



Recent Small Asteroid and Boulder Radar Observations

Michael Busch, for the Asteroid Radar Team

Review of Radar Astronomy



Goldstone 70-m



Goldstone 34-m antenna



Arecibo 305-m



Klystron Tube; astronomer for scale

Small Objects

- Range resolution:
 - 7.5 m Arecibo.
 - 3.75 m Goldstone, upgrade to < 2 m pending.
- Can get equatorial dimensions to < ±2 m given spin period, pole constraint.
- Best results with several observations from multiple directions; requires rapid scheduling.
- Speckle tracking helpful only for strongest targets.

Small Objects



2013 EC20

- Discovered 2013 March, observed with Arecibo while outbound.
- 2-3 m diameter, 10-30 ton mass.
- High-albedo silicates.
- <u>Too small to be worth retrieval?</u>
- Most recent small target: 2014 FO38
 - Observed last week with Arecibo.
 - Spin period ~5.4 min.
 - d = 10-15 m, based on 1 day of observations.
 - Not a retrieval target.

Boulders

- Seen on ~10 asteroids observed with radar; on all visited with spacecraft.
- More visible in higher-resolution and higher-SNR images.
- Not identified on objects <80 m in diameter (may still be present).









Image from Chang'e 2 *(Huang et al.)*

<u>Toutatis</u>

Images from Goldstone.

Asteroid is 4.5 km long. Some boulders visible here also show up in Chang'e 2 pictures.



A Thought On Cohesion

- Objects appear to be gravity-dominated above ~200 m diameter. On 200 m object, g ~0.03 mm/s². Gravitational force on blocks on surface: ~0.03 N / ton.
- Spin rates of some ~30 m objects are ~300 s. a ~5 mm/s². Cohesive force on the outside of such a pile: ~5 N / ton. Larger objects fracture along weak zones.
- Full shapes of blocks, actual cohesive forces, and fracture conditions unpredictable.

Picking up a block will require a lot more force than bagging one.