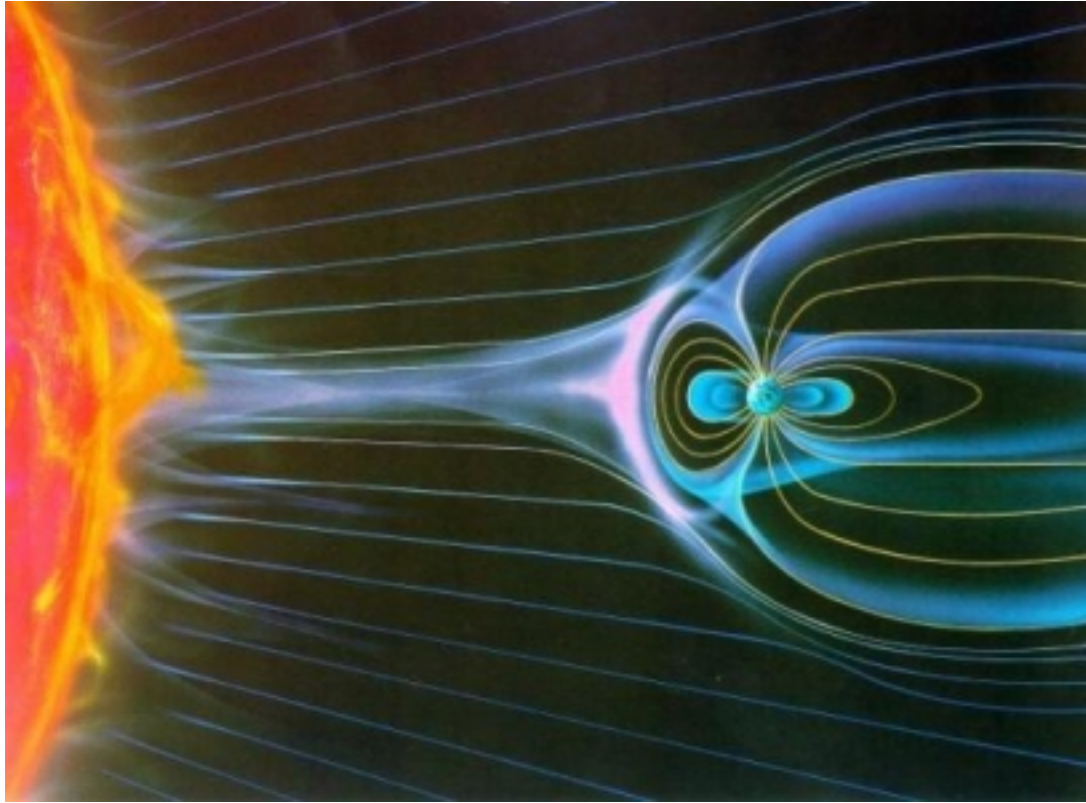


*Star-Planet Interactions Discussion*

# Lessons from Our Star and Solar System



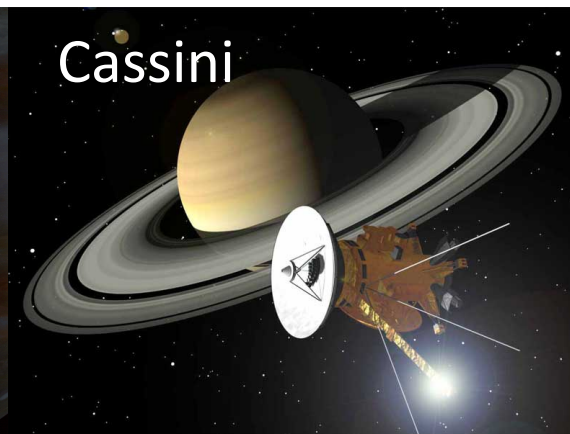
Paulett Liewer, Jet Propulsion Laboratory, California Institute of Technology  
Planetary Magnetic Fields: Planetary Interiors and Habitability  
Keck Institute for Space Studies,  
August 12 - 16, 2013





## **ANTENNAS IN SPACE!**

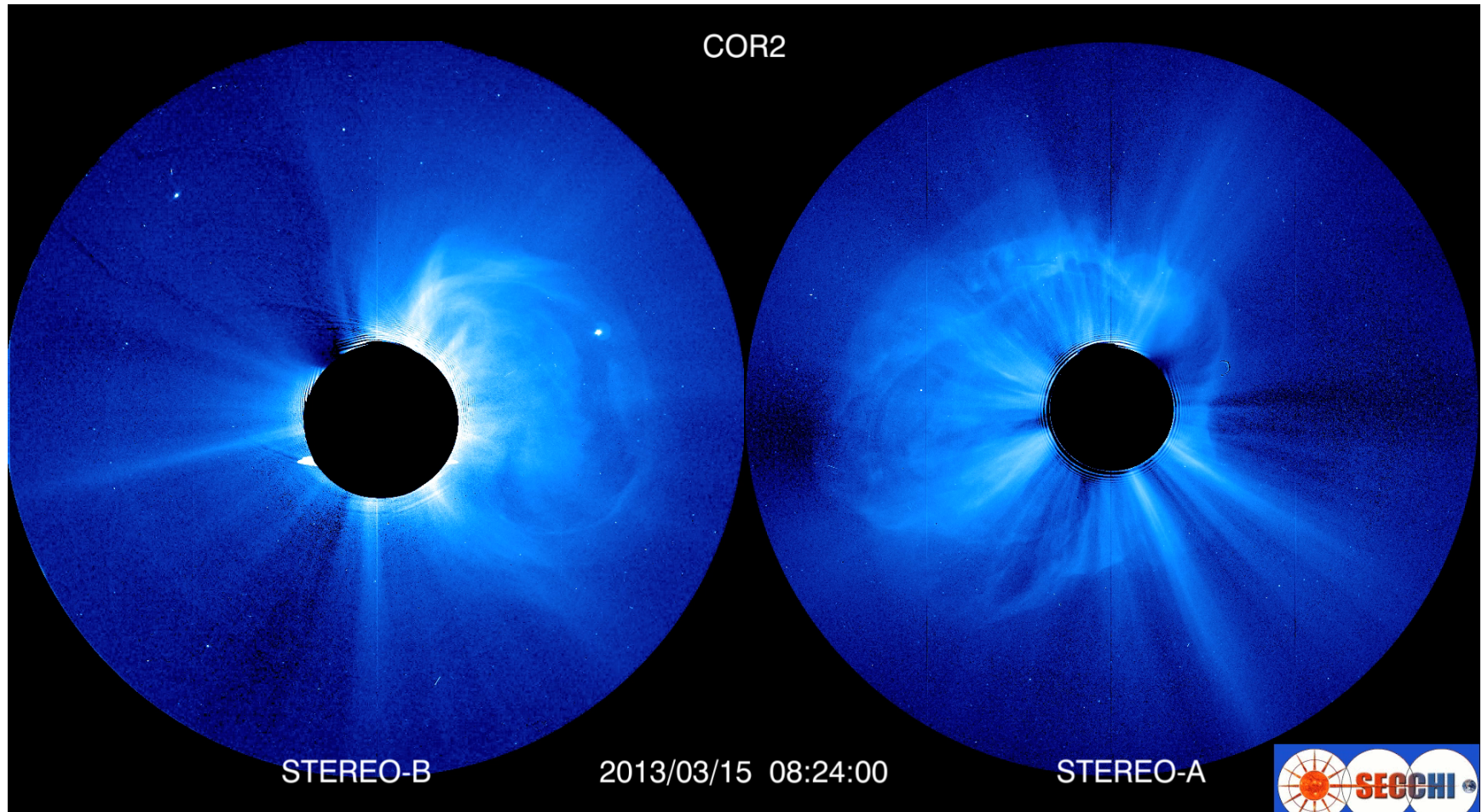
- Star-planet emission contrast in our solar system most favorable in radio[  $\sim 1$  for Jovian decameter]
- Jupiter's emission seen from Earth
- Voyager detected kilometric radiation from poles of Saturn, Uranus and Neptune





# STEREO views of Solar Corona

(add movie later)





# Solar Wind – Magnetosphere Interactions

*Unmagnetized flow impinging on magnetized obstacle*

- Radio and UV emissions from planets with magnetospheres increase with increased solar wind (SW )kinetic energy flux
  - Seen at Jupiter (Zarka&Genova 1983; Gurnett+ 2002; Hess+ 2012) & Saturn (Prange+2004)
- Process
  - SW energy dissipated in **magnetic reconnection** which **accelerates electrons**, leading to coherent **electron cyclotron maser emission** near and auroral ultraviolet emission in polar regions
  - Radiometric Bode's Law introduced in 1984 (Desch & Kaiser)



# Moon – Magnetosphere Interactions

*Magnetized flow impinging on an obstacle*

- Interaction of moons with planetary magnetospheres also causes strong radio and UV emissions
  - Seen at Jupiter & Saturn (*what about Neptune & Uranus?*)
- Magnetized obstacle (Jupiter-Ganymede; “dipolar”)
  - Continuous reconnection between 2 magnetospheres accelerates electrons causing CMI in auroral regions
  - *Analogy to Star-planet reconnection?*
- Unmagnetized obstacle (Jupiter- Io; “unipolar”)
  - Relative motion of moon and magnetosphere causes electron acceleration via electric field (waves or induced) causing electron cyclotron maser emission
  - *Analogy to unmagnetized hot Jupiters or planets around white dwarfs ?*



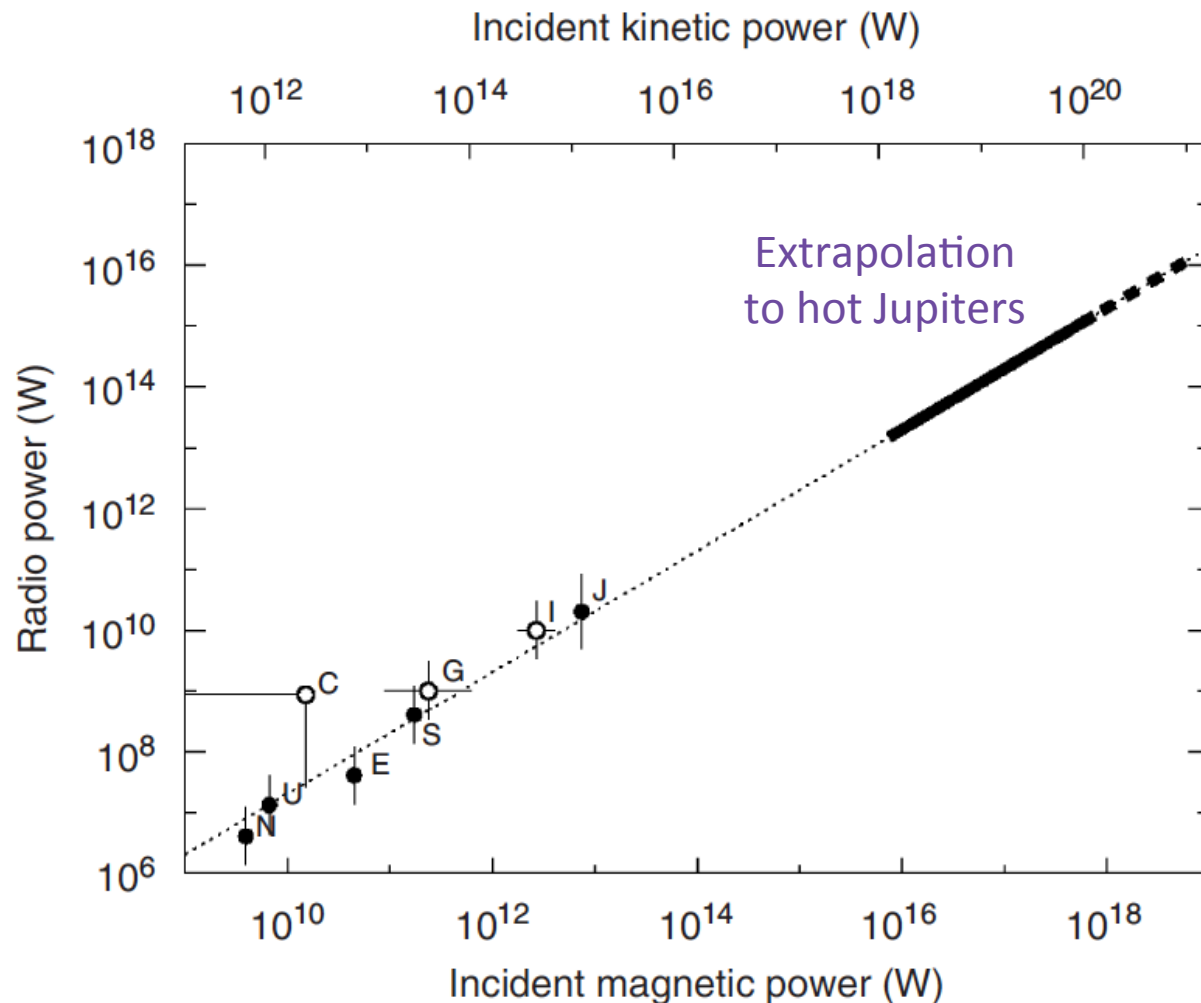
# Generalized Radio Bode's Law - Zarka (2007)

Radio power vs. incident kinetic power or Poynting Flux

- *Captures both types of interactions*
- *Kinetic-to-radio efficiency is  $10^{-5}$ , magnetic-to-radio efficiency is  $2 \times 10^{-3}$*

- E, J, S, U, N: 5 radio planets

○ Io, Ganymede, Callisto (upper limit)





# Extrapolating to Extra-solar systems

- Various other extrapolations from our solar system appear in the literature
  - Refs Farrell+ (1999); Lazio+ (2004), Hess & Zarka (2011)
  - How good are the extrapolations?
- What if a star has superflares?
  - Largest known Earth event: Carrington event  $\sim 10^{32}$  ergs
  - Interplanetary shocks lead to big enhancements in radio emission
  - Superflares ( $10^{33}$ - $10^{36}$  ergs) in Kepler data analyzed (Maehara +, Nature 2012 & Notsu+ Ap, 2013)
    - Found 365 superflares ( $>10^{33}$  ergs) on 148 solar-type star)
    - Statistics not consistent with hot Jupiter as cause
    - Stars appear to have larger starspots than our Sun