



Metrology for Space-based Large Apertures

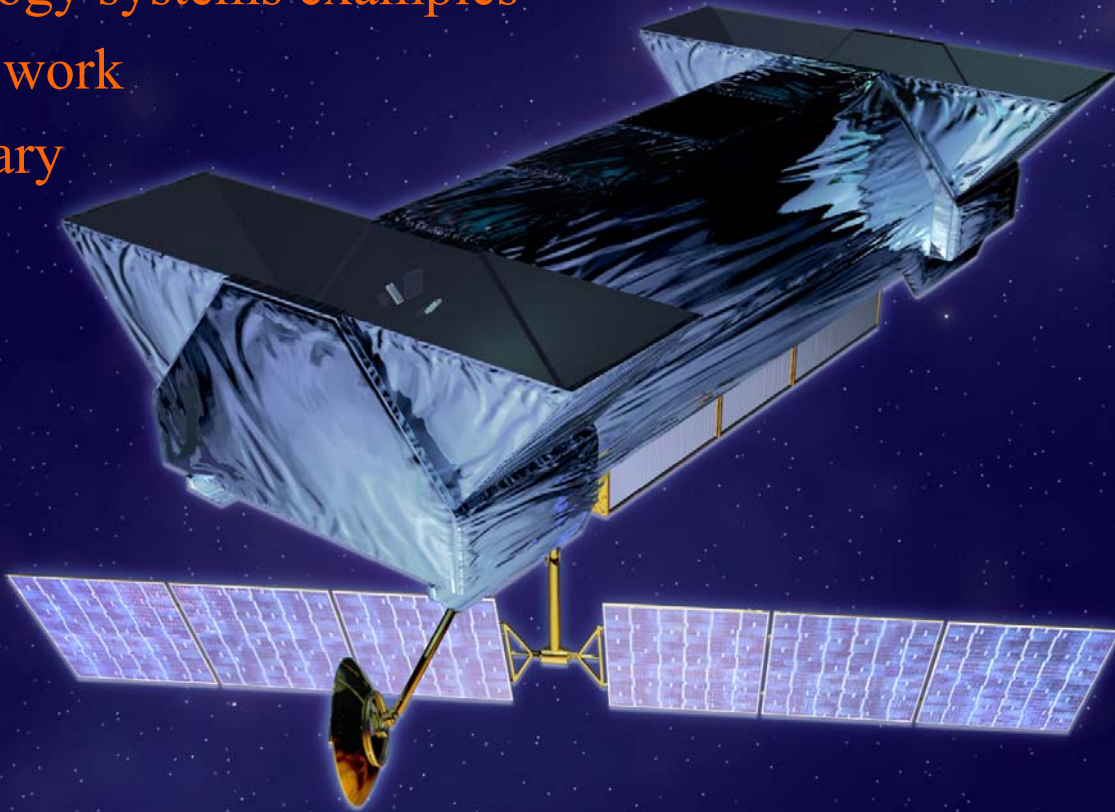
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California Institute of Technology

November 10 -- 11, 2008

- Role of metrology in large space based structures
- Metrology systems examples
- Future work
- Summary



Option Space:

A: Light-weight, deployable, stable structure with no or limited metrology, to

Z: Low-cost structure and a very capable metrology system to correct for deformations.

Where is the optimum?

- Mechanical scaling law:

- Mechanical optical error scales to the 4th power of aperture diameter (Peterson, et al.)

$$\delta \propto \frac{D^4}{h^2} \frac{1}{E/\rho} f(\nu) g(P)$$

- The larger the aperture, the more (a lot more!) challenge on structures

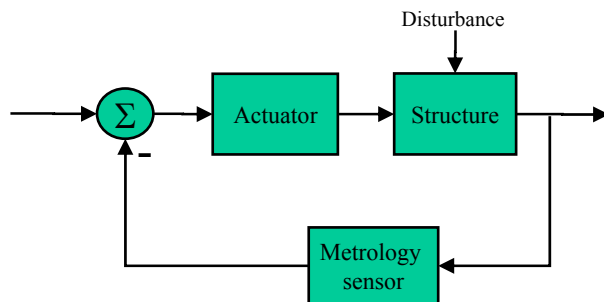
- Metrology sensor scaling law:

- Metrology sensor error scales linearly with the dimension

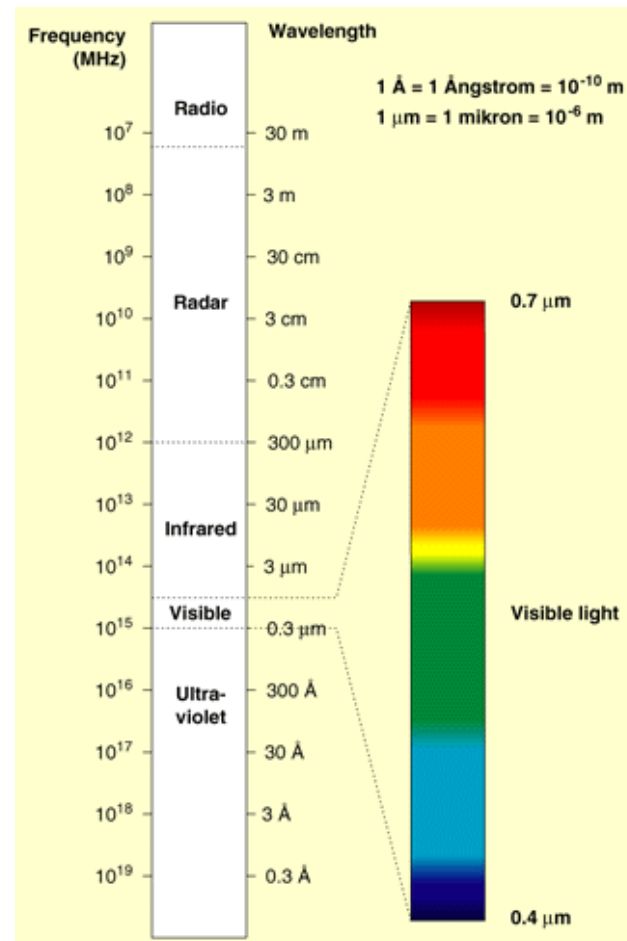
$$\delta x_M = \frac{\delta x_R}{\Delta \phi_R} \Delta \phi_M = \frac{\delta x_R}{x_R} x_M$$

- Save your \$\$ and kg from structures to build metrology?

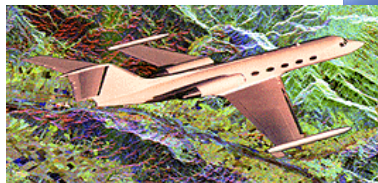
- Open-loop: metrology provides knowledge of structure dimension over time and metrology measurement is used in post-processing
- Closed-loop: metrology provides structure deformation measurement which is then used to maintain structure rigidity via controllers and actuators



- Dimension of large apertures/structure are stabilized (at least have the knowledge) to a small fraction of the observation wavelengths
- Primary driver of metrology performance:
 - Radar: mm to um metrology accuracy
 - IR/VIS: um to sub-nm
 - Dimension: typically >10m
- Types of metrology measurement:
 - Provide knowledge of absolute distance (abs MET)
 - Provide knowledge of change of distance - displacement (rel MET)



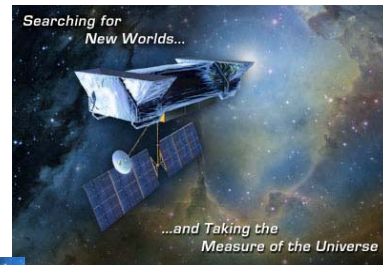
GeoSAR



SRTM



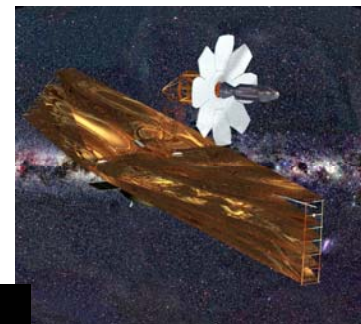
SIM



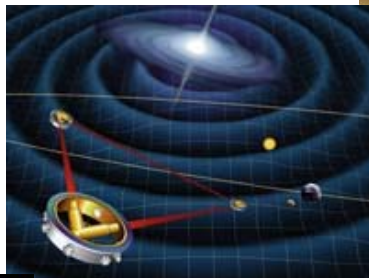
TPF



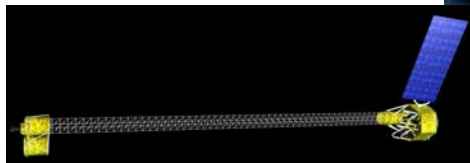
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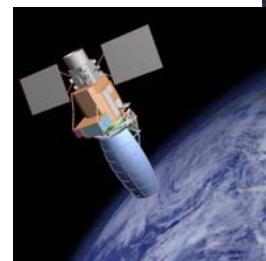
LISA



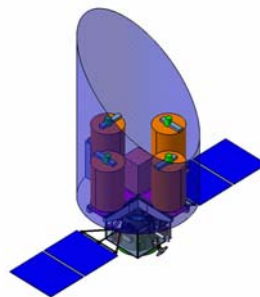
NuSTAR



CAMEO



DAVINCI



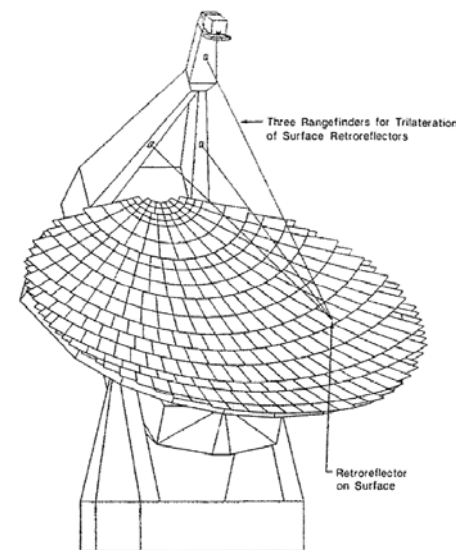
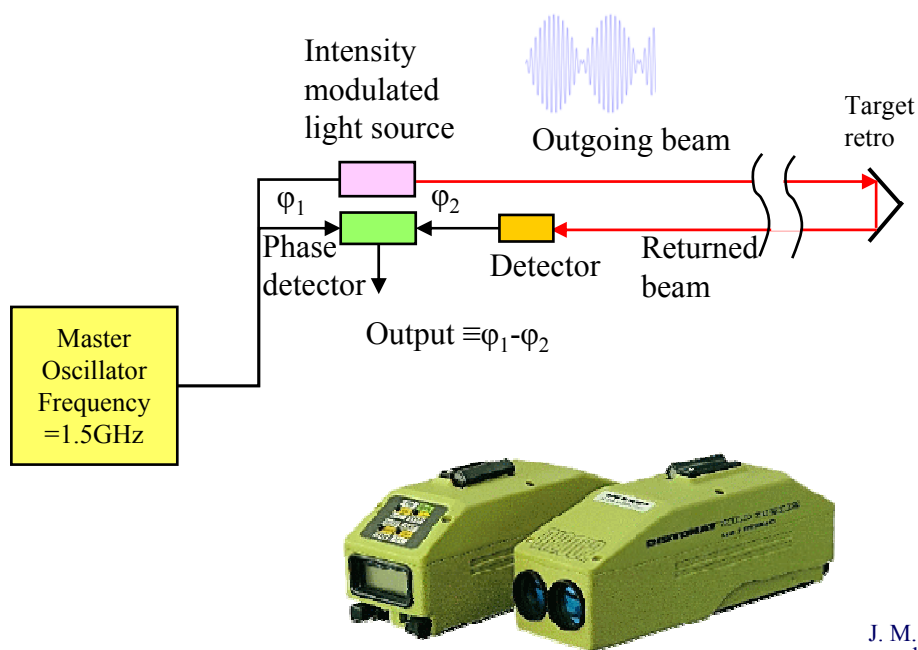


Examples of Metrology Systems



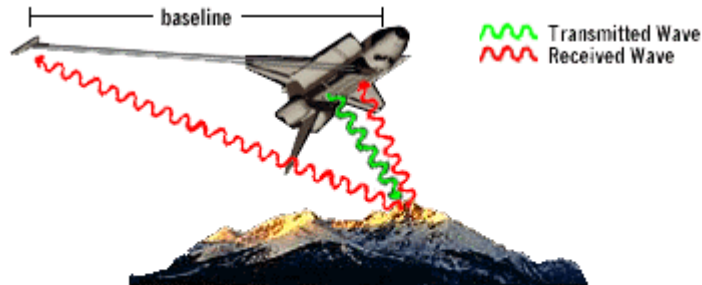
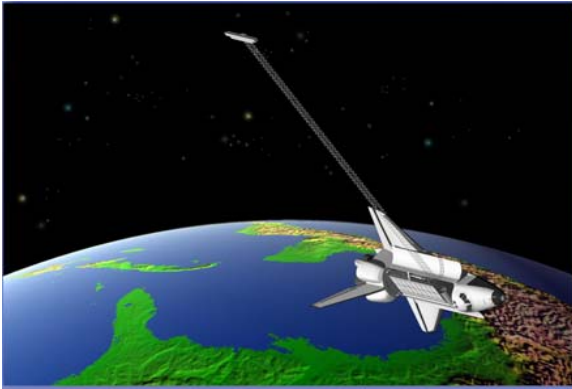
- Amplitude-modulated metrology
 - Application example: SRTM
- Laser heterodyne metrology
 - Application example: SIM

- Ranging accuracy and resolution: mm - um
- Simple and lost cost
 - “Non-interferometric”
 - Light source can be “low coherence”, e.g. LED
- Most commercial surveying instruments based on this principle



J. M. Payne, D. Parker, and R. F. Bradley, “Rangefinder with fast multiple range capability”, *Rev. Sci. Instrum.* 63 (6), pp 3311, June 1992

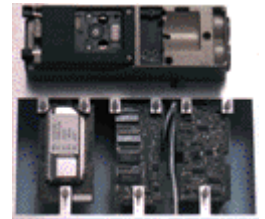
- Illustration of State of the Art



Radar signals being transmitted and received in the SRTM mission (image not to scale).

- Longitudinal (base-line length): modified commercial range finder (inboard)

- resolution: 0.5 mm
- Accuracy: 2mm
- operating range: ~60m
- retro-reflectors (outboard)



- Lateral:
 - Star tracker inboard
 - 3 LED point source outboard
- Open loop

Riley M. Duren and Eldred F. Tubbs, "A Modified Commercial Surveying Instrument For Use as a Spaceborne Rangefinder", *IEEE Aerospace Conference* (2000, Big Sky, Montana)

- Want point to point displacement at \sim nm to pm resolution
- Use laser heterodyne approach
- Have high bandwidth
 - kHz or higher
 - Low latency
- Good for closed-loop control of high freq jitter
- Three types of interferometers:

1. Polarization separation

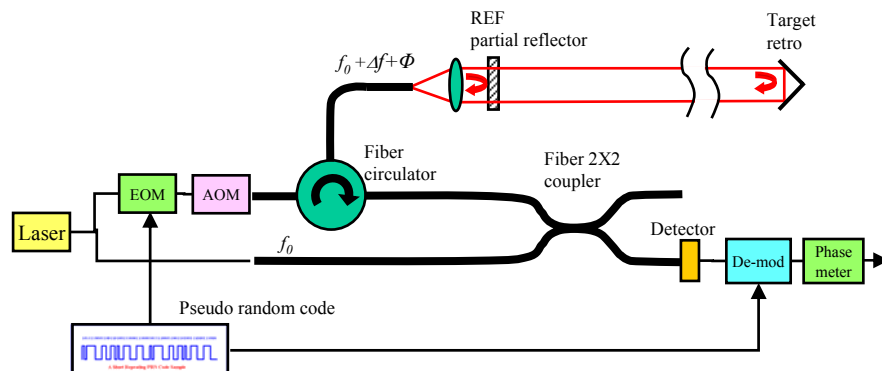
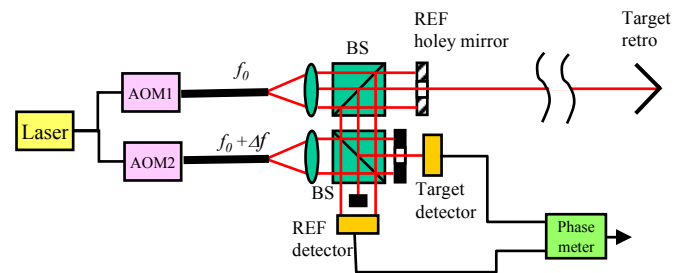
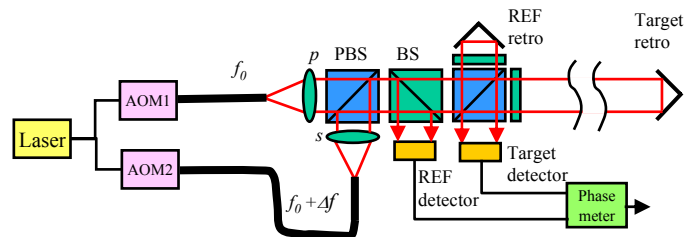
- N. Bobroff, “Recent advances in displacement measuring interferometry”, *Meas. Sci. Technol.*, vol. 4, pp. 907-926, 1993

2. Spatial separation

- F. Zhao, “Development of high precision laser heterodyne metrology gauges,” *Proc. SPIE*, Vol. 5634, pp. 247-259, 2005.

3. Temporal separation

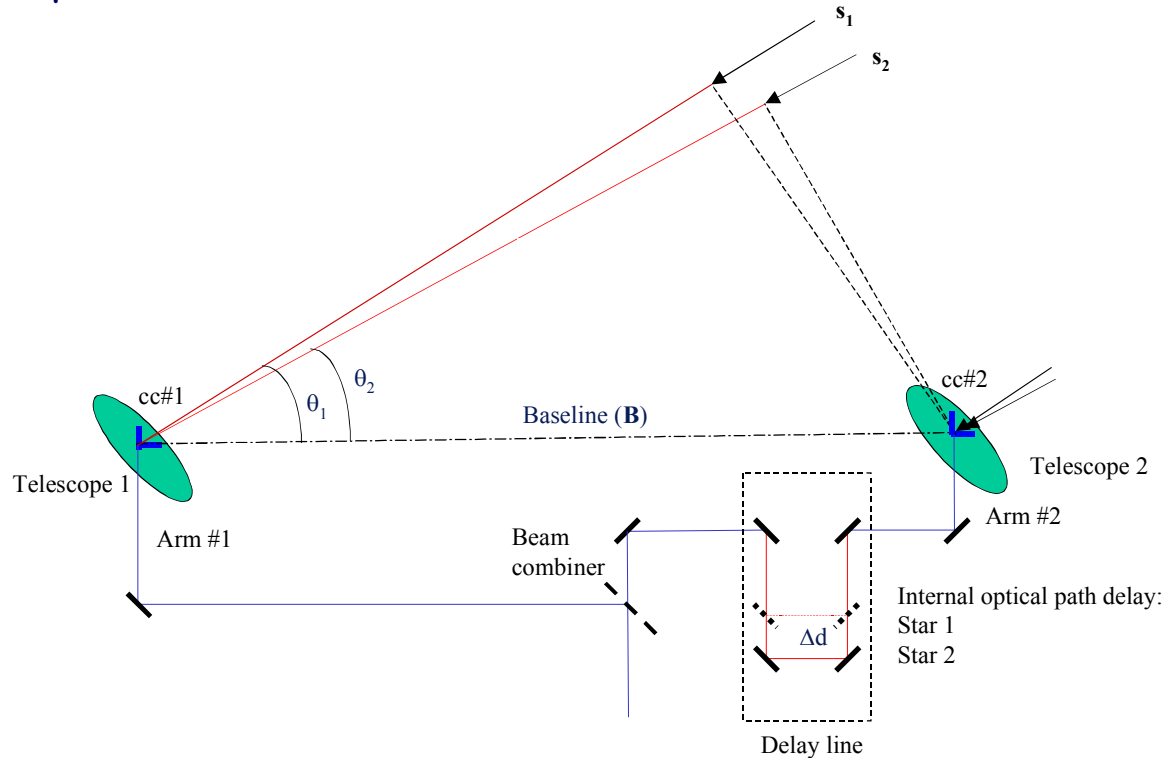
- O. P. Lay, et al, “Coherent range-gated laser displacement metrology with compact optical head”, *Opt. Lett.* Vol. 32, No. 20 October 15, 2007



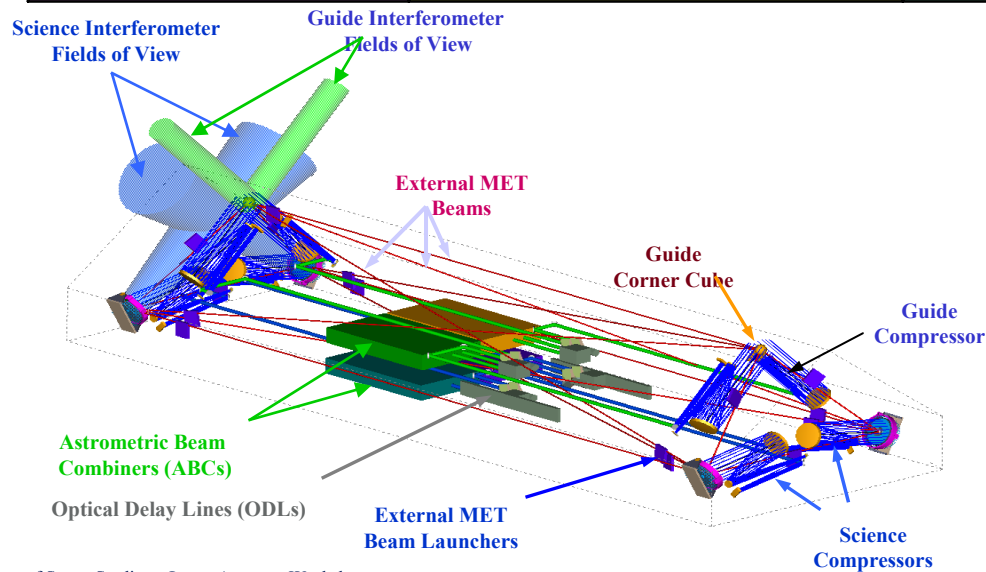
- SIM model in its simplest form:

$$\Delta d = \vec{B} \bullet (\vec{s}_1 - \vec{s}_2)$$

- Laser metrology gauges to measure:
 - Δd (Internal Metrology) – displacement to 3pm RMS
 - \vec{B} (External Metrology) – displacement to 8pm pm RMS and absolute distance to $3\mu\text{m}$ RMS



	Internal metrology (Δd)	External metrology (B)
Number of gauges	3 total (1 science, 2 guide)	14 total
Fiducial distance	8m	2.5 – 9m
Range of motion	+/-1.2m (ODL)	~100 μ m
Velocity	200mm/s while slewing 500 μ m/s while observing	~1 μ m/s
Accuracy (absolute)	Not needed	3 μ m
Accuracy (relative)	3pm (90 s), 45pm (1 hr)	8pm (90 s), 57pm (1 hr)
Telemetry rate	16kHz	1kHz





- Picometer knowledge
 - Picometer laser metrology
 - Picometer starlight fringe position measurement
 - Picometer deformation of large (35 cm) optics
- Nanometer control (for dim star fringe sensing)
 - Vibration isolation
 - High bandwidth active optics

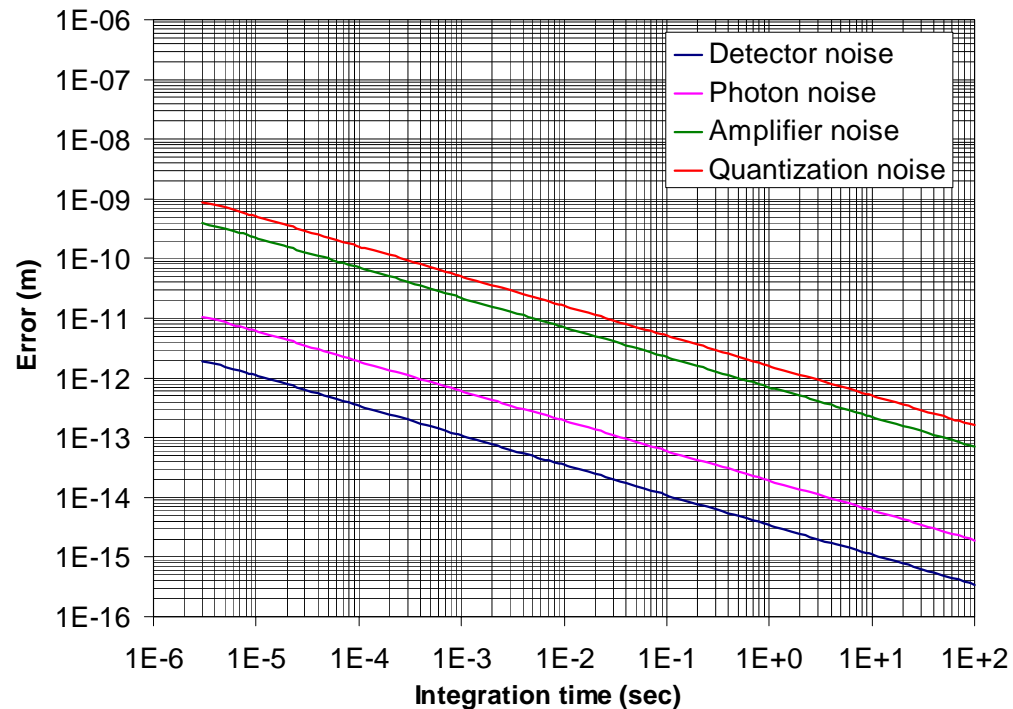
- Metrology gauge errors:
 - Random noise (error)
 - Systematic error:
 - Diffraction error
 - Cyclic error (periodic nonlinearity)
 - Environmental error (thermal drift)
- Alignment error:
 - Cosine error (pointing);
 - Abbé error (overlap);

Detector noise	$\varepsilon = \lambda \cdot NEP \cdot \frac{\sqrt{B}}{4\pi\eta P}$
Photon noise	$\varepsilon = \frac{1}{4\pi} \sqrt{\frac{2hc\lambda B}{\eta P}}$
Amplifier noise	$\varepsilon = \frac{\lambda}{2\pi} \frac{\sqrt{4kTRB}}{\eta PR}$
Phase meter quantization noise (zero-crossing)	$\varepsilon = \frac{\lambda}{\sqrt{12}} \frac{f_{het}}{f_{clock}}$

- $P=200\text{nW}$
- $NEP=5 \times 10^{-15} \text{ W}/\sqrt{\text{Hz}}$,
- $R=100\text{k}\Omega$,
- heterodyne $f_{het}=300\text{kHz}$,
- clock $f_{clock}=128\text{MHz}$,
- bandwidth $B=540\text{kHz}$,

- Use higher laser power
- Use lower heterodyne frequency
- Use longer integration time (averaging)

$$\langle \delta x \rangle = \frac{\varepsilon}{\sqrt{N}} = \frac{\varepsilon}{\sqrt{f_h \tau}}$$

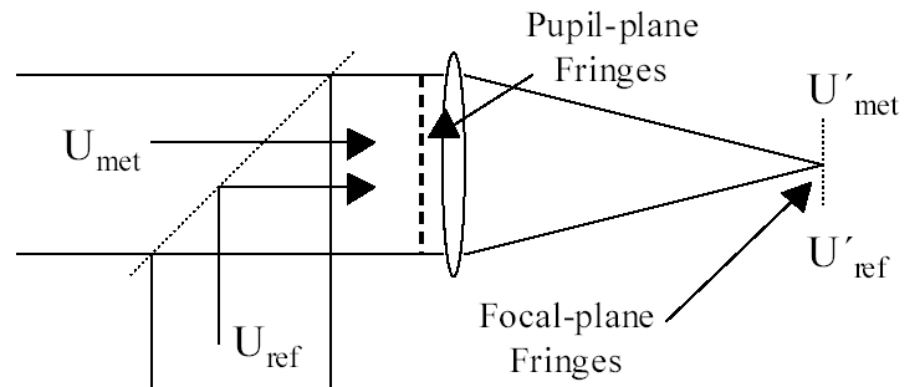


- Geometric path length \neq average phase
 - Due to beam diffraction
- Fresnel (paraxial) approximation

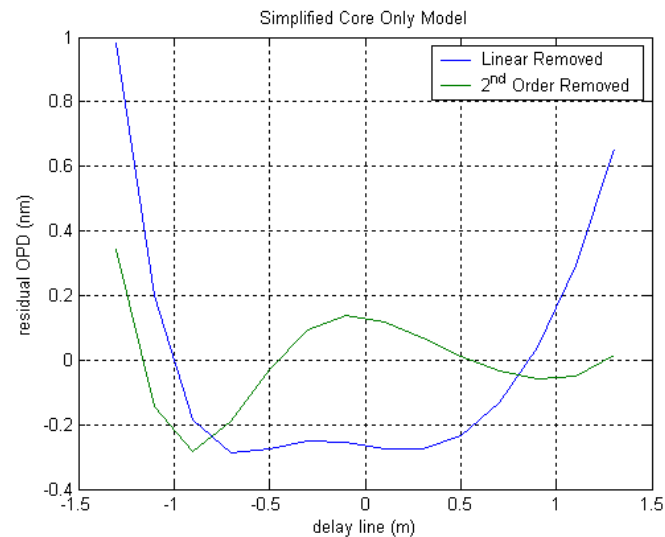
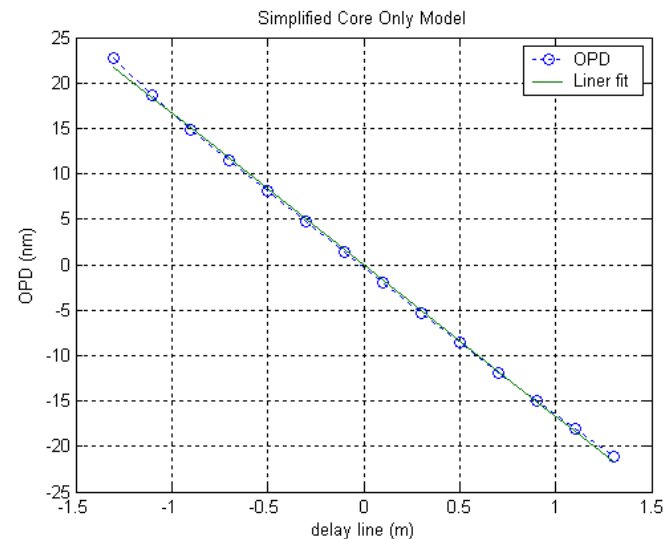
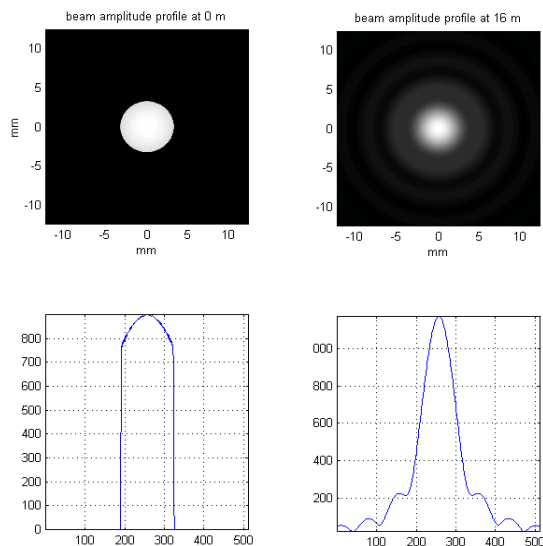
$$U(x, y; z) = \frac{\exp(ikz)}{i\lambda z} \int \int_{-\infty}^{\infty} U_l(x_l, y_l) \exp\left\{i \frac{k}{2z} \left[(x - x_l)^2 + (y - y_l)^2\right]\right\} dx_l dy_l$$

- Average phase

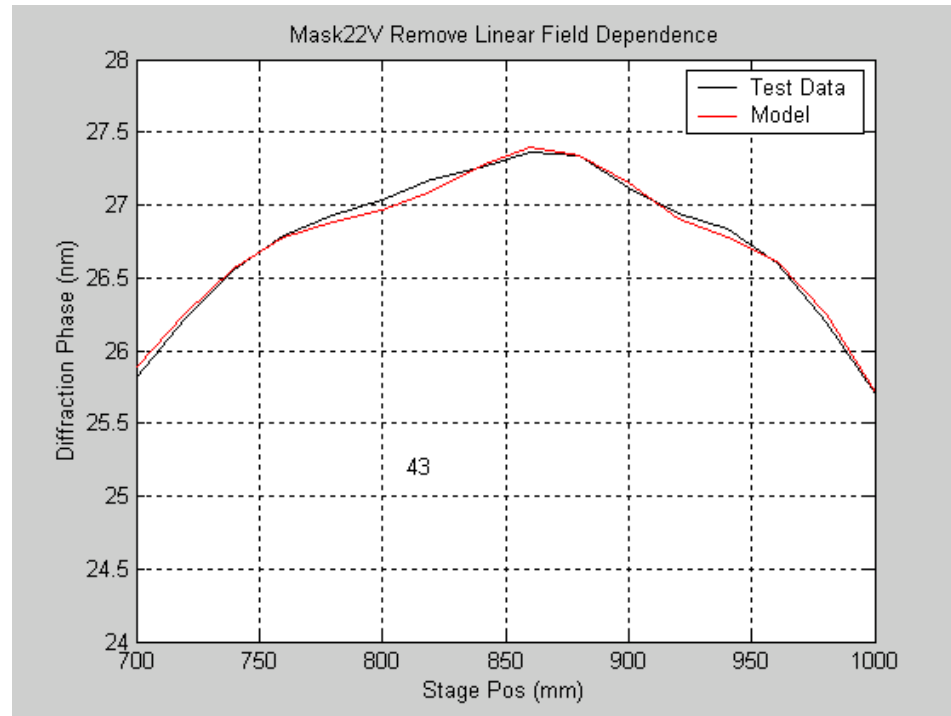
$$[\phi]_{av} = \arg \left[\iint_{det} U_{met}(x, y) U_{ref}^*(x, y) dx dy \right]$$



- Apply diffraction correction (calibration) based on model calculation
- Example: truncated Gaussian



- Successfully used in several ground testbed demonstrations (MAM, DTB)

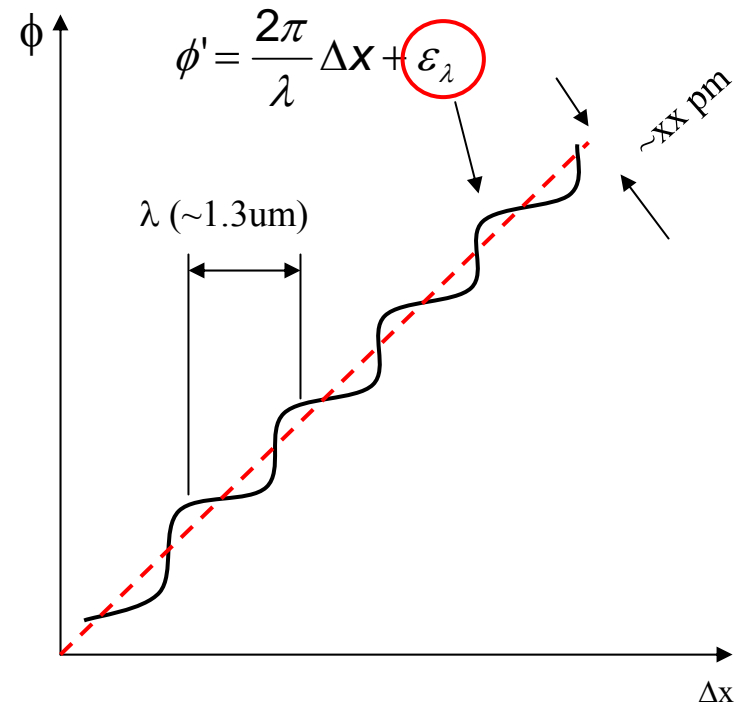
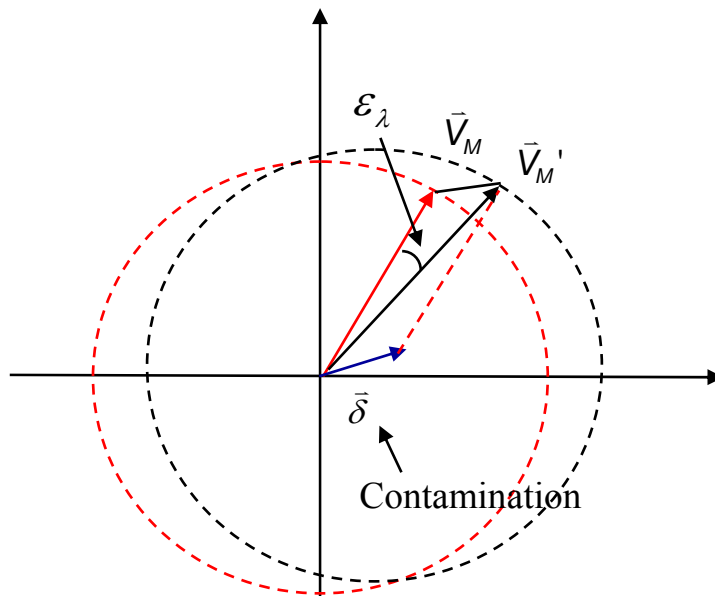


D. B. Schaefer, et al, "Diffraction hardware testbed and model validation,"
Proc. SPIE vol.4852, pp. 302-314, 2002.

- Defined as “periodic (λ) nonlinearity” between measured phase and actual displacement
- Caused by “contamination” to the heterodyne signals

$$V_M' = V_M \sin\left(2\pi f_h t + \frac{2\pi}{\lambda} \Delta x\right) + \delta \sin(2\pi f_h t + \phi_1)$$

$$\bar{V}_M' = \bar{V}_M + \bar{\delta}$$

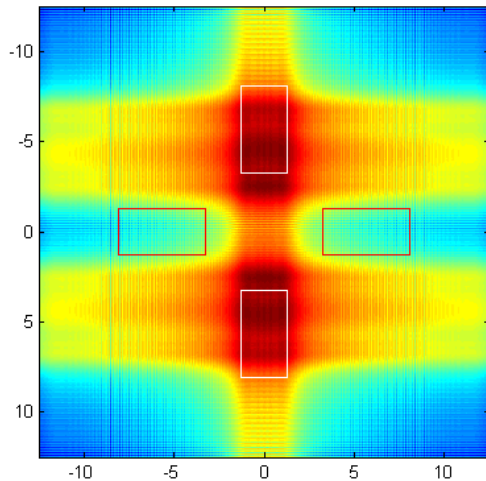


Cyclic error:

$$\varepsilon_{\lambda} = \frac{\lambda}{2\pi} \sqrt{\frac{P_{xtalk}}{P_{signal}}}$$

For example:

-90dB cross-talk results in
6.5pm

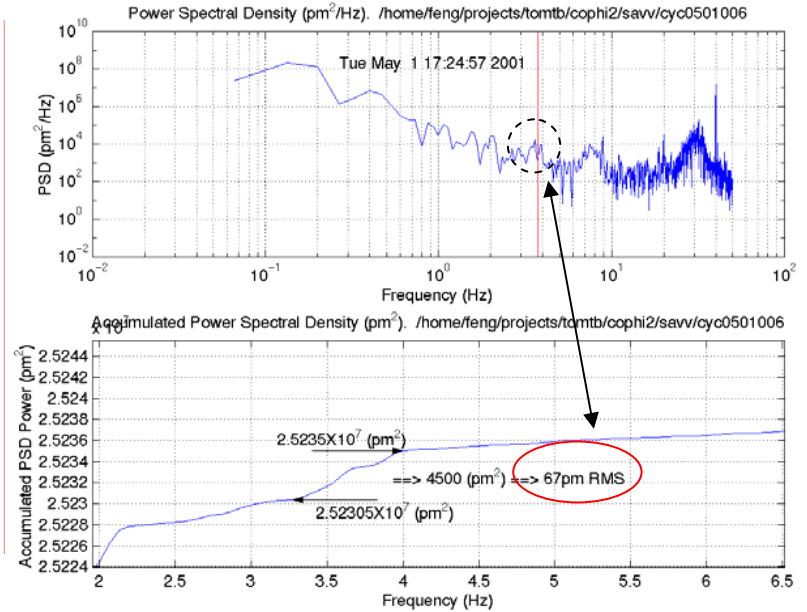
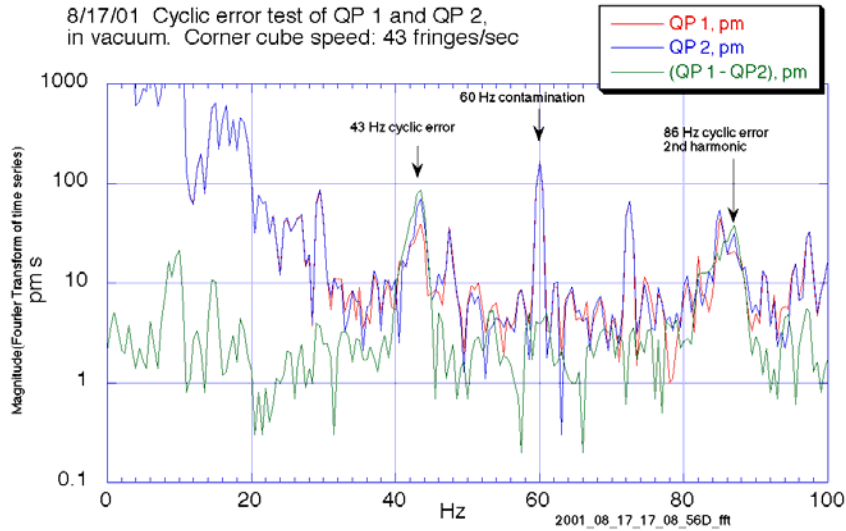


Mask patterns to minimize
diffraction cross-talk

Common sources of cyclic
error:

- Optical cross-talk:
 - Diffraction
 - Ghost reflections
 - Scattering
 - Polarization leakage (commercial DMI)
- Electrical cross-talk
 - EMI in AOM RF drivers
 - EMI between MEAS and REF signals
 - Ground coupling

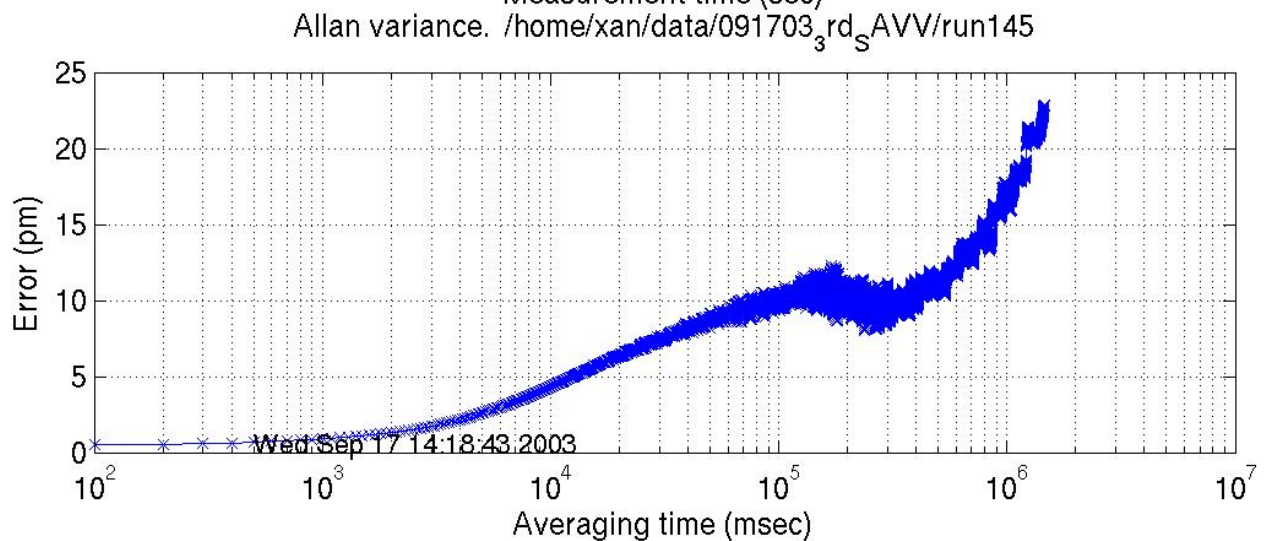
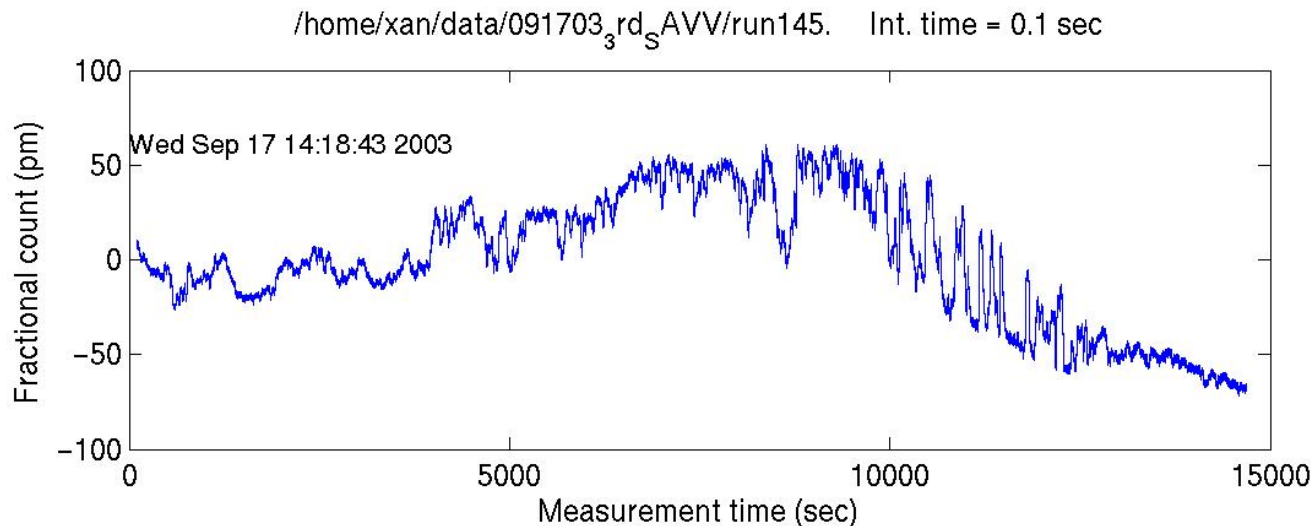
F. Zhao, et al, "SIM internal metrology beam launcher development," *Proc. SPIE* vol.4852, pp. 370-379, 2002.



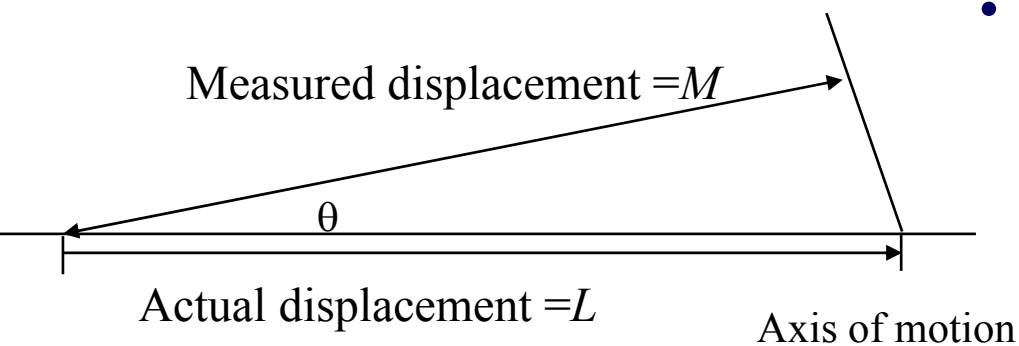
QP2	QP3	QP8	QP12	QP16	SAVV	Milestone #1
44	16	45	28	29	67	100 (pm, RMS)



Environmental Errors – SAVV Beam Launchers

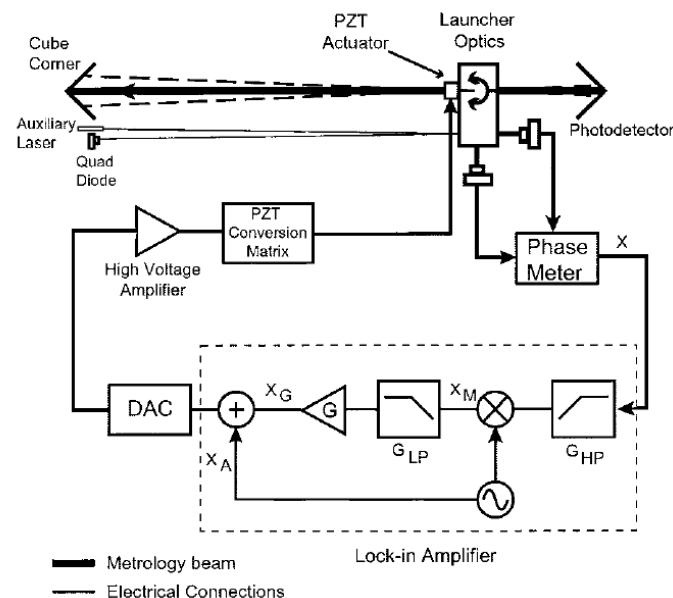


- Measurement along a non-parallel axis

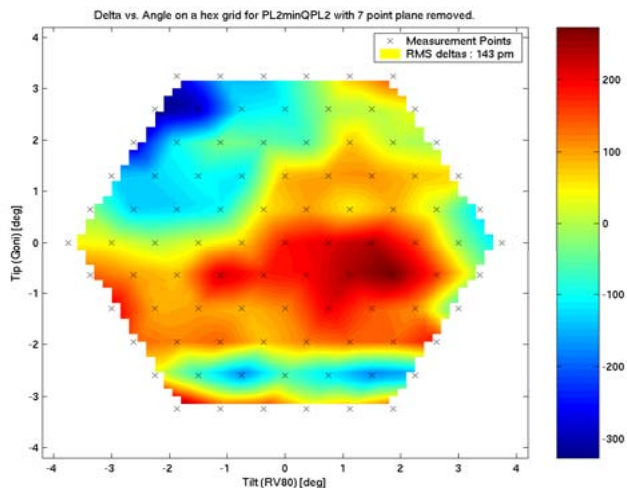


$$\varepsilon = L - M = L(1 - \cos \theta) \approx \frac{1}{2} L \theta^2$$

- Solution: dither in laser beam pointing, lock to minimum optical path length
- Demonstrated $\sim 0.5 \mu\text{rad}$ pointing accuracy in SIM



Logan, J.A., et al, "Automatic Alignment of a Displacement-measuring Heterodyne Interferometer," Appl. Opt., 41, p. 4316 (2002)



Narrow Angle (90% conf.)

[pm]

SIM Goal

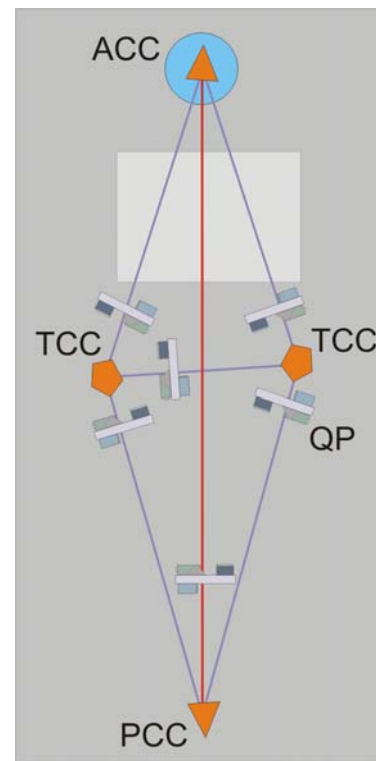
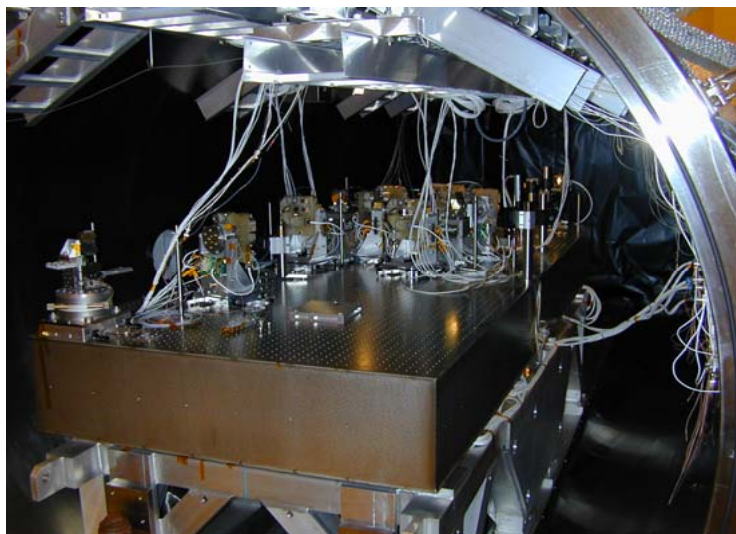
8

Without optical averaging

23

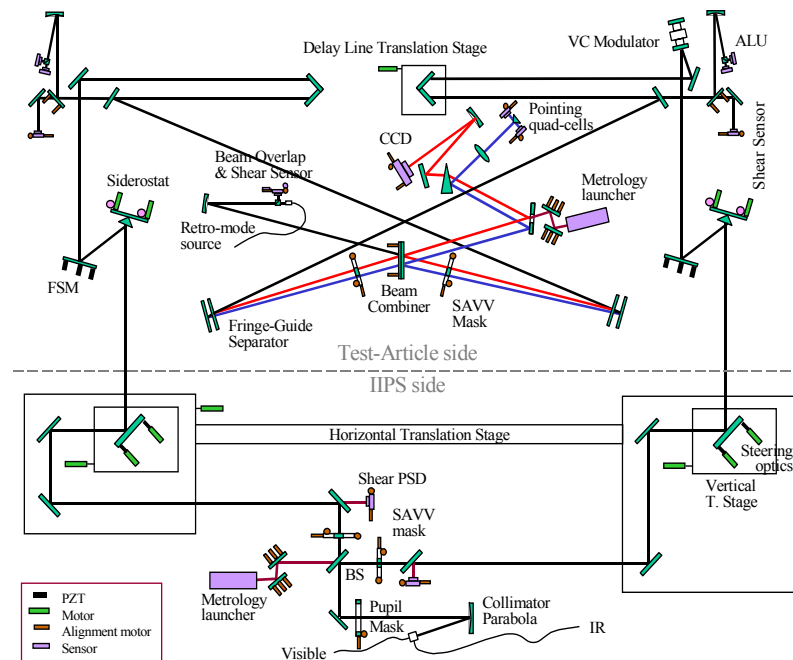
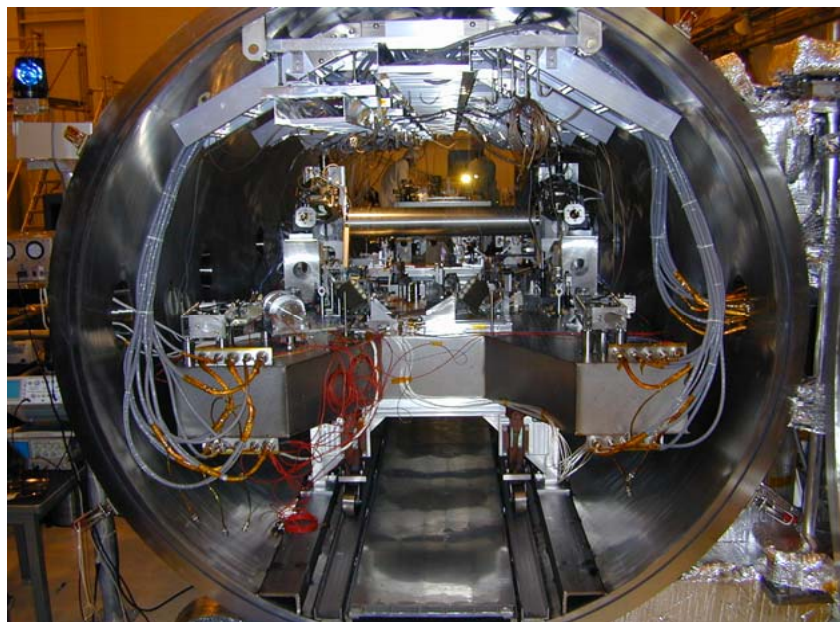
With optical averaging

11

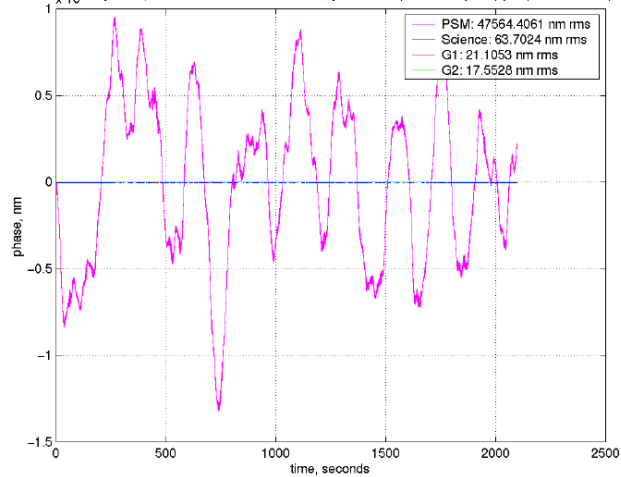


Narrow Angle Requirements

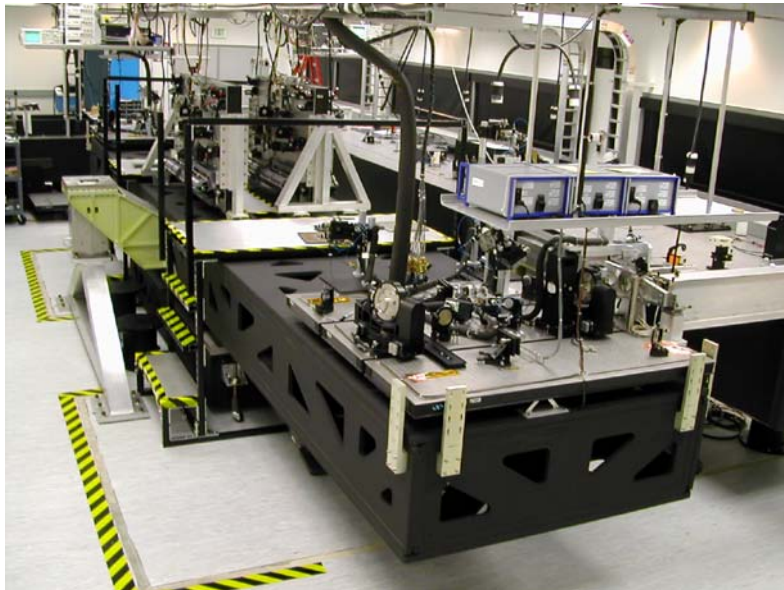
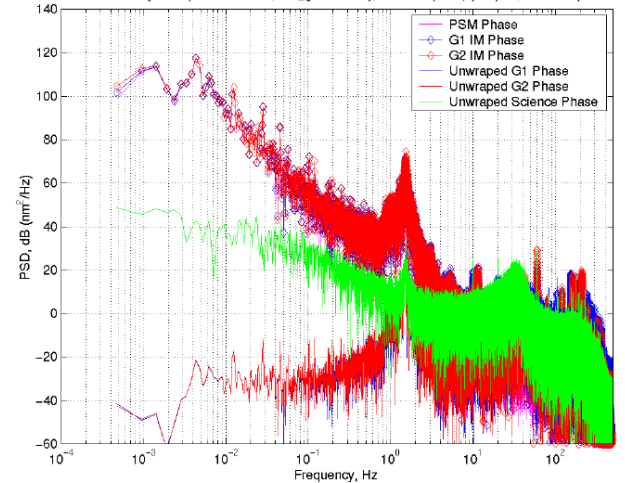
MAM	Goal	Baseline		Testbed
Astrometry	1.1	3.0	uas	Results
FD	17.9	99.0	pm	
FI	16.0	38.0	pm	13.7



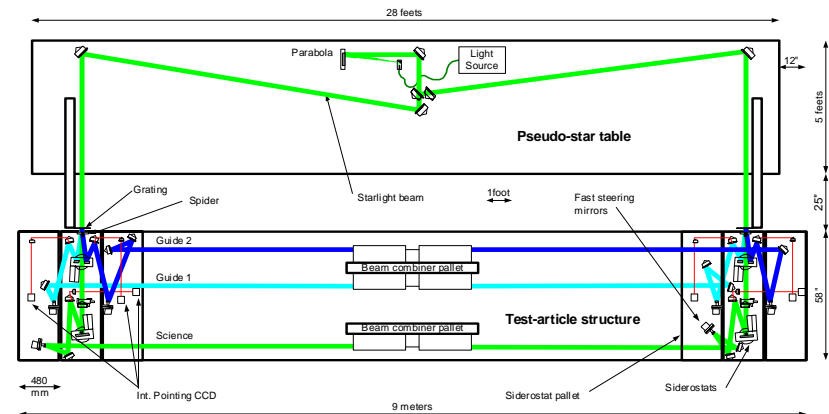
test021302_0.mat: PSM OPD, Science, G1 & G2 Unwrapped Phase while on PFF Mode
Disturbance=Z-rot by ACS (On-Orbit ACS disturbance, Gui_gain = 0.092), 1938000 point(s) skipped, 2097152 points usec.



test021302_0.mat: Power Spectra for, PSM, G1, G2 & Science Phase while on PFF Mode
Disturbance=Z-rot by ACS (On-Orbit ACS, Gui_gain=0.092), 1938000 point(s) skipped, 2097152 points usec

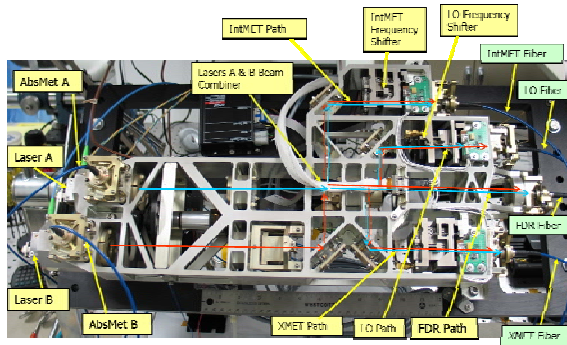


Starlight system optical layout - phase 2

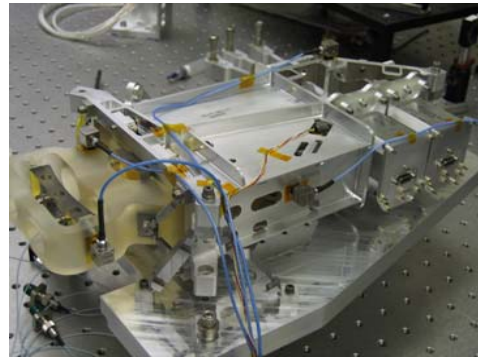


Performance demonstrated after environmental tests (random vibe, thermal vac)

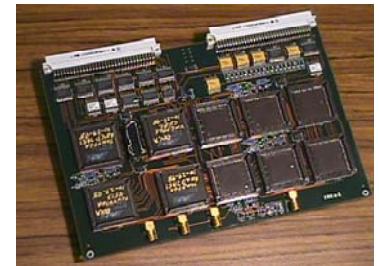
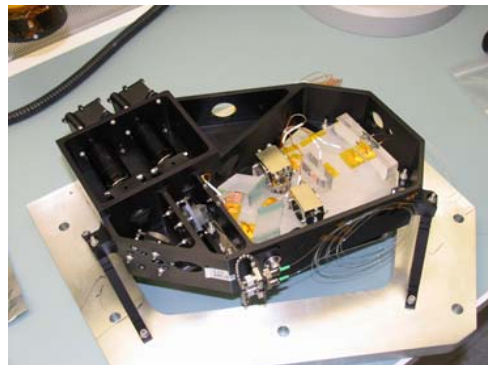
Metrology source brassboard



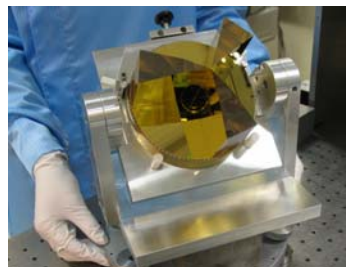
External beam launcher brassboard



Internal beam launcher brassboard



Phase meter electronics



Metrology fiducials

- Metrology performance ranging from mm to pm
- Various technical maturity
 - Sub-millimeter flight technology demonstration example SRTM
 - Nanometer and picometer component technology development is complete -- all components have achieved TRL-5 (SIM)
 - Nanometer closed-loop control system demonstration is complete -- at TRL-6 (SIM)
 - Picometer sensing at system level demonstration is completed -- TRL-6 (SIM)
- Design tools (such as diffraction code), test/diagnostic tools established