Ultraviolet Observations for Lunar Volatile Detection and Mapping

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KISS: New Approaches to Lunar Ice Detection and Mapping

Ultraviolet defined here as FUV through UVA measurements

- FUV ("vacuum UV"): 100-200 nm
- UVV:200-280 nm
- UVB: 280-315 nm
- UVA: 315-400 nm



Perovich & Govoni (1991)

Cumulative Polar Nightside Lya Maps



• PSRs have substantially darker FUV albedos than their surroundings

From Gladstone, 2011

Diviner maximum bolometric surface temperature (E. Sefton-Nash, LPSC 2013)



Hayne, 2013



Hayne, 2013

LAMP and Diviner Overlay



Hayne, 2013

South Pole



Top layer: LAMP Ly-alpha Albedo

South Pole



Top layer: LAMP Ly-alpha Albedo

Example of UV/Vis gas emission sensitivities

The Gas Emission Line Brightness for several Required Species and NEP (Ra/nm)



LADEE UVS Observations of OH & H2O+



$\Delta t x # coadded$	For SNR=5 (#/cm ²)	For SNR=5 (#/cm ³)		
1x1	5.48E+11	12894.78		
1x10	5.48E+10	1289.48		
10x10	1.73E+10	407.77		
10x100	5.48E+09	128.95		
10x400	2.74E+09	64.47		

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	For SNR=5	For SNR=5	
$\Delta t x \# coadded$	(#/cm ²)	(#/cm ³)	
1x1	1.80E+11	4246.51	
1x10	1.80E+10	424.65	
10x10	5.71E+09	134.29	
10x100	1.80E+09	42.47	
10x400	9.02E+08	21.23	

Possible Applications

- FUV: Similar to LAMP, using Lyman-a as a source to see in the dark, but for longer times for better SNR.
- Dedicated small orbiter or lander/rover
- Monitor impacts (e.g., LRO LAMP)

UVC/B: Monitor for OH mobility

- Beyond LADEE lower altitudes and higher latitudes
- Dedicated small orbiter or lander/rover
- Monitor impacts (e.g., LCROSS VSP)

Strength:

• Good way to look for mobile OH/H2O

Weakness:

• Surface reflectance may be ambiguous