

Exoplanets from Sounding Rockets: The PICTURE Experience and Some Random Thoughts

Supriya Chakrabarti
Center for Space Physics
Boston University

The Team

- **Boston University**
 - John Atkinson
 - Thomas Bifano
 - Supriya Chakrabarti
 - Timothy Cook
 - Paul Jung
 - Nicholas McConnell
 - Donald Wroblewski
 - Brian Hicks
 - Christopher Mendillo
- **JPL**
 - Francisco Aguyo
 - Randy Bartos
 - Gary Brack
 - Jim Moore
 - Eduoard Schmidlin
 - Michael Shao
 - B. Martin Levine
 - Shanti Rao
 - J. Kent Wallace
- **Boston Micro Machines**
 - Paul Bierden
- **NASA WFF**
 - Ted Gass
 - John Gsell
 - Valerie Gsell
 - Matt Peterson
- **MIT**
 - Geoff Crew
 - Rick Foster
 - Bob Goeke
 - Dorothy Gordon
 - Jimmy O'Connor
 - Matt Smith
- **GSFC**
 - Scott Antonille
 - David Content
 - Joseph Davila
 - Douglas Rabin
 - Tom Wallace
- **Draper laboratory**
 - Benjamin Lane
- **Northrup Grumman**
 - Ronald Polidan
 - Rocco Samuele
- **Lockhead Martin**
 - Andy Clavins
 - John Miles
 - Bill Sable
 - Dominic Tennerili

When normalizing to the nominal values for a discovery class mission: $D=1.5$ m, $T=1$ h, $Eff=5\%$ and $sl=10^{-9}$ (assuming an integrated null depth of $1e-6$), we get for a Jupiter analog orbiting a Sun-like star at a distance (Dist) of 10 pc:

$$SNR = 14.2 \left(\frac{D_{in.m}}{1.5} \right) \cdot (T_{in.h})^{0.5} \cdot \left(\frac{Eff}{0.15} \right)^{0.5} \cdot \left(\frac{10^{-9}}{sl} \right)^{0.5} \cdot \left(\frac{10}{Dist_{in.pc}} \right) \cdot \left(\frac{1.4}{Reso} \right)^{0.5} \quad (8)$$

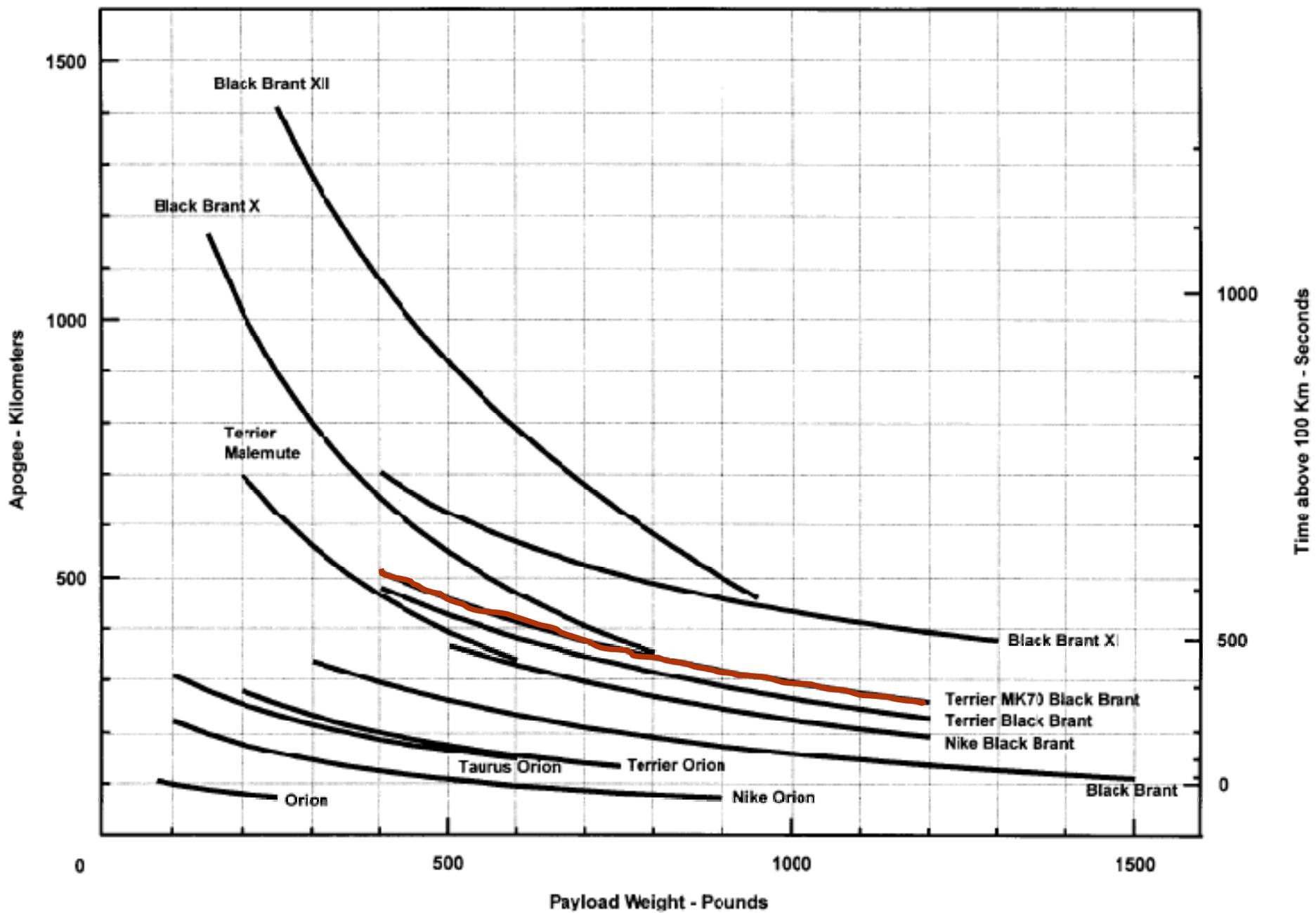
Looking at the general case of an exo-Jupiter with a given separation Sep in A.U. from a central star with flux F^* , and assuming pure stellar reflection with Jupiter's albedo, we get:

$$SNR = 14.2 \left(\frac{D_{in.m}}{1.5} \right) \cdot (T_{in.h})^{0.5} \cdot \left(\frac{Eff}{0.15} \right)^{0.5} \cdot \left(\frac{10^{-9}}{sl} \right)^{0.5} \cdot \left(\frac{10}{Dist_{in.pc}} \right) \cdot \left(\frac{1.4}{Reso} \right)^{0.5} \cdot \left(\frac{5.2}{Sep} \right)^2 \cdot \left(\frac{F^*}{F_{Sun}} \right)^{0.5} \quad (9)$$

and the integration time T_{int} for broadband detection of exo-Jupiters is:

$$T_{int} = 7 \text{ min} \cdot \left(\frac{SNR}{5} \right)^2 \cdot \left(\frac{1.5}{D} \right)^2 \cdot \left(\frac{0.15}{eff} \right) \cdot \left(\frac{sl}{10^{-9}} \right) \cdot \left(\frac{Dist_{in.pc}}{10} \right)^2 \cdot \left(\frac{Reso}{1.4} \right) \cdot \left(\frac{Sep}{5.2} \right)^4 \cdot \left(\frac{F_{sun}}{F^*} \right) \quad (10)$$

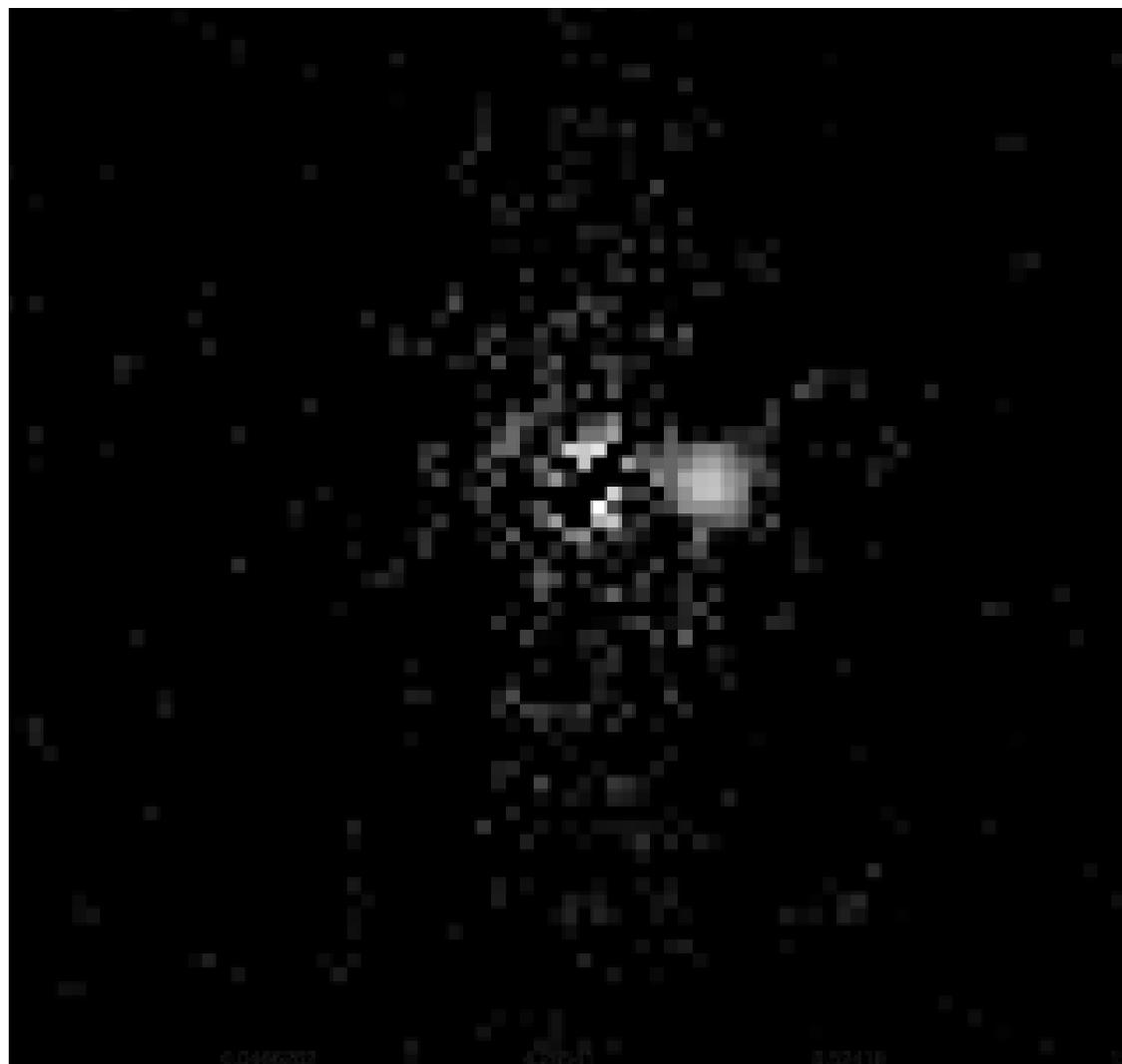
One can note the strong dependency on the planetary separation. A Jupiter “analog” orbiting at 2 AU from a solar type star 10 pc away can be detected at 5σ in less than a minute.



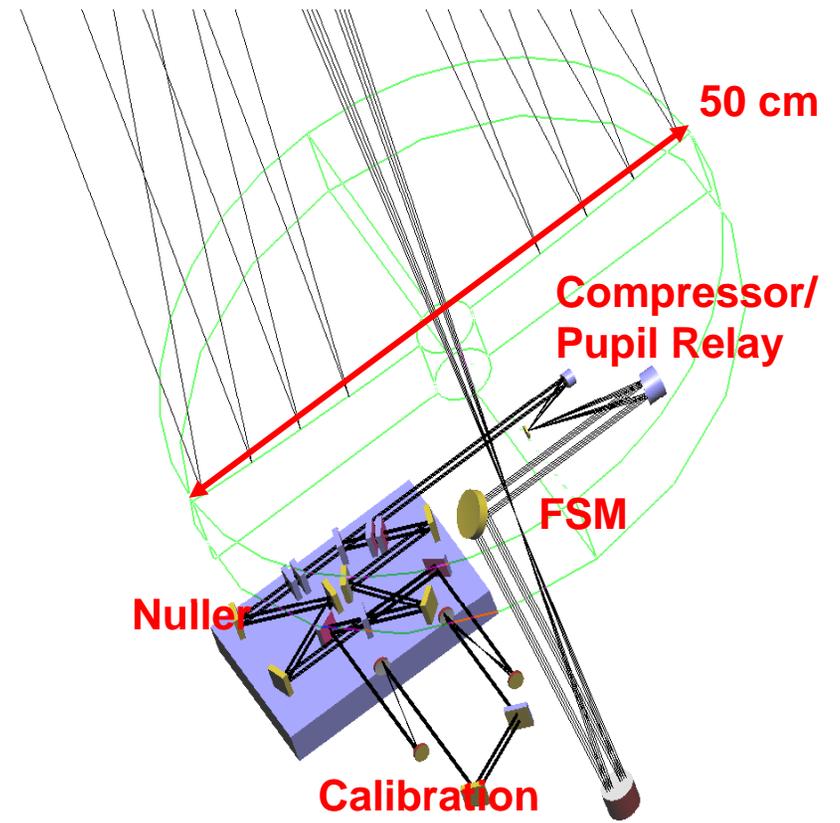
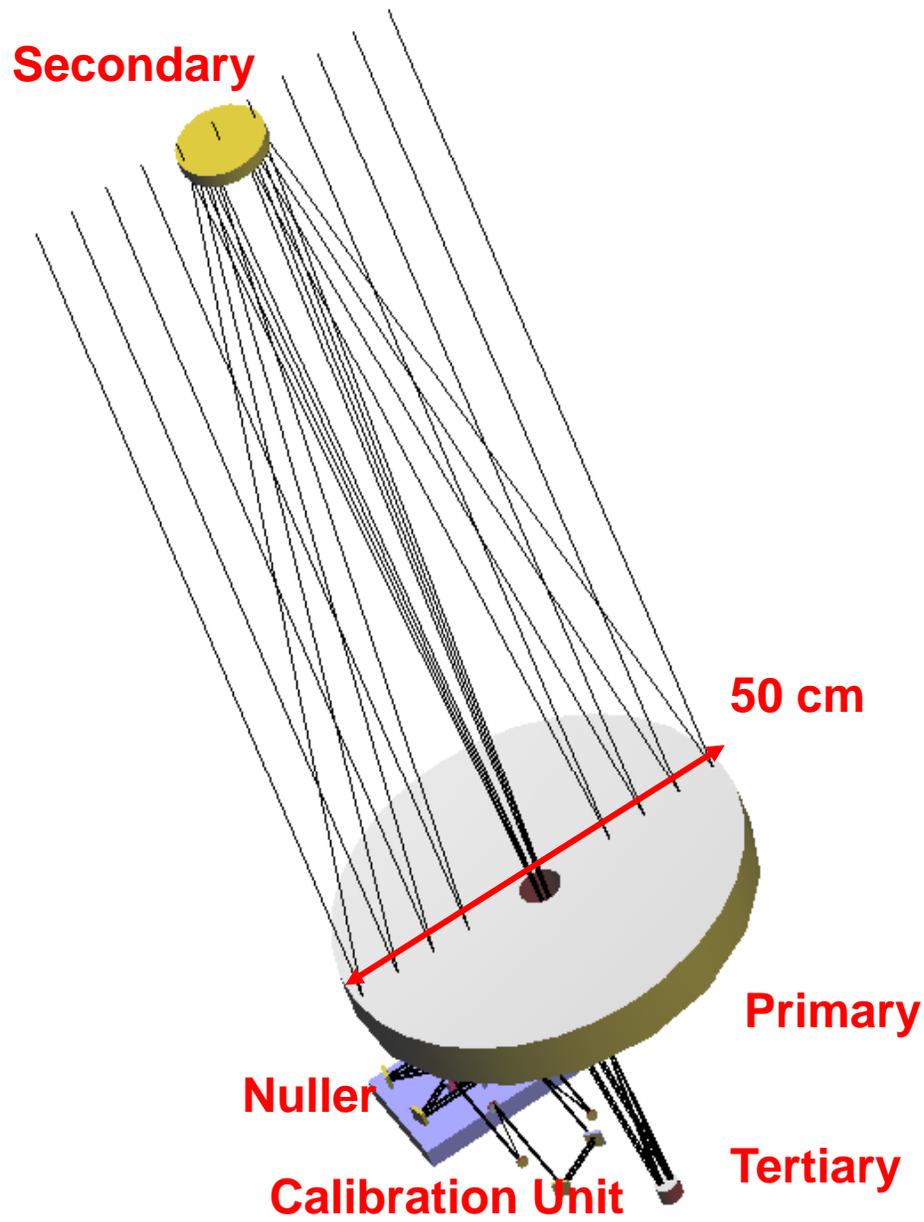
PICTURE Plans

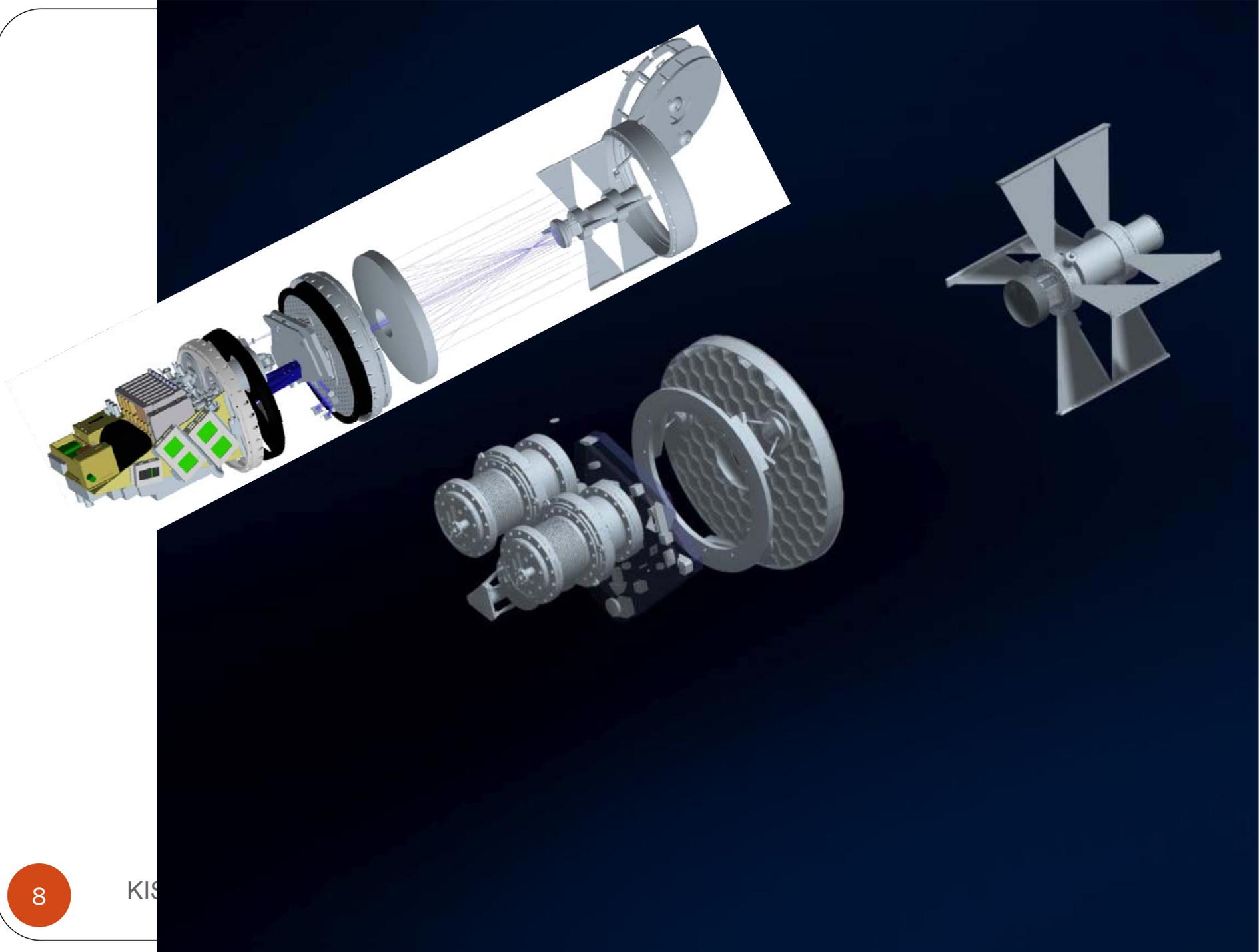
Event	Time (sec)	Altitude (km)
Launch	0	0
Burn out	57	70
Separation	75	100
ACS on Rigel	105	165
End in-flight Cal.	145	250
ACS on ϵ ERI	165	340
Apogee	340	425
Shutter Close	620	100

Expected Image



Optics Overview

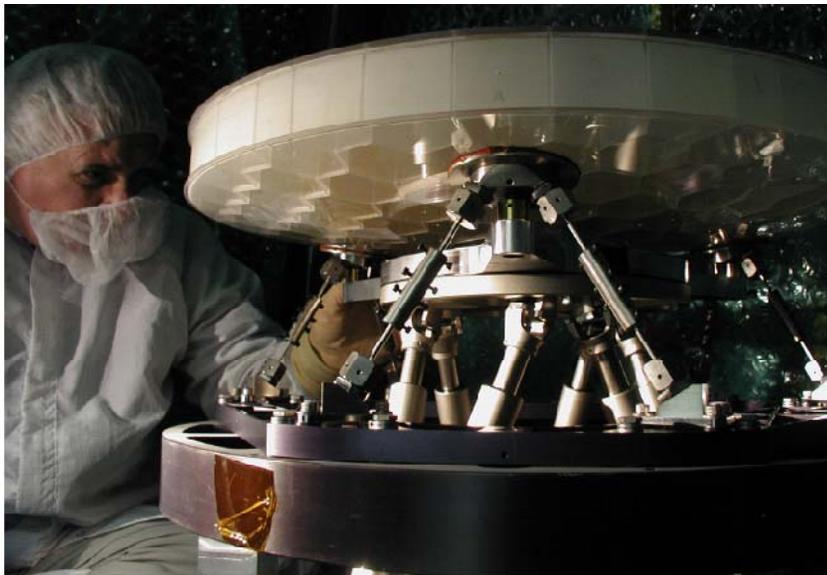




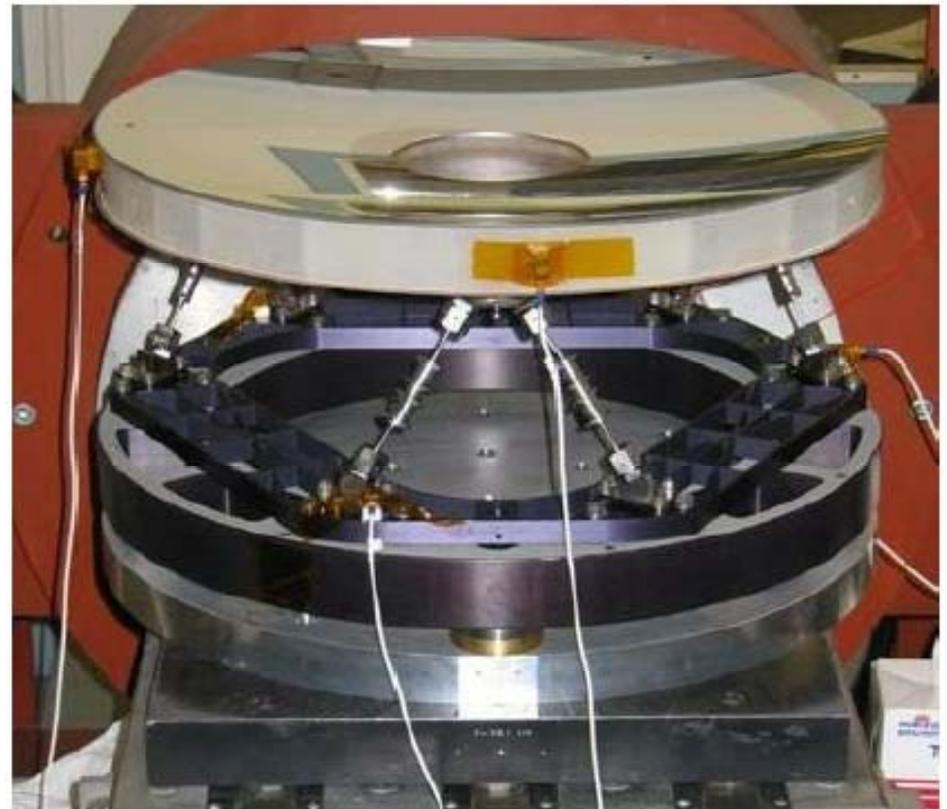
The PICTURE primary mirror

Solar High Angular Resolution Photometric Imager (SHARPI)

– designed for 0.2" res. over 70" FOV at 120 nm (7 nm rms figure error)

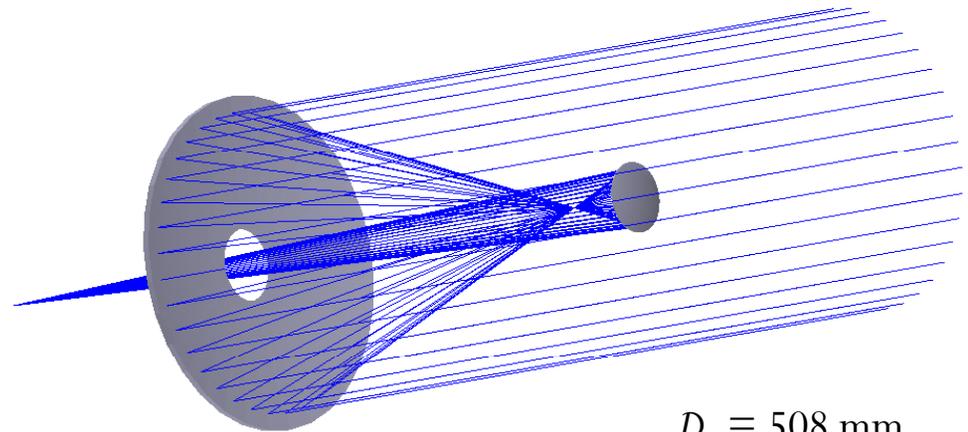


- Protected silver coating
- Lightweight (~ 5 kg)



The PICTURE telescope

- Gregorian design ($k_p = -1$, $k_s = -0.66$)
 - invar, Ti, carbon fiber structure



$$D_p = 508 \text{ mm}$$

$$F_p = 1.26$$

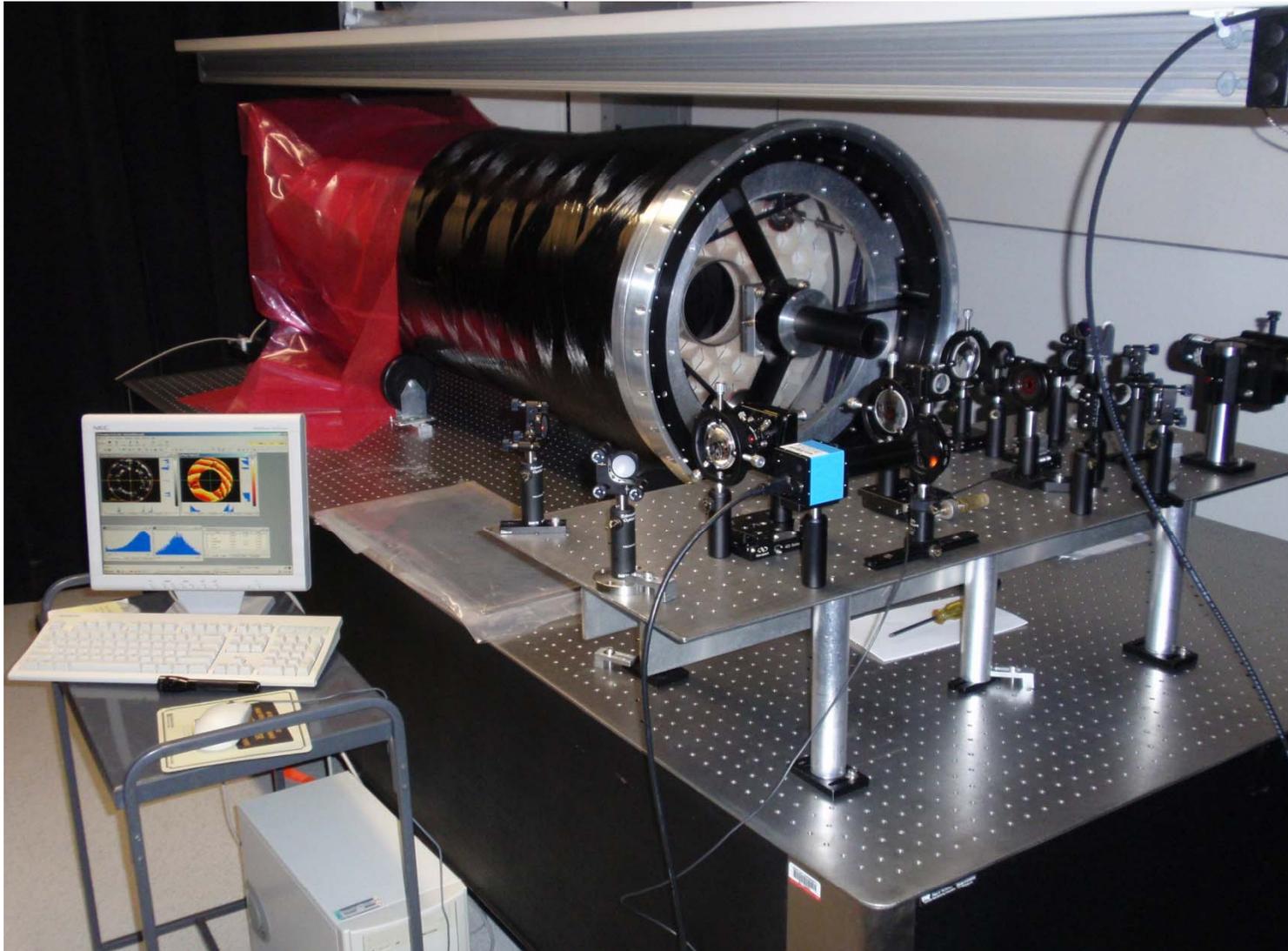
$$F = 12.3$$

$$S = 763 \text{ mm}$$

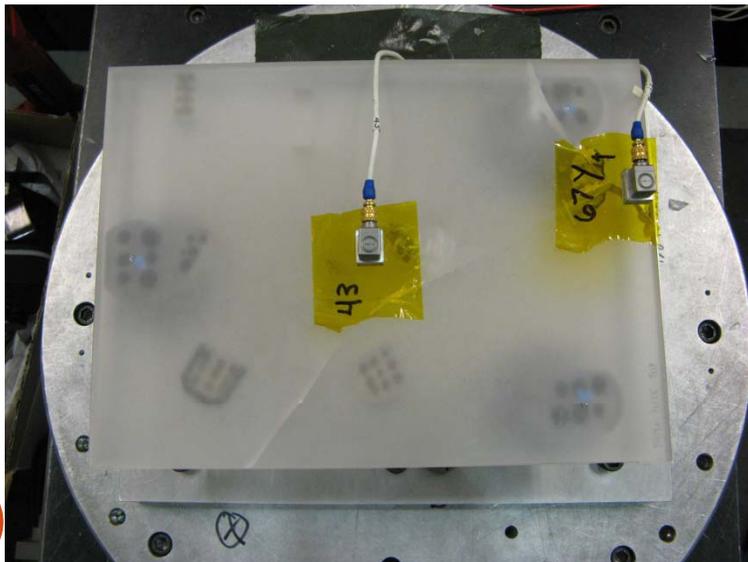
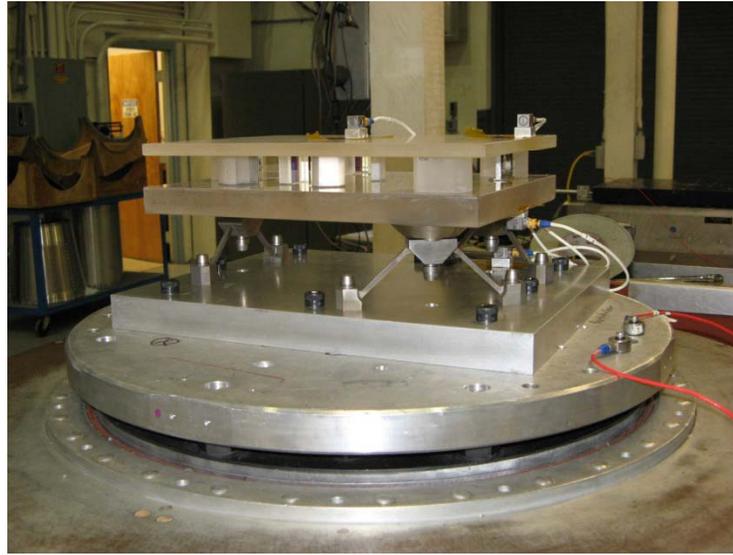
$$B = 459 \text{ mm}$$

$$\varepsilon = 0.25$$

Horizontal alignment in 1g



Development and tests of the nuller



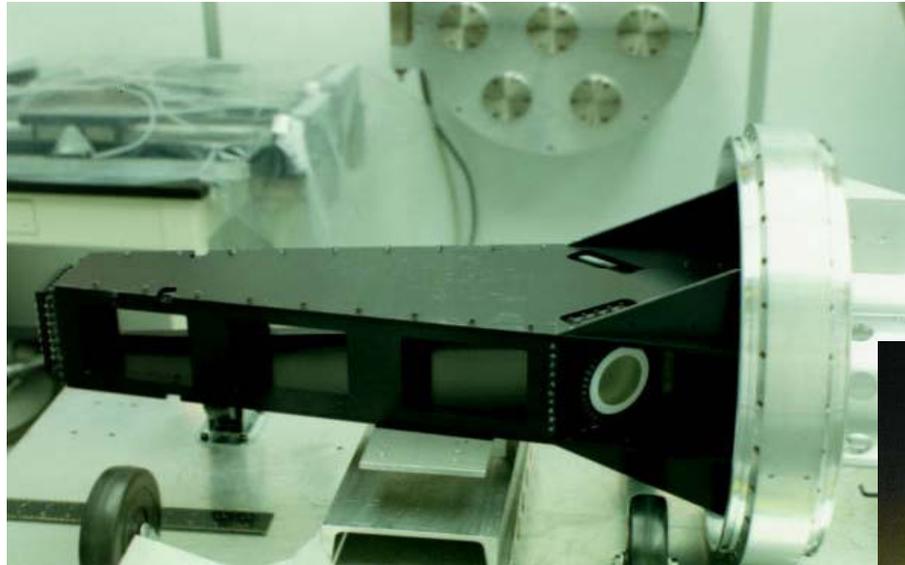
PICTURE Current Status

- Started 2005
- Met with NASA Wallops – people responsible for rocket support
 - Almost got cancelled – our required pointing stability (0.2”) unavailable
- The team relaxed pointing requirement to 2”
- Telescope assembly begins at GSFC – unfortunately, runs out of funds
- 2008: Planet not observable until 2013-4 – new science: exo-zodi
- Telescope aligned by BU after building necessary test equipment
- NASA Wallops flight-demonstrated sub-arc-sec pointing
- Nuller developed, tested and delivered by JPL in Sept, 09
- Integration on-going, hoping to launch in 2010

Now, some random thoughts

.... On Enabling Technologies

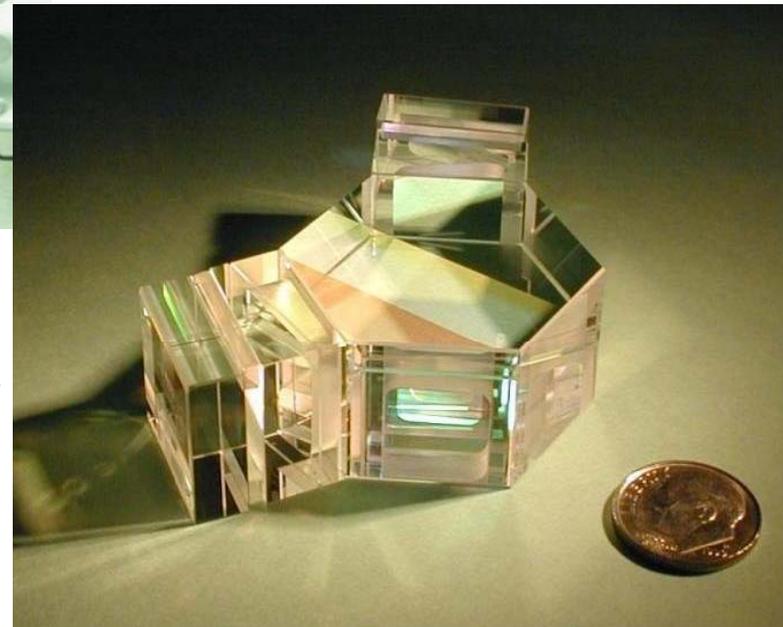
SCARI/SHS to SHIMMER



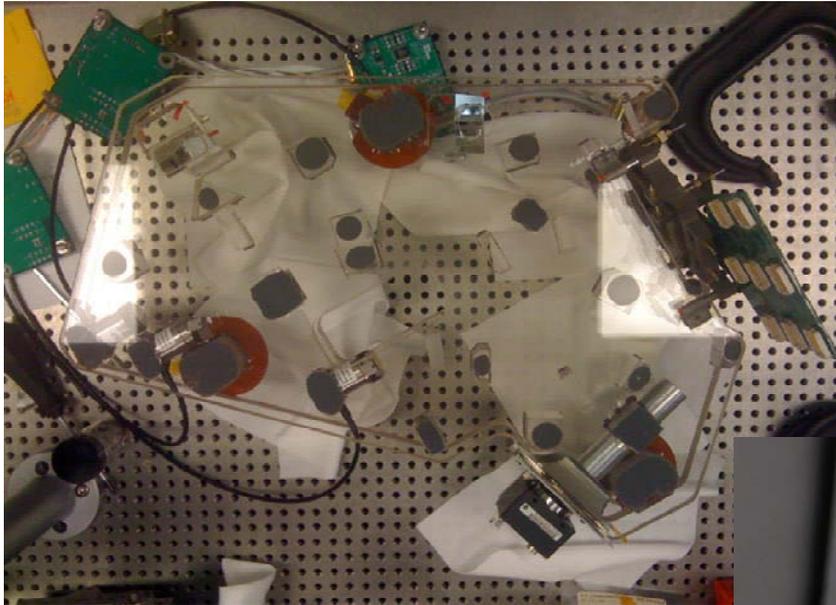
← Self Compensating All-Reflection Interferometer (SCARI) - 1216Å
(Spatial Heterodyned Spectrometer)

SHIMMER (Spatial Heterodyne Imager
for Mesospheric Radicals):

Monolithic construction
Suitable for high resolution spectroscopy
Flew on STS-112 in 2002



PICTURE Nuller to MANIC



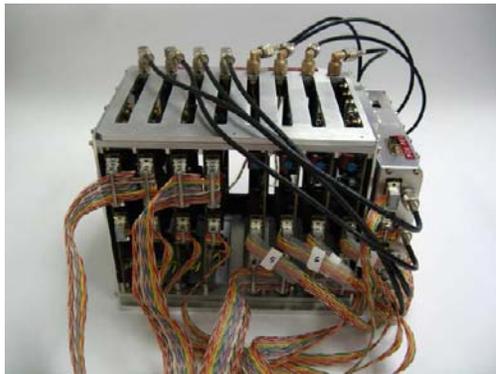
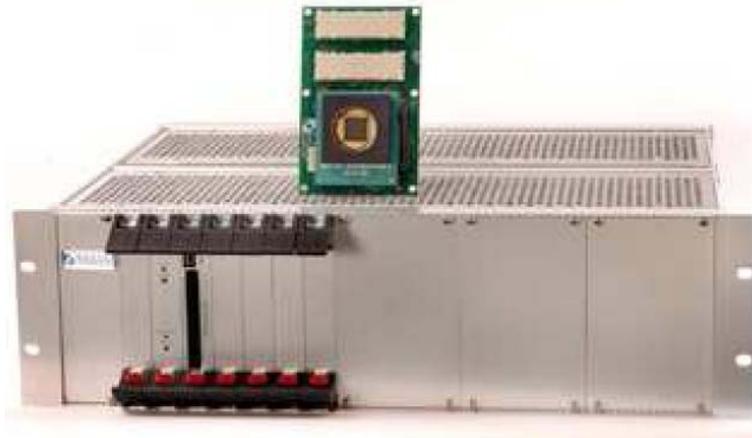
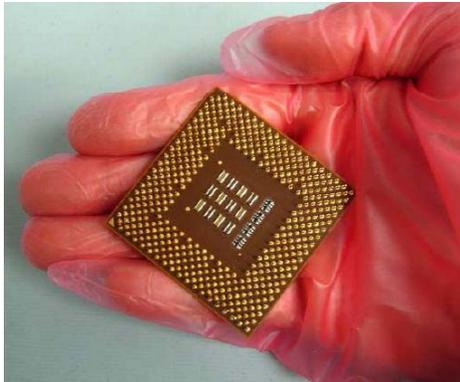
- Rotationally shearing nulling interferometer

Monolithic achromatic nulling interference coronagraph: design and performance

Brian Hicks,^{1,*} Timothy Cook,¹ Benjamin Lane,² and Supriya Chakrabarti¹
10 September 2009 / Vol. 48, No. 26 / APPLIED OPTICS 4963



Another Key technology Development



**A Novel ASIC
Integrating Latching
(NAIL) deformable
mirror controller**

**A NASA TDEM
proposal currently
under consideration**

Finally, Extended Duration Sounding Rocket: Yet Another Technology on the Horizon

- **Experiment Mass:**
 - ~ 150 kg (without subsystems)
- **Experiment Volume:**
 - 2.2 m length, 0.56 m diameter
- **Data Rate (primary science):**
 - 0.5 -- 1 Gbyte day⁻¹
- **Experiment Power (orbit avg)**
 - 3.25 Amps @ 28 Volts
- **Pointing**
 - 1" -- 2" stability per orbit
 - 1' accuracy
- **Duration**
 - 1 day for minimum success
 - 1 month for comprehensive success



Table 2 – Recurring Cost Estimates

SYSTEM	RECURRING COST
WFF Support Systems	\$ 3,300 K *
Launch Vehicle & Range Cost	\$ 10,000 K **
Experiment Cost (mod for orbital flight)	\$ 1,000 K
On-Orbit Ops Cost	\$ 100 K
Misc / Contingency	\$ 1,000 K
TOTAL	\$ 15,400 K

*Cost Based on WFF OSR Study, 1/15/2009. Development+1st flight demonstration = \$26M

**Launch vehicle cost based on Space X response to inquiry.

Two-burn Performance to LEO

