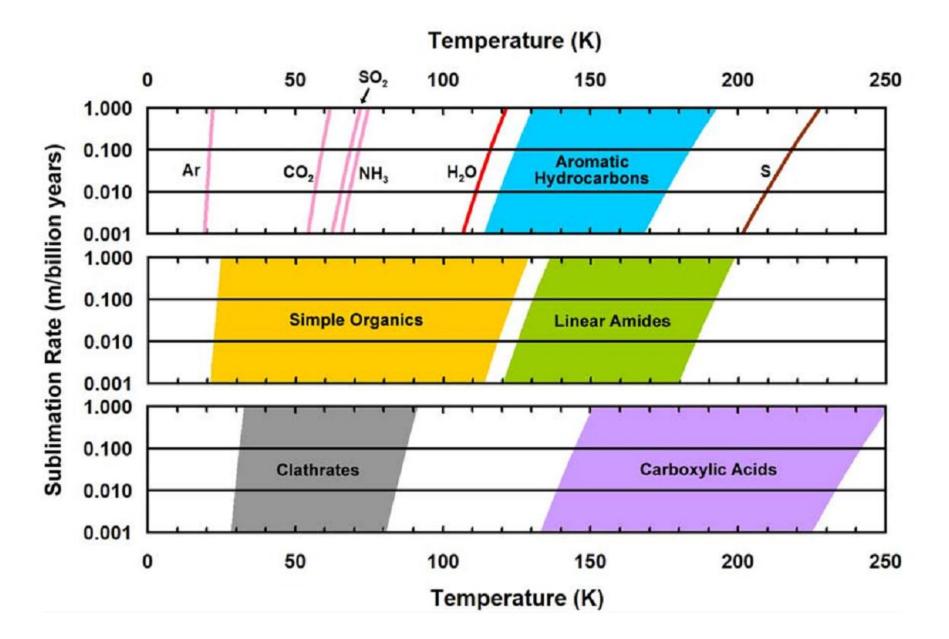
Keck Institute for Space Studies Caltech, Pasadena, CA July 22, 2013

### **Recent Results: Thermal**

David A. Paige Dept. of Earth, Planetary and Space Sciences UCLA

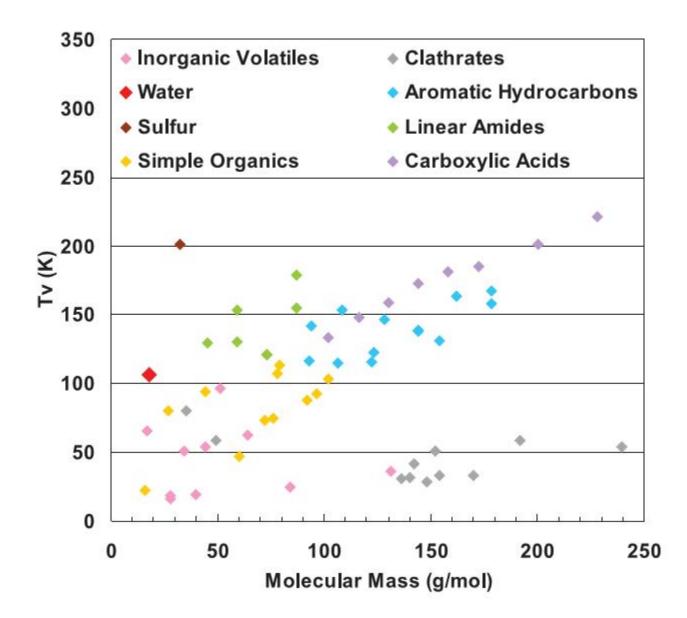
# Temperature - The Master Variable

- Thermal Stability of Volatiles
- Spatial Distribution of Cold Traps
- Thermal Evolution of Cold Traps
- Correlations with Other Datasets and Models
  - Radar
  - Neutrons
  - Surface Reflectance
  - In-Situ
  - Future....



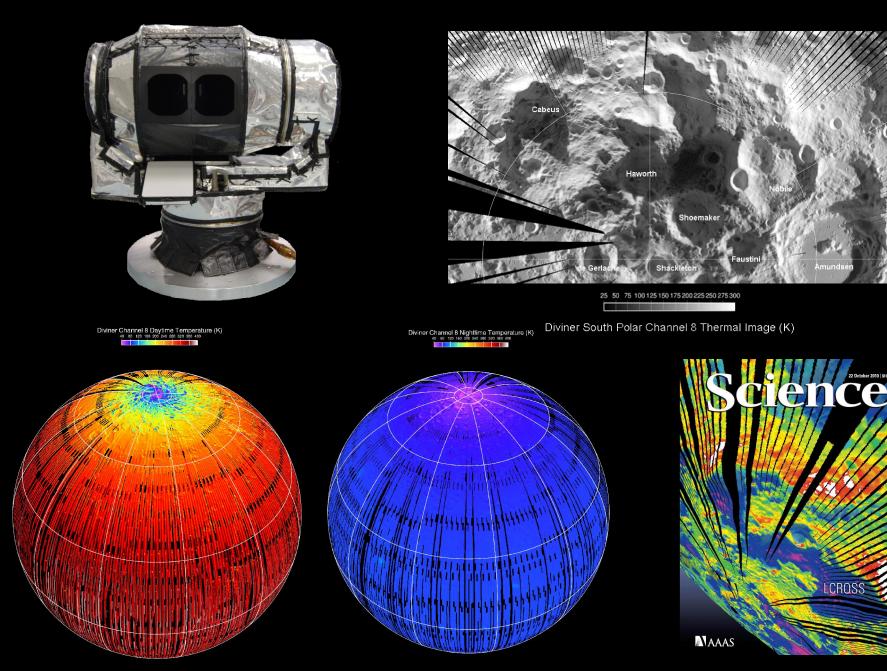
Zhang and Paige, Geophysical Research Letters 36, L16203, 2009.

	Chemical Formula	Name	M (g/mol)	Т <sub>v</sub> (K)		Chemical Formula	Name	M (g/mol)	T <sub>v</sub> (K)
	N <sub>2</sub>	Nitrogen	28.02	16.2	su	C <sub>7</sub> H <sub>6</sub> O	Benzaldehyde	106.13	114.6
Inorganics	СО	Carbon monoxide	28.01	18.2		C <sub>7</sub> H <sub>6</sub> O <sub>2</sub>	Salicylaldehyde	122.12	115.6
	Ar	Argon	39.95	19.5		C <sub>6</sub> H <sub>7</sub> N	Aniline	93.13	116.5
	Kr	Krypton	83.80	24.5	g	C <sub>6</sub> H <sub>5</sub> NO <sub>2</sub>	Nitrobenzene	123.11	122.8
	Xe	Xeon	131.30	36.1	Aromatic Hydrocarbons	C <sub>12</sub> H <sub>10</sub>	Biphenyl	154.20	131.3
	H <sub>2</sub> S	Hydrogen sulfide	34.09	50.6		C <sub>10</sub> H <sub>8</sub> O	1-Naphthol	144.16	137.7
	CO <sub>2</sub>	Carbon dioxide	44.01	54.3		C <sub>6</sub> H <sub>6</sub> O	Phenol	94.11	141.9
	NH <sub>3</sub>	Ammonia	17.03	65.5		C <sub>10</sub> H <sub>8</sub> O	2-Naphthol	144.16	139.1
	NH₄SH	Ammonium hydrosulfide	51.12	96.1		C <sub>10</sub> H <sub>8</sub>	Naphthalene	128.16	146.1
	SO <sub>2</sub>	Sulfur dioxide	64.07	62.3		C <sub>7</sub> H <sub>8</sub> O	Benzyl alcohol	108.13	153.3
	H <sub>2</sub> O	Water	18.02	106.6		C <sub>14</sub> H <sub>10</sub>	Phenanthrene	178.22	157.8
	S	Sulfur	32.07	201.5		C <sub>12</sub> H <sub>18</sub>	Hexamethylbenzene	162.26	163.2
Simple Organics	CH₄	Methane	16.04	22.0		C <sub>14</sub> H <sub>10</sub>	Anthracene	178.22	167.5
	OCS	Carbonyl sulfide	60.08	46.8	in the second	C <sub>3</sub> H <sub>7</sub> NO	Dimethylformamide	70.10	120.7
	C <sub>5</sub> H <sub>12</sub>	Pentane	72.15	73.6	Linear Amides	C <sub>3</sub> H <sub>7</sub> NO	Methylacetamide	73.10	120.8
	CS <sub>2</sub>	Carbon disulfide	76.15	74.4		C <sub>2</sub> H <sub>5</sub> NO	Formamide	45.04	129.8
	HCN	Hydrogen cyanide	27.03	80.5		CH <sub>3</sub> NO	Methylformamide	59.07	130.1
	C <sub>7</sub> H <sub>8</sub>	Toluene	92.13	87.6		C <sub>2</sub> H <sub>5</sub> NO	Acetamide	59.07	153.3
	C <sub>5</sub> H <sub>10</sub> O	3-Pentanone	96.21	92.8		C <sub>4</sub> H <sub>9</sub> NO	Dimethylacetamide	87.12	154.9
	NH₄CN	Ammonium cyanide	44.06	93.8		C₄H <sub>9</sub> NO	Methylpropanamide	87.12	179.1
	C <sub>5</sub> H <sub>10</sub> O <sub>2</sub>	Ethyl propanoate	102.13	103.6	Carboxylic Acids	C <sub>5</sub> H <sub>10</sub> O <sub>2</sub>	Valeric acid	102.13	133.6
	NH <sub>4</sub> CO <sub>2</sub> NH <sub>2</sub>	Ammonium carbonate	78.08	107.4		C <sub>6</sub> H <sub>12</sub> O <sub>2</sub>	Caproic acid	116.16	148.4
	NH <sub>4</sub> HCO <sub>3</sub>	Ammonium bicarbonate	79.06	113.3		C <sub>7</sub> H <sub>14</sub> O <sub>2</sub>	Enanthic acid	130.18	159.1
Clathrates	Ar·6H <sub>2</sub> O	Argon clathrate	148.05	28.9		C <sub>8</sub> H <sub>16</sub> O <sub>2</sub>	Caprylic acid	144.23	172.4
	N <sub>2</sub> ·6H <sub>2</sub> O	Nitrogen clathrate	136.12	30.5		C <sub>9</sub> H <sub>18</sub> O <sub>2</sub>	Pelargonic acid	158.23	181.1
	O <sub>2</sub> ·6H <sub>2</sub> O	Oxygen clathrate	140.10	31.9		C <sub>10</sub> H <sub>20</sub> O <sub>2</sub>	Capric acid	172.26	185.4
	CO <sub>2</sub> ·7H <sub>2</sub> O	Carbon dioxide clathrate	170.12	33.4		C <sub>12</sub> H <sub>24</sub> O <sub>2</sub>	Lauric acid	200.31	201.3
	N <sub>2</sub> ·7H <sub>2</sub> O	Nitrogen clathrate	154.13	33.4		C <sub>16</sub> H <sub>32</sub> O <sub>2</sub>	Palmitic acid	228.36	221.4
	CH <sub>4</sub> ·7H <sub>2</sub> O	Methane clathrate	142.15	41.8		C <sub>18</sub> H <sub>36</sub> O <sub>2</sub>	Stearic acid	256.42	224.3
	CO <sub>2</sub> ·6H <sub>2</sub> O	Carbon dioxide clathrate	152.12	50.9	Fullerene		Fullerene	720.60	465.0
	Xe·6H <sub>2</sub> O	Xeon clathrate	239.40	53.9					
	2NH <sub>3</sub> ·H <sub>2</sub> O	Ammonia clathrate	191.90	58.4		C <sub>60</sub>			
	Kr·6H₂O	Krypton clathrate	49.06	58.6					
	NH <sub>3</sub> ·H <sub>2</sub> O	Ammonia clathrate	35.05	80.3					



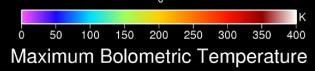
Zhang and Paige, Geophysical Research Letters 36, L16203, 2009.

#### LRO Diviner Lunar Radiometer Experiment



## Diviner Observed Maximum Temperature North Pole South Pole

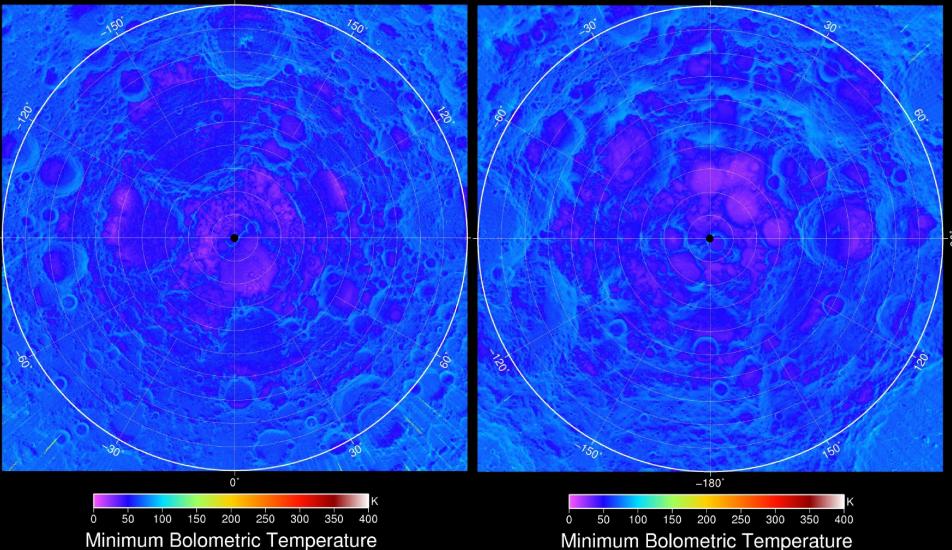
–180° 0 –180



o 50 100 150 200 250 300 350 400 Maximum Bolometric Temperature

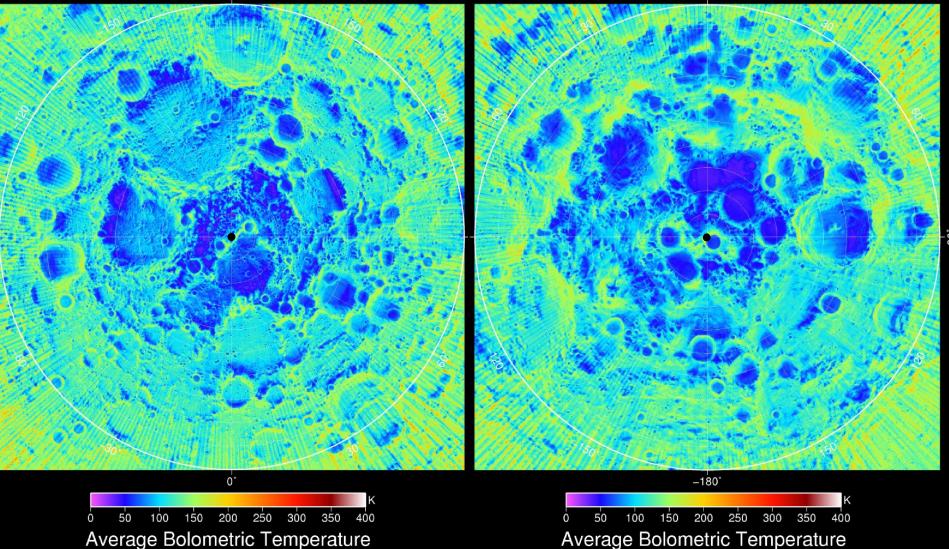
## Diviner Observed Minimum Temperature North Pole South Pole

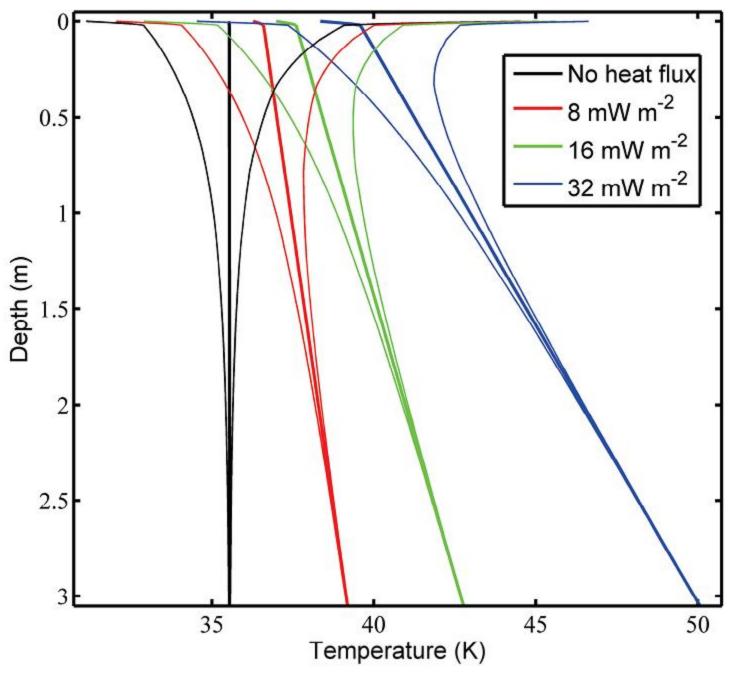
–180°



## Diviner Observed Average Temperature North Pole South Pole

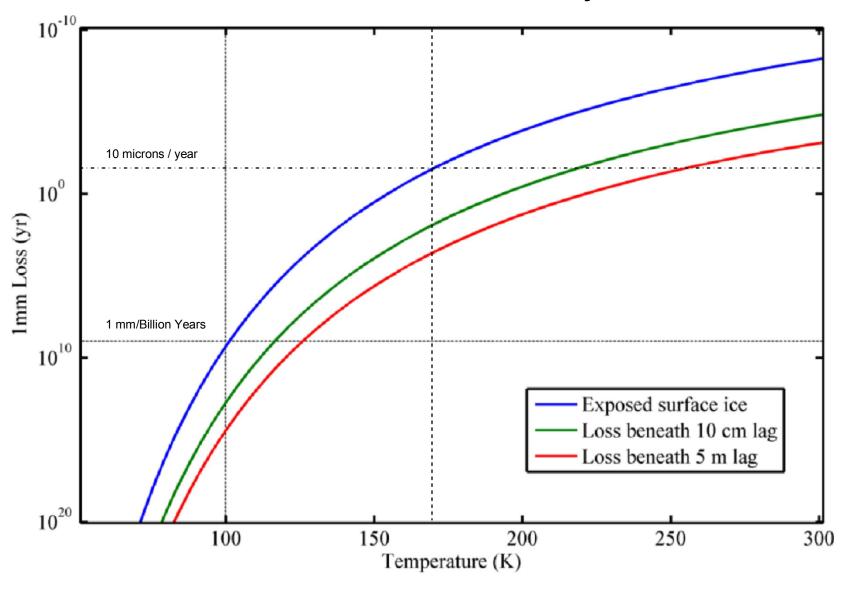




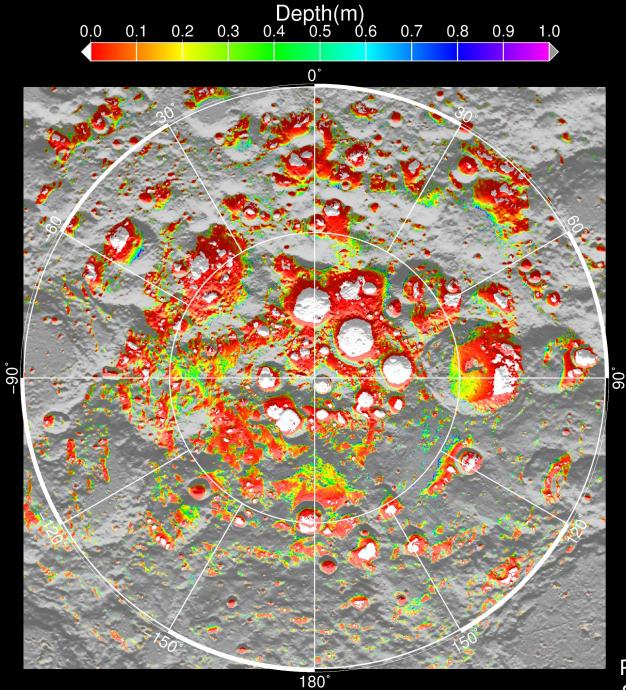


Paige et al., Science, 2010

Water Ice Thermal Stability



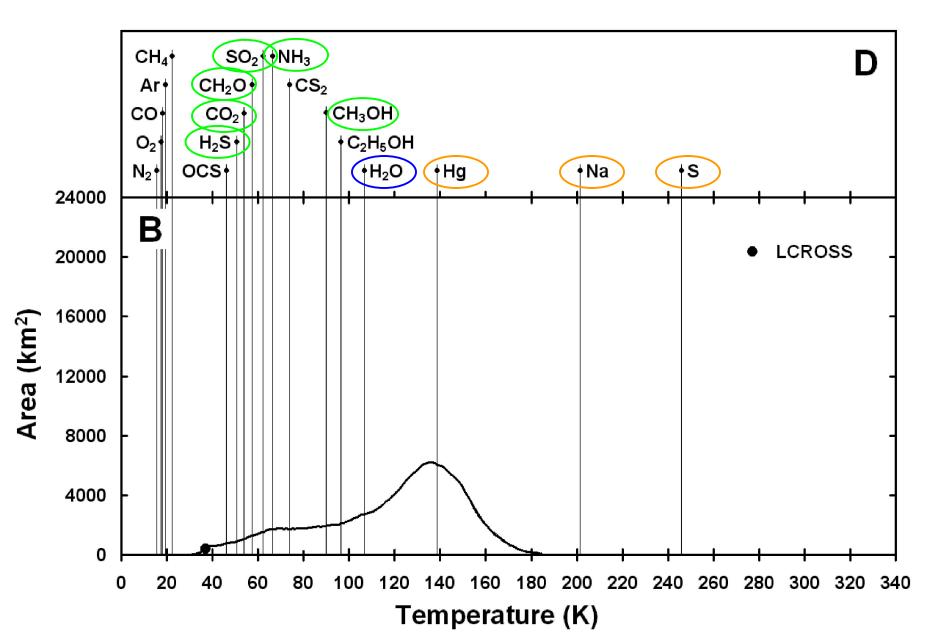
Paige et al., Science 339, 300-302, 2013

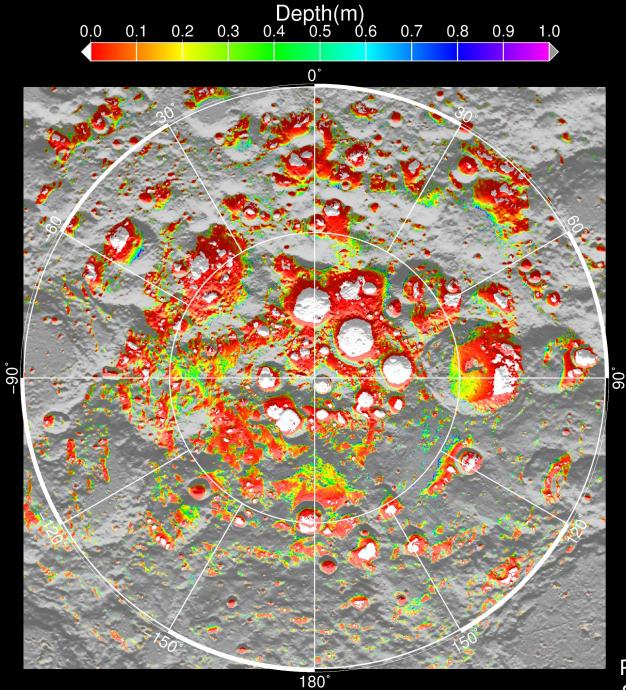


Depth to 1 mm water ice sublimation in 1 BY

Paige et al. *Science* (2010)

#### South Polar Annual Average Temperature Histogram

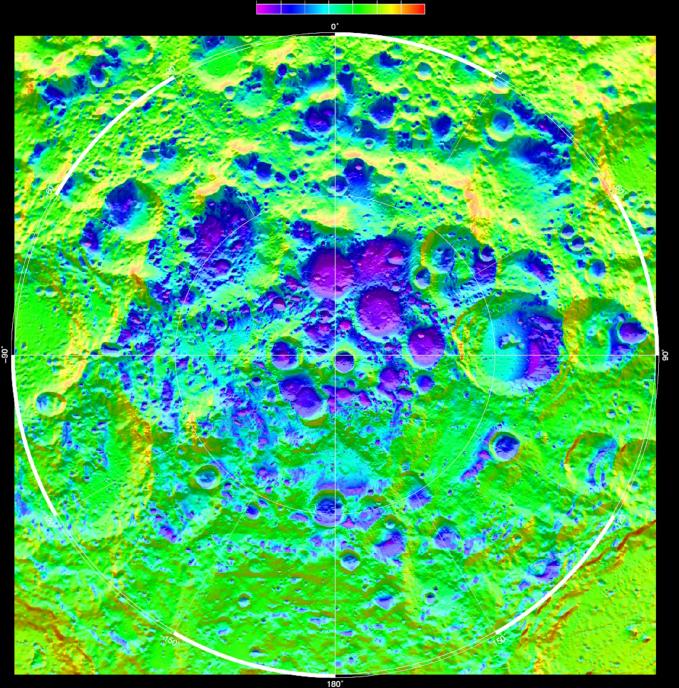




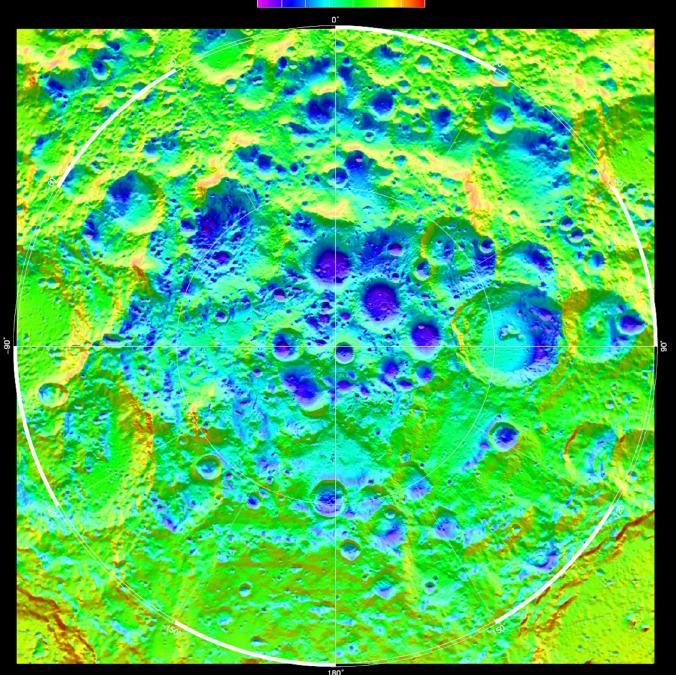
Depth to 1 mm water ice sublimation in 1 BY

Paige et al. *Science* (2010)

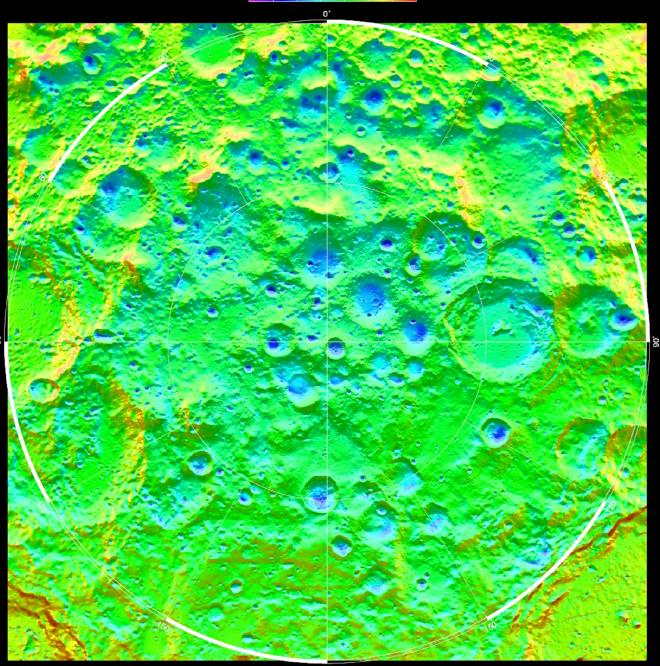
Model Calculated 80s\_dtop.avgminmax South Polar Annual Average Surface Temperatures



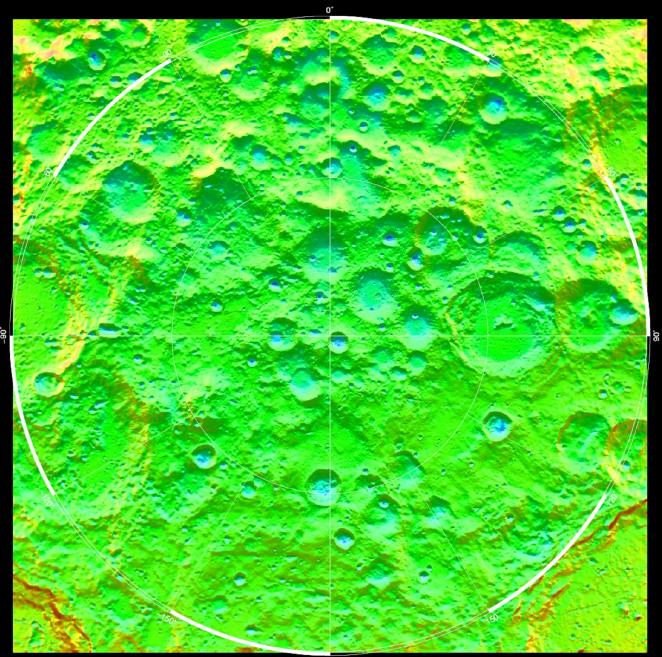
Model Calculated Annual Average Temperature  $\Theta$ =1.54° Model Calculated 80s\_4.00.avgminmax South Polar Annual Average Surface Temperatures



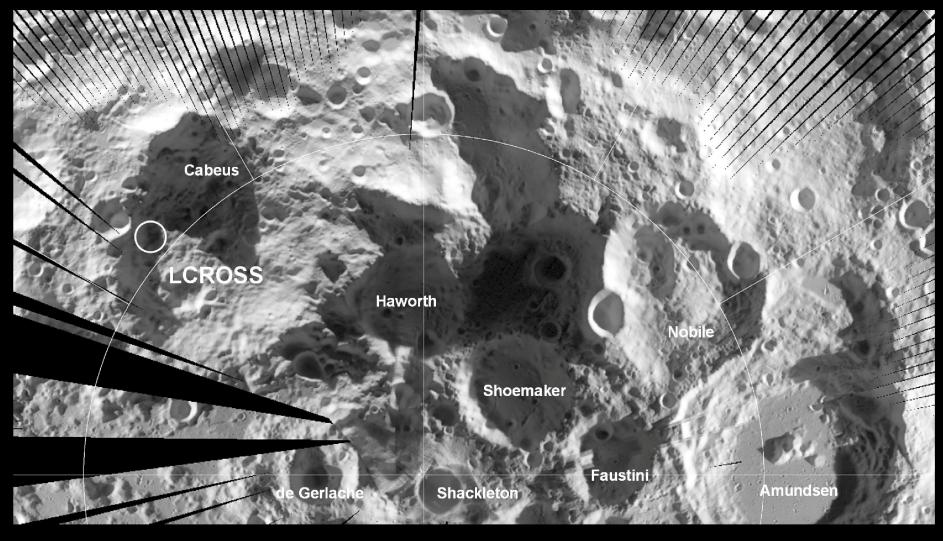
Model Calculated Annual Average Temperature  $\Theta$ =4° Model Calculated 80s\_8.00.avgminmax South Polar Annual Average Surface Temperatures



Model Calculated Annual Average Temperature  $\Theta=8^{\circ}$  Model Calculated 80s\_12.0.avgminmax South Polar Annual Average Surface Temperatures



Model Calculated Annual Average Temperature  $\Theta$ =12°



 $25 \ 50 \ 75 \ 100 \ 125 \ 150 \ 175 \ 200 \ 225 \ 250 \ 275 \ 300$ 

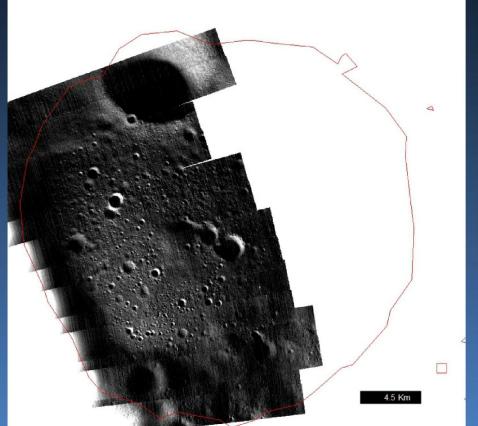
Diviner Channel 8 Brightness Temperature Map (K)

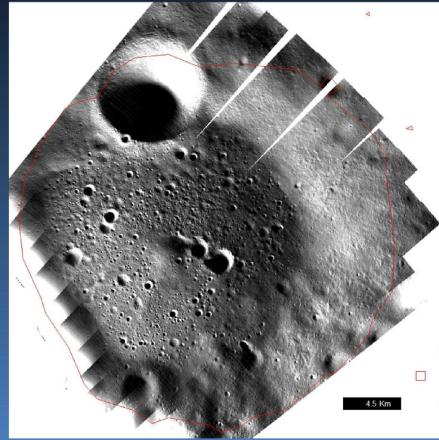
# Faustini Crater 2009 vs. 2012

Koeber et al, NLSI LSF, 2013

2009

### 2012

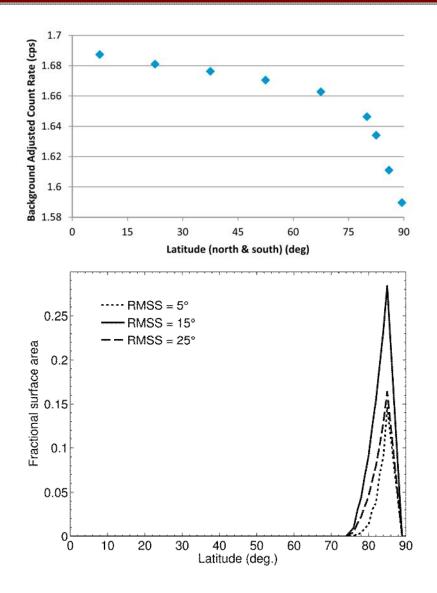




NAC Exposure time 12 ms, resampled to 10-m mosaic

NAC Exposure time 24 ms, resampled to 20-m mosaic

## Comparison with "Background" Epithermal Neutron Suppression



LEND background neutron suppression (Boynton et al., 2012)

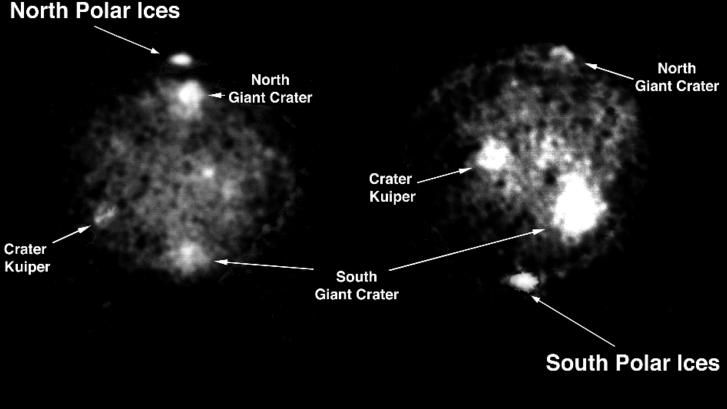
Fractional surface area with temperatures conducive to stable surface water ice Thermal Stability of Volatiles in the North Polar Region of Mercury



### **Polar Caps on Mercury?**

#### Goldstone/VLA Radar Maps of Mercury's Unimaged Side

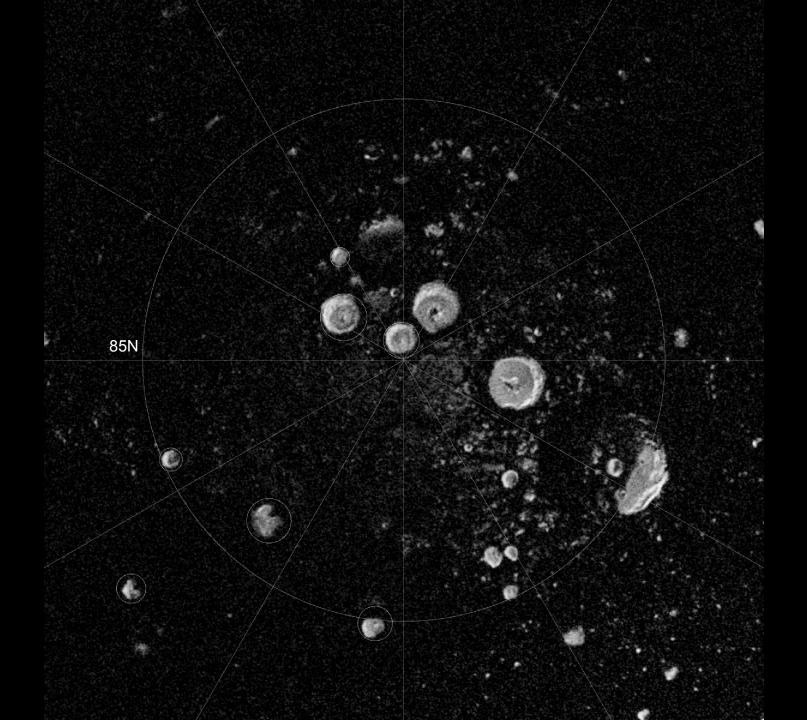
(courtesy of B. J. Butler, M. A. Slade, D. O. Muhleman)

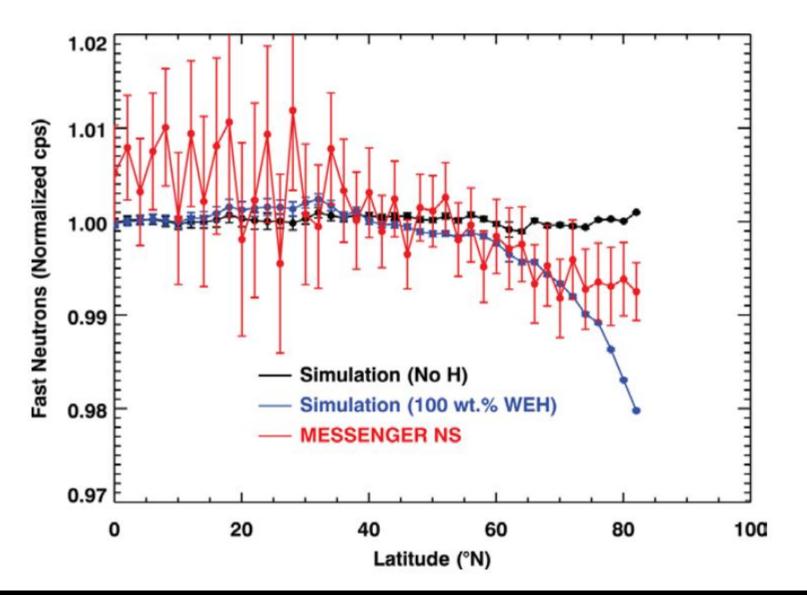


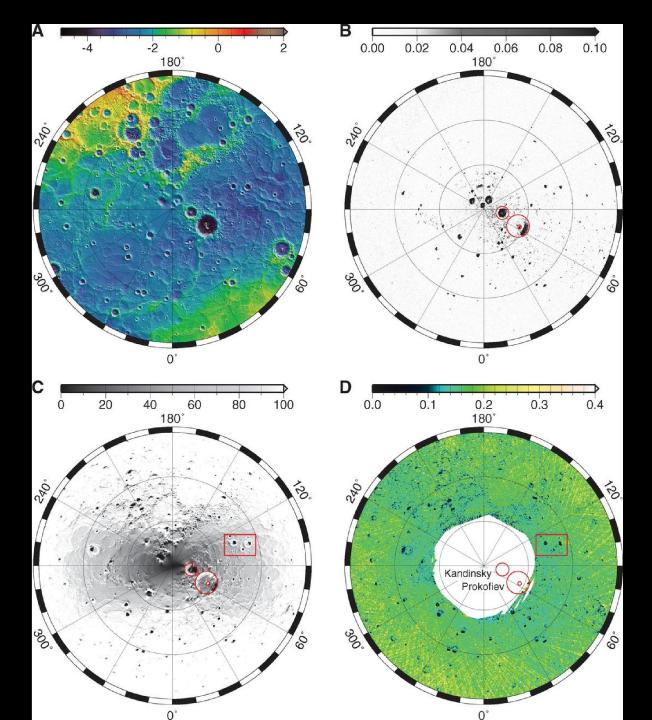


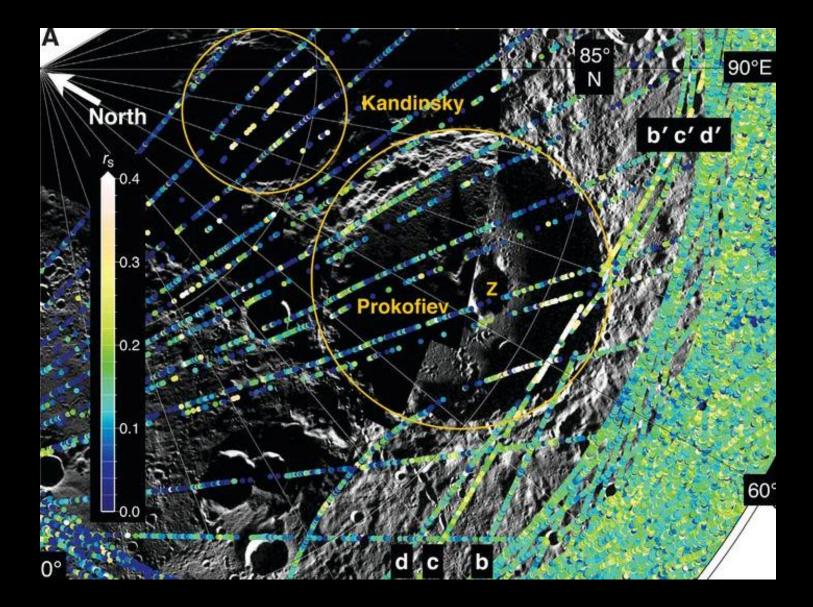
Feb. 21, 1994

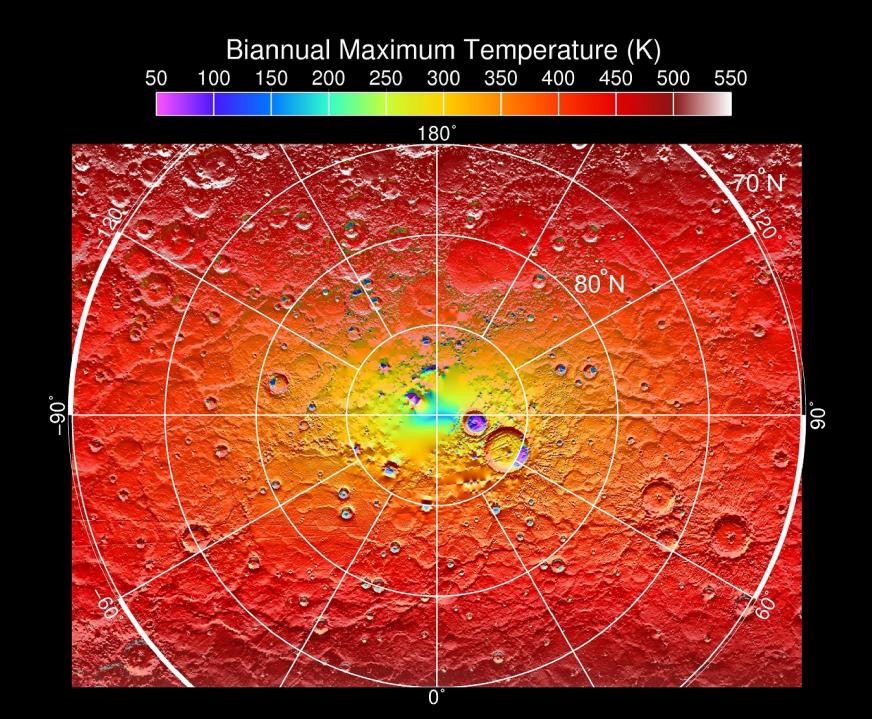
- Strong depoloarized echoes from both polar regions
- Mercury polar radar signatures suggest presence of thick, nearly pure water ice deposits below a thin (< 1 meter) covering of soil</li>

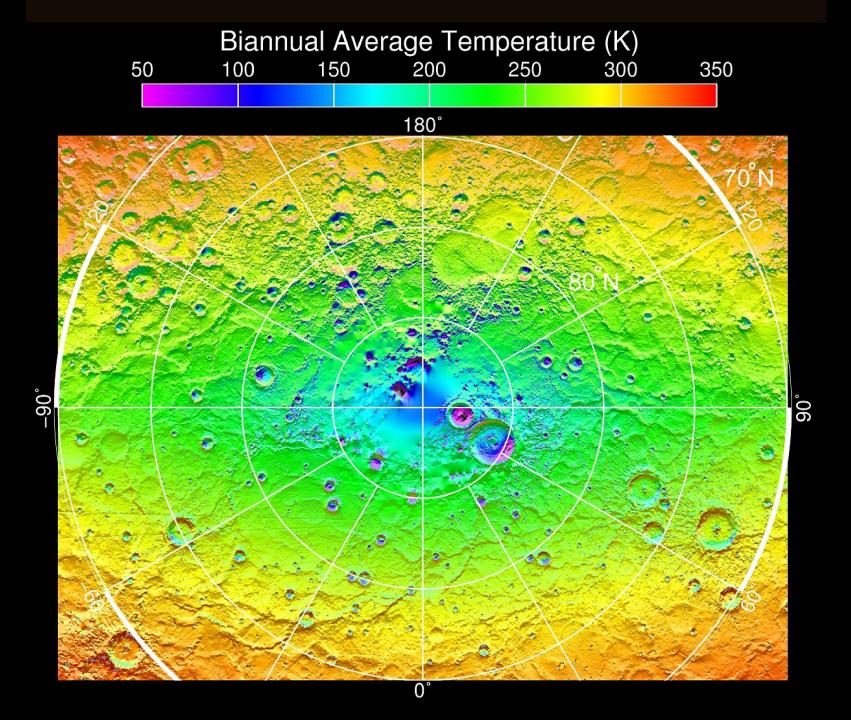


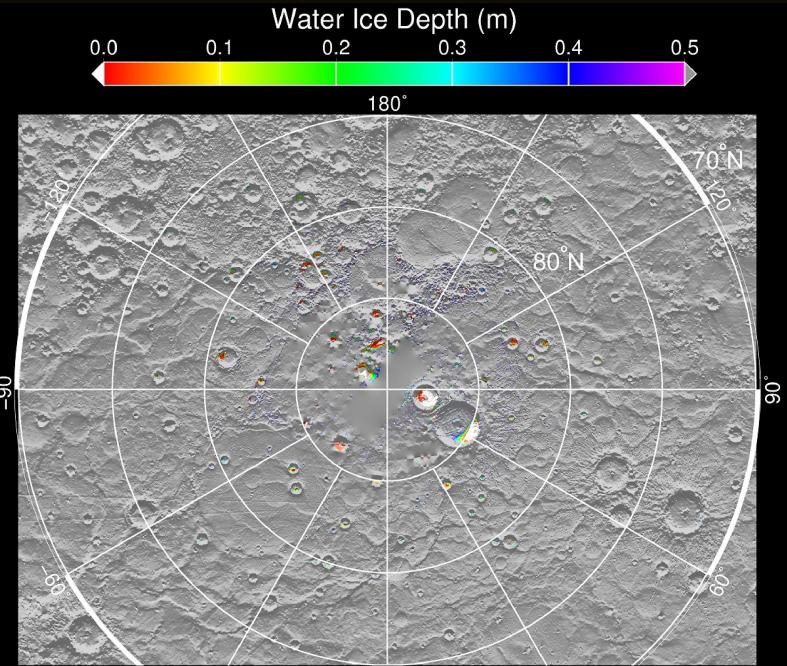




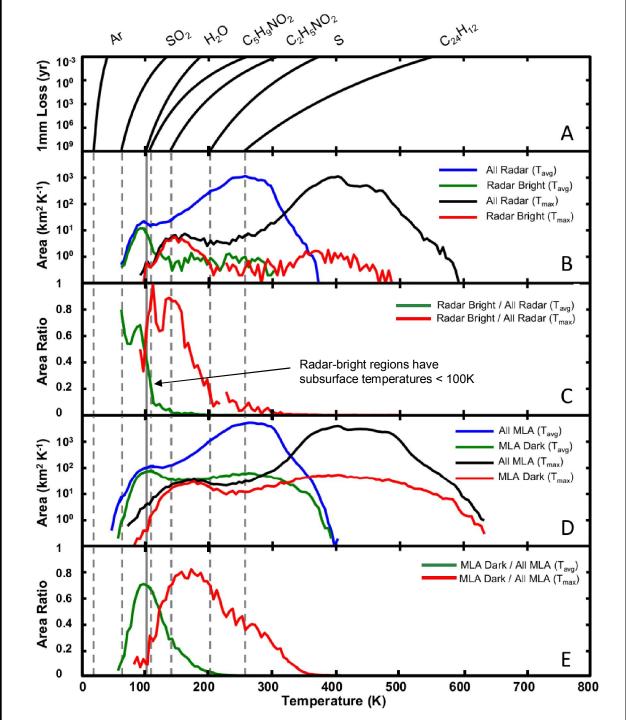


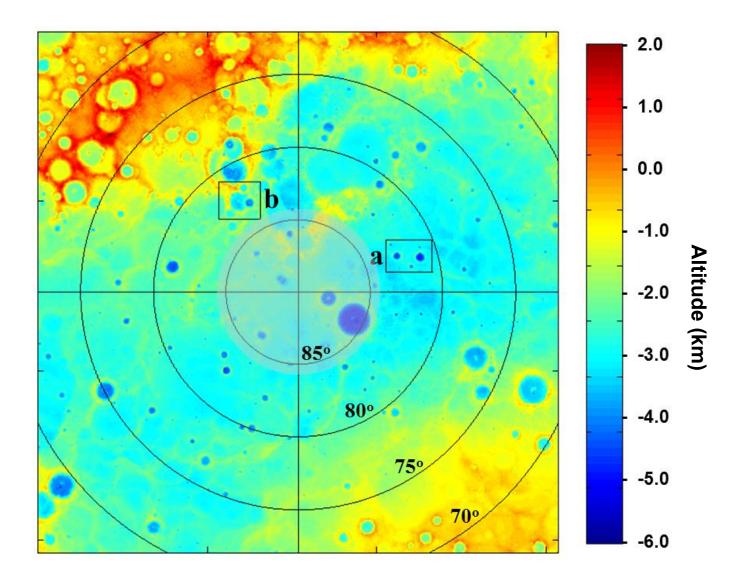


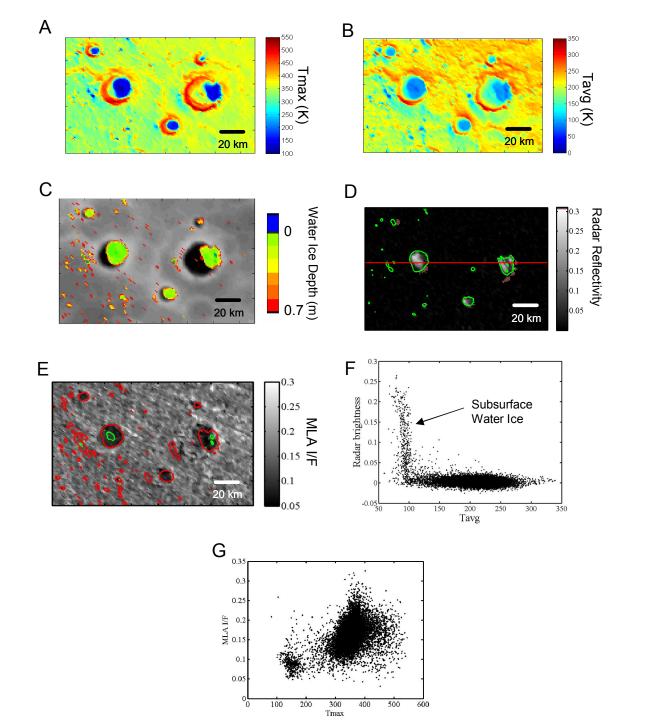




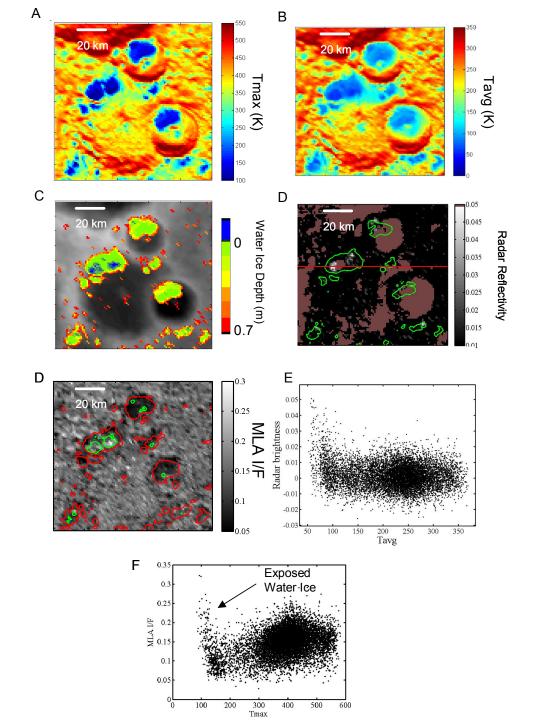
。 06-





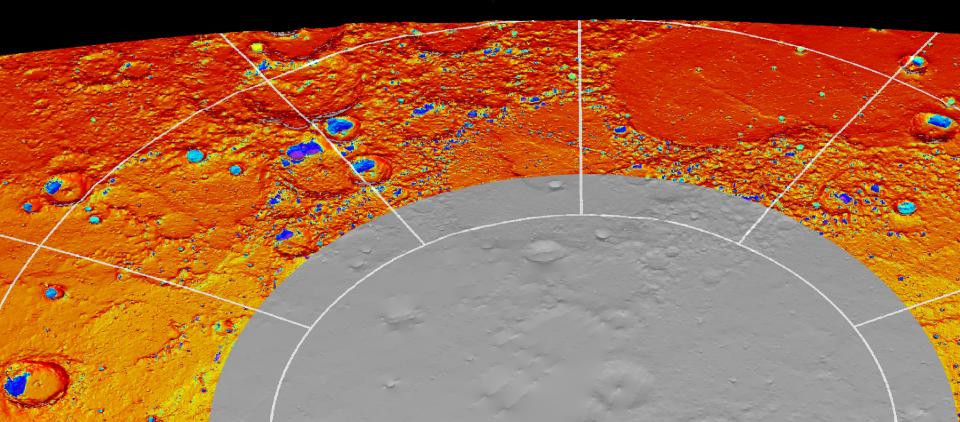


Region A

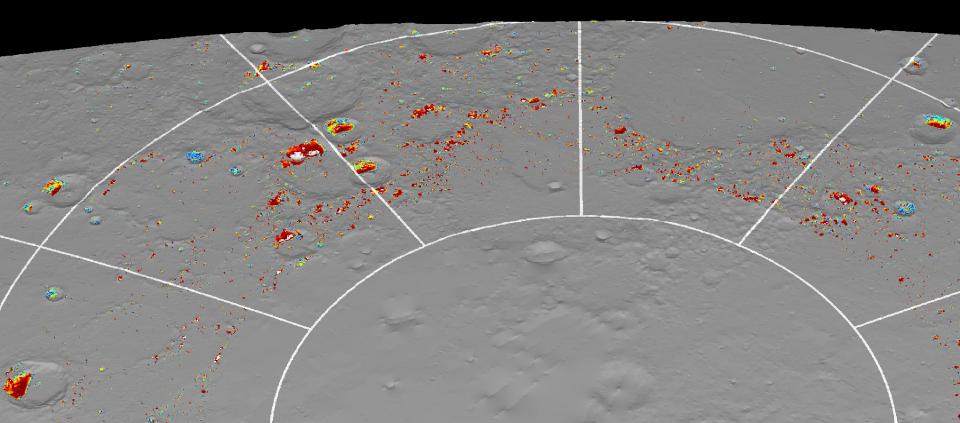


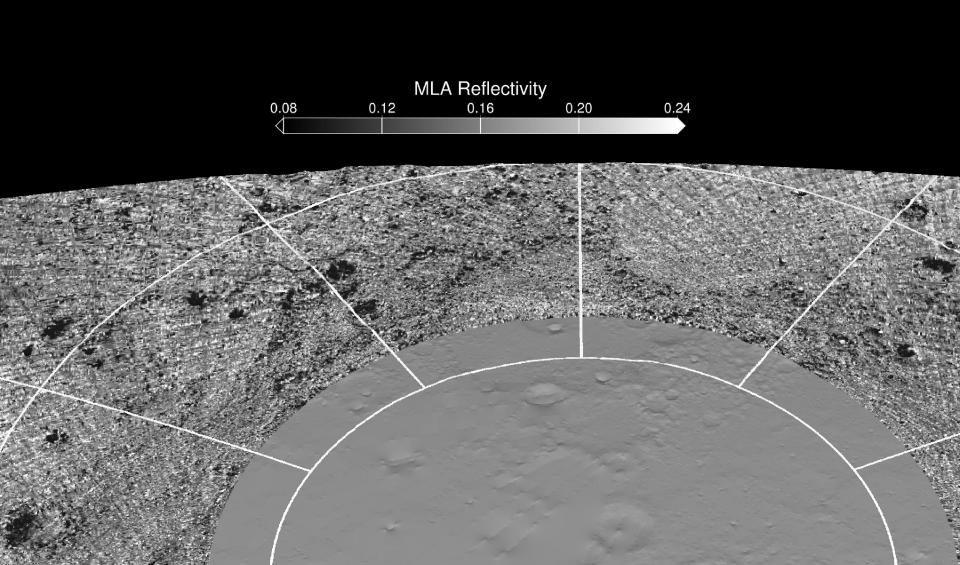
**Region B** 

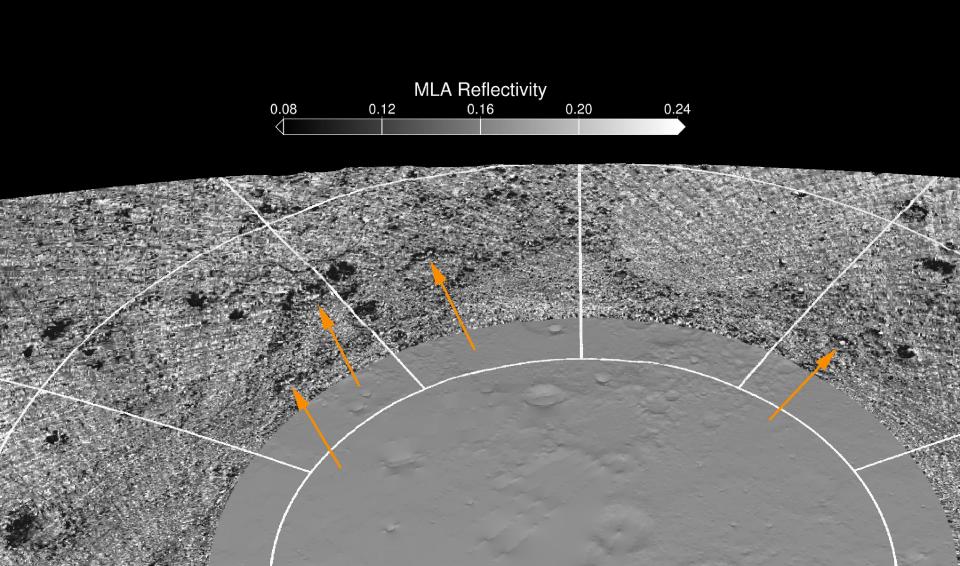




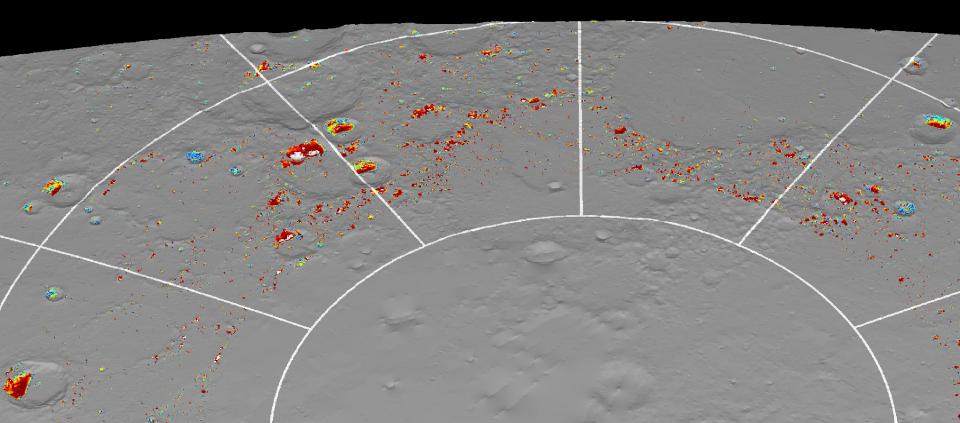














MESSENGER MErcury Surface, Space Environment, GEochemistry, and Ranging

Prokofiev

Sunlit mosaic + shadowed terrain!

Chabot et al, 2013



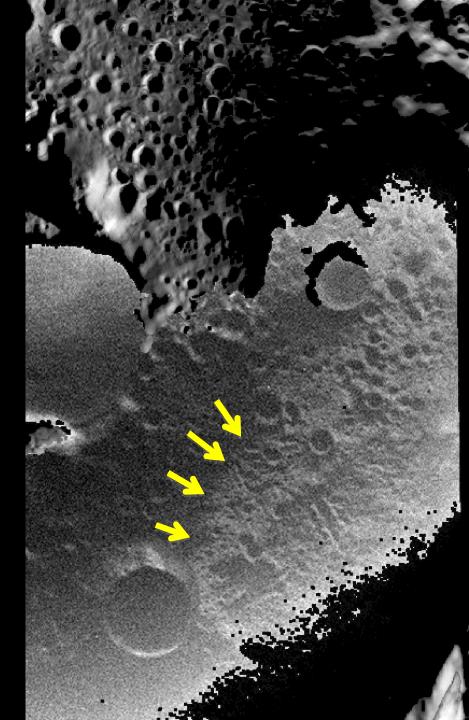
# Prokofiev

## Zoom in

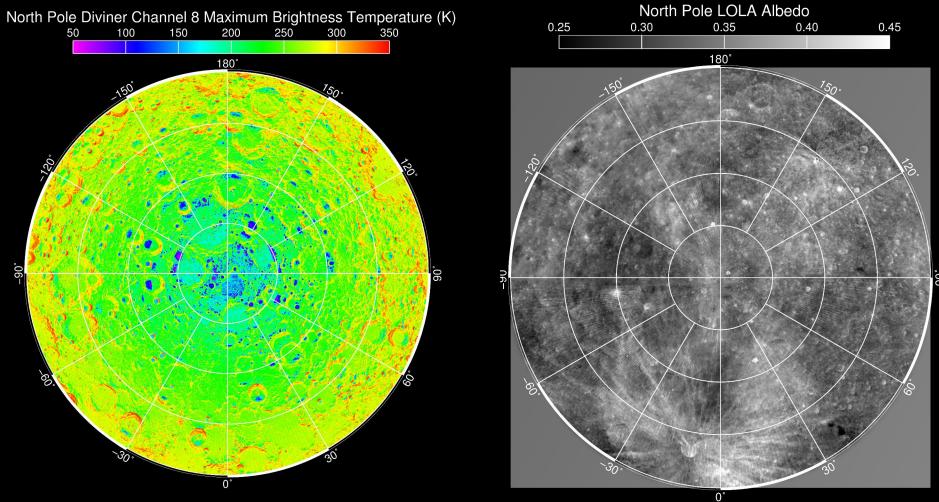
Sunlit mosaic + shadowed terrain

Terrain imaged is covering radar-bright, shadowed region

2<sup>nd</sup> observation – evidence for an albedo boundary that is consistent with boundary of the radar-bright material

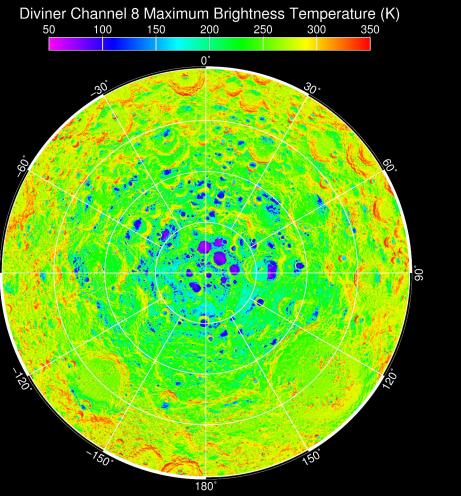


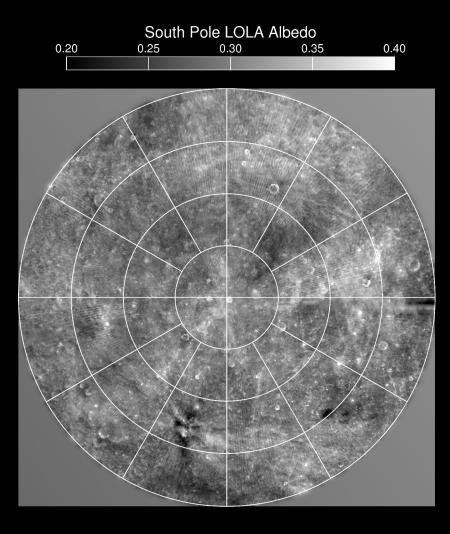
#### North Pole

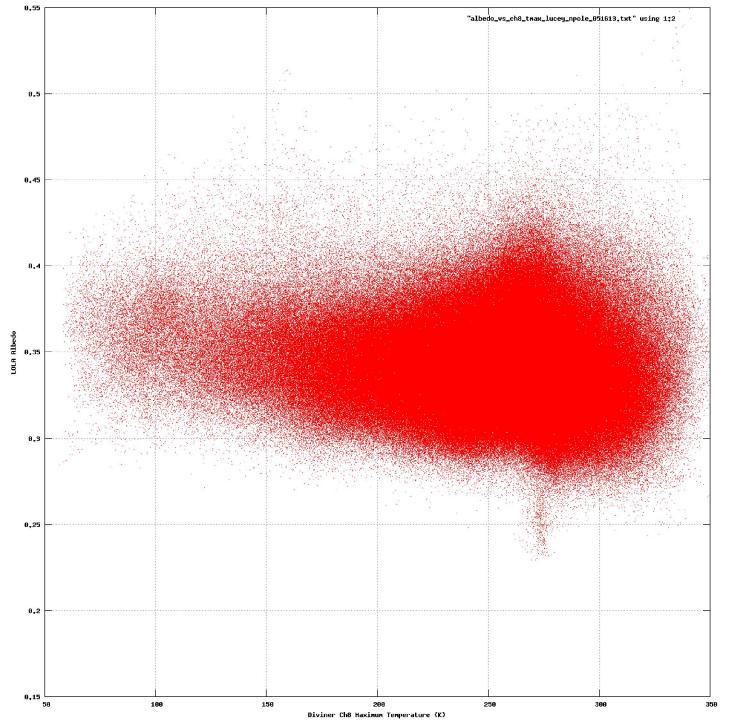


. 06

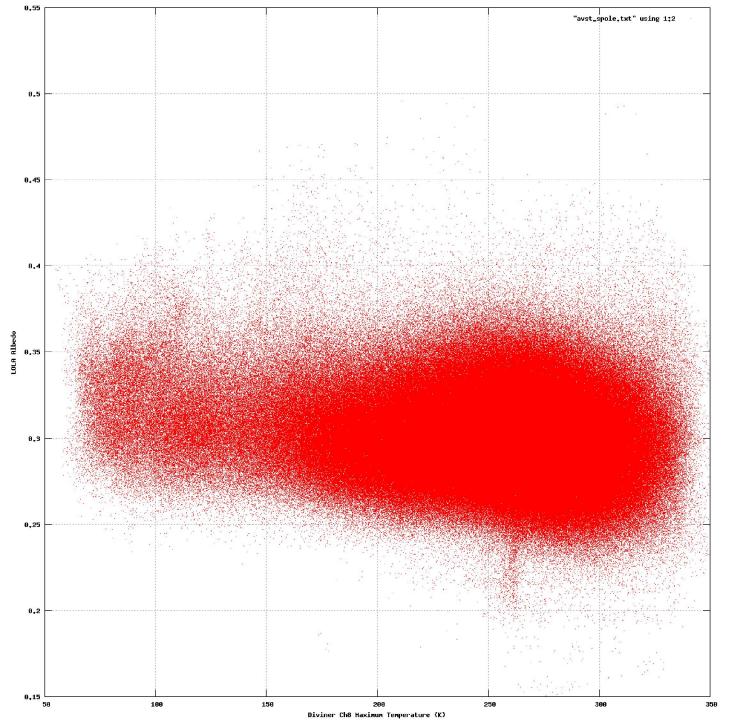
#### South Pole



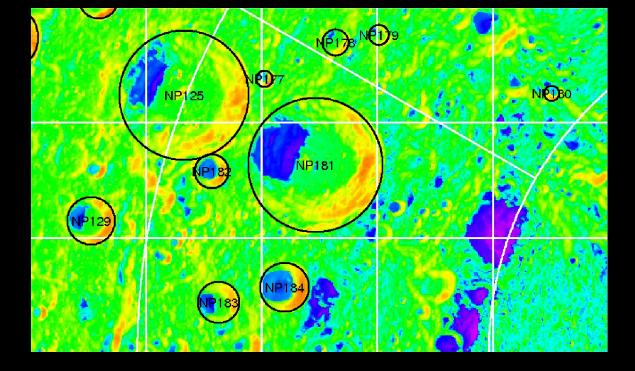




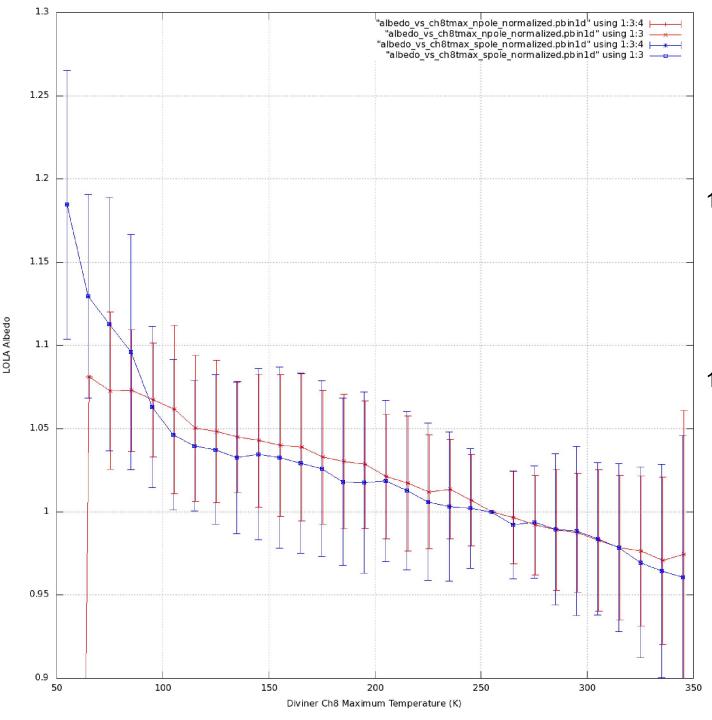
### North Polar Region



## South Polar Region



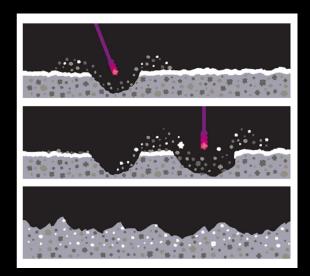


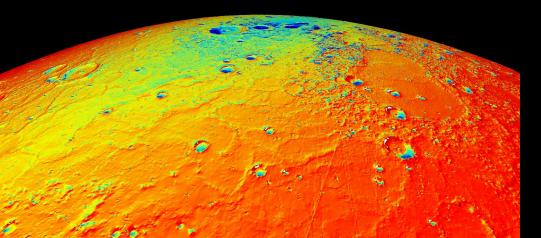


Preliminary Conclusions from Normalized Crater Reflectance Study

- 1. Strong, systematic increase in crater normal reflectance deviations with decreasing temperature
- Possible high reflectance deviations within the coldest south polar craters.

# Mercury vs. The Moon





 Ice deposits on the Moon and Mercury are destroyed by UV photolysis, sputtering and impact gardening

•The highly organized present-state of Mercury's polar ice deposits suggests that the sources of water and the mobility of water in Mercury's environment are sufficiently robust to overcome the combined effects of all other processes that would tend to destroy and disrupt them

• The relative scarcity and apparently disorganized state of ice on the Moon suggests that the converse is true