

Quantum Limits on Sensing and Imaging

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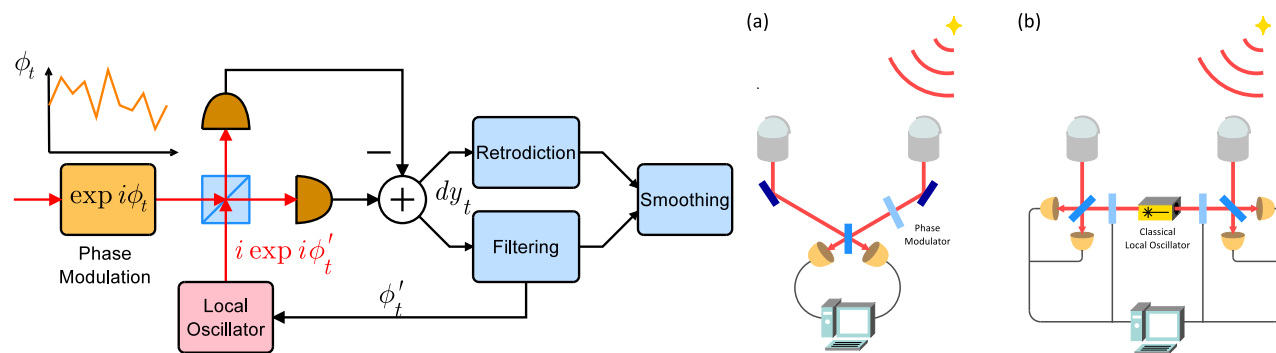
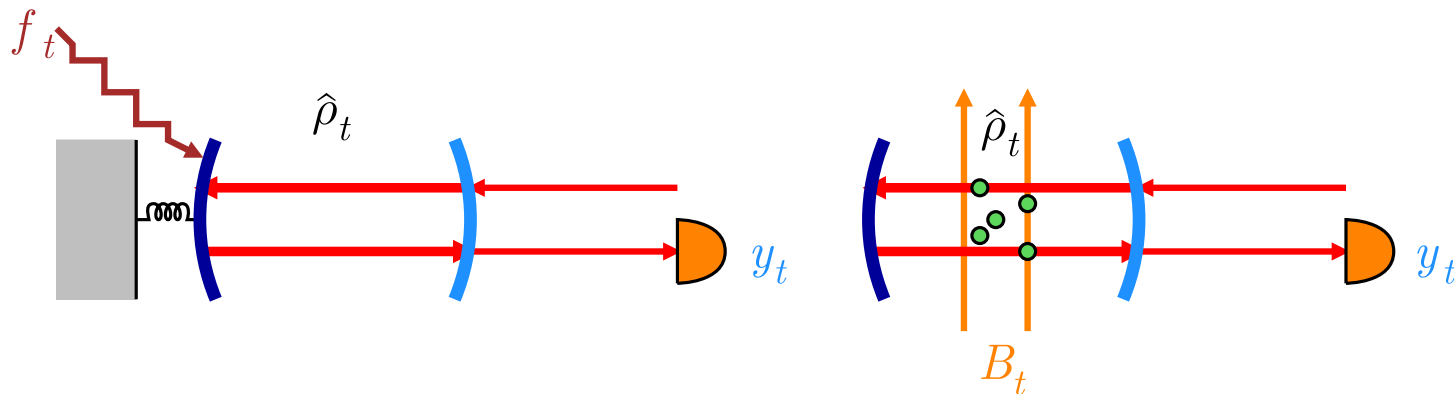
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Quantum Waveform Sensing



- Estimation/detection of **classical waveforms** $x(\mathbf{r}, t)$ using quantum systems
- Examples: optical interferometry, optical imaging, optomechanical force sensing (gravitational-wave detection), atomic magnetometry, gyroscopy, etc.

Optical Phase and Frequency Estimation

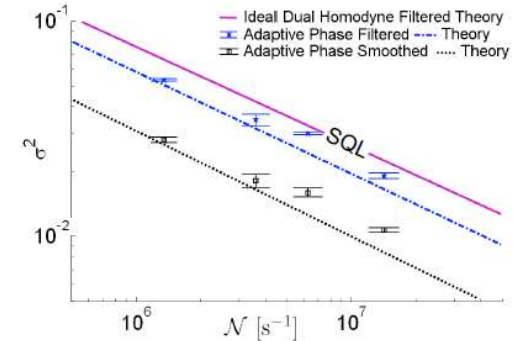
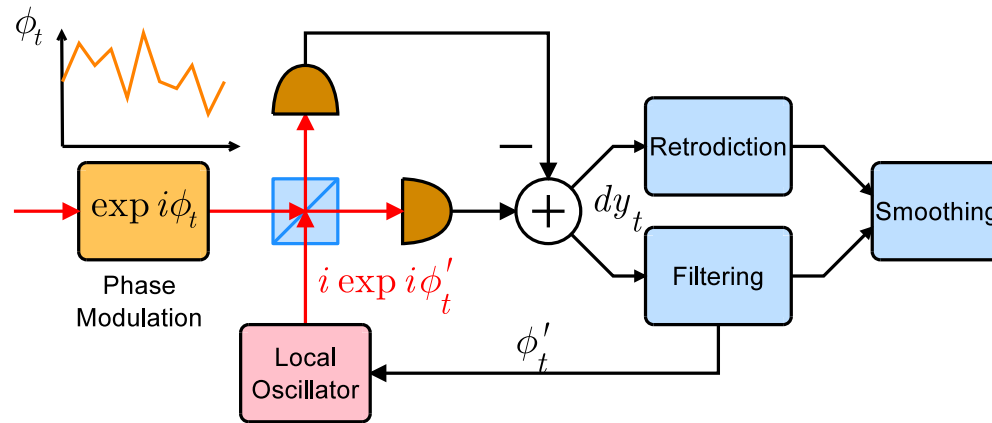


FIG. 3: The variance σ^2 of the adaptive phase estimation for quantum filtering and smoothing as a function of photon number \mathcal{N} , compared to the relevant theoretical predictions, and the theoretical predictions for nonadaptive measurements.

QCRB [Tsang, Wiseman, and Caves, PRL **106**, 090401 (2011); unpublished]:

$$\phi(t) = \int_{-\infty}^{\infty} dt' h(t-t')x(t'), \quad \langle \delta x^2 \rangle \geq \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{1}{4|h(\omega)|^2 S_{\Delta \hat{I}}(\omega) + 1/S_x(\omega)},$$

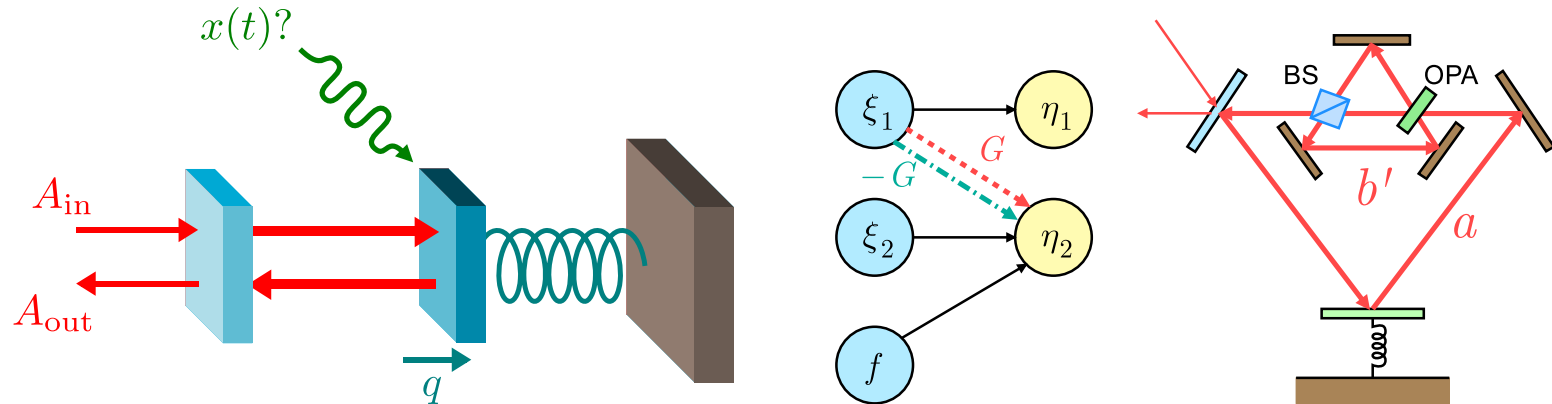
$$\text{e.g. } S_{\Delta \hat{I}}^{\text{coh}}(\omega) = \frac{\bar{P}}{\hbar\omega_0}, \quad S_x^{\text{OU}}(\omega) = \frac{\kappa}{\omega^2 + \epsilon^2}. \quad (1)$$

Achieved by **homodyne phase-locked loop** + **Smoothing** [Personick IEEE TIT **17**, 240 (1971); Tsang, Shapiro, and Lloyd, PRA **78**, 053820 (2008); **79**, 053843 (2009)]

Wheatley *et al.*, PRL **104**, 093601 (2010).

Interferometry, ranging, velocimetry, clock synchronization, coherent comm., etc.

Optomechanical Force Sensing



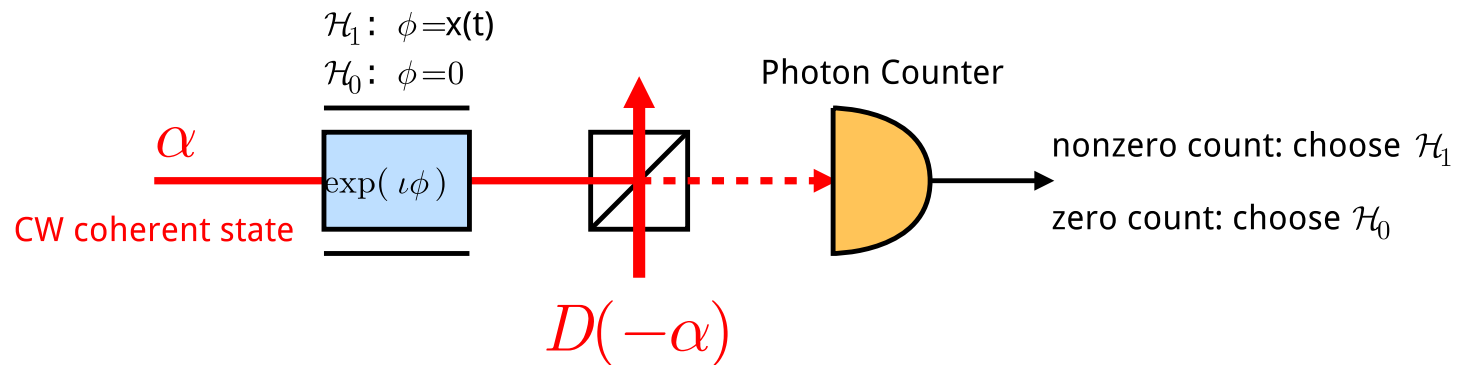
- QCRB [Tsang, Wiseman, and Caves, PRL **106**, 090401 (2011)]:

$$\langle \delta x^2 \rangle \geq \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{1}{(4/\hbar^2) S_{\Delta \hat{q}}(\omega) + 1/S_x(\omega)}. \quad (2)$$

- Achieved by [Quantum Backaction Noise Cancellation](#) [Kimble *et al.*, PRD **65**, 022002 (2001); Chen, PRD **67**, 122004 (2003); Tsang and Caves, PRL **105**, 123601 (2010)] + [Smoothing](#) [Tsang, PRL **102**, 250403 (2009)].

Optical Phase Waveform Detection

- Binary hypothesis testing $\mathcal{H}_0 : \phi(t) = 0$, $\mathcal{H}_1 : \phi(t) = x(t)$
- Tsang, unpublished:

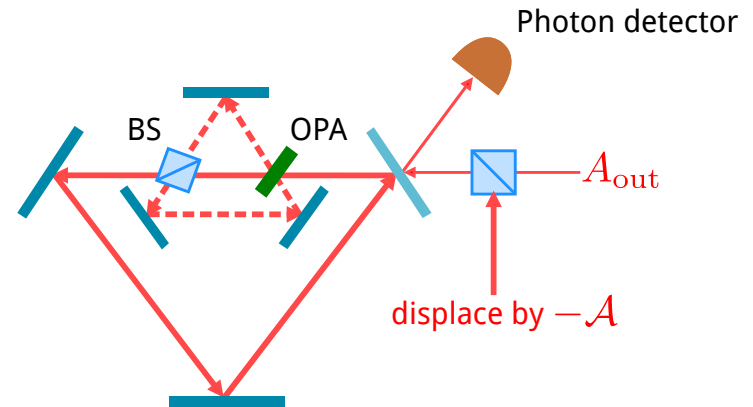
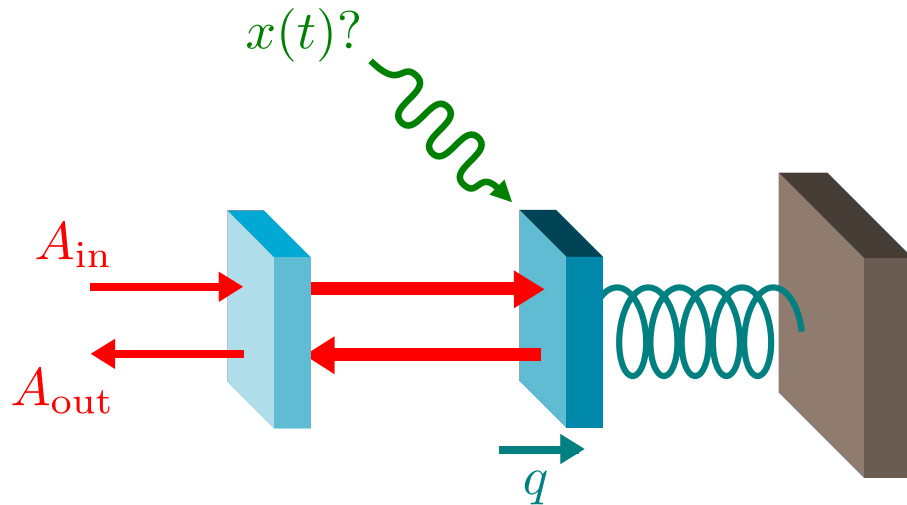


$$P_e \geq \frac{1}{2} \left(1 - \sqrt{1 - 4P_0P_1F} \right), \quad (3)$$

$$F = \int Dx P[x] \left| \langle \psi | \exp \left[i \int_{t_0}^T dt \hat{I}(t)x(t) \right] | \psi \rangle \right|^2. \quad (4)$$

- Kennedy receiver** has optimal error exponent for stochastic waveform detection with coherent state.
- Homodyne performance depends on **prior waveform statistics**

Optomechanical Force Detection



- Tsang and Nair, arXiv:1204.3697 (2012):

$$P_e \geq \frac{1}{2} \left(1 - \sqrt{1 - 4P_0 P_1 F} \right), \quad F = \int Dx P[x] \left| \langle \psi | \mathcal{T} \exp \left[i \int_{t_0}^T dt \hat{q}_0(t) x(t) \right] | \psi \rangle \right|^2 \quad (5)$$

- Because the bound is **achievable** for deterministic $x(t)$ and not limited by backaction noise, SQL can definitely be overcome
- QNC + Kennedy receiver achieves optimal error exponent

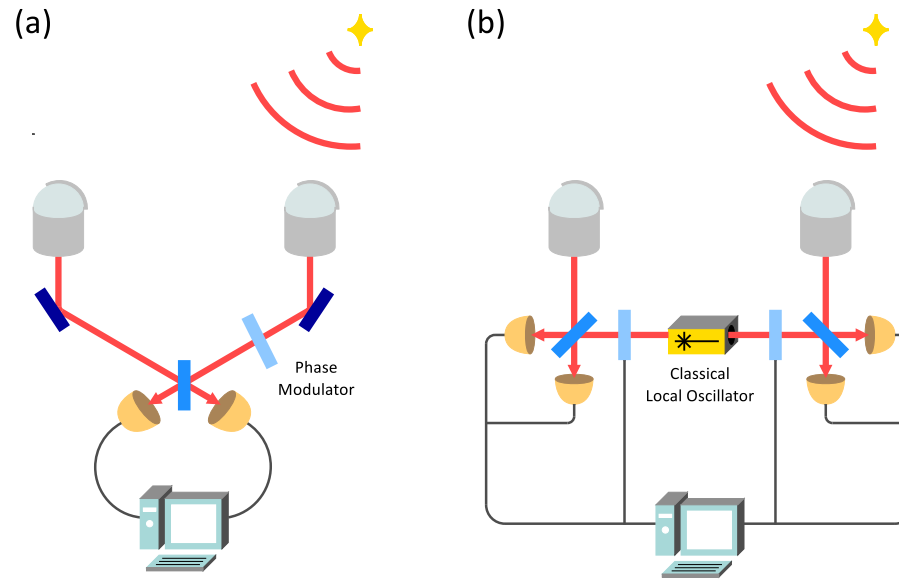
Decoherence

- Significant decoherence/loss rules out any significant quantum enhancement using **squeezed state/nonclassical state of light** (Durkin *et al.*/Escher *et al.*)
- **No-go** for nonclassical light in space optics applications
- No result yet about decoherence in **waveform estimation/detection**, surprise unlikely
- Quantum illumination (Lloyd/Erkmen/Guha/Shapiro/Giovannetti *et al.*): up to **6 dB improvement in error exponent!**
 - Producing squeezed state requires strong pump with way more photons
 - can be achieved by coherent state with 6 dB more photons
 - Known receivers can't get to 6 dB
 - Low-photon-number regime only
 - Gaussian noise strengths are assumed to be different under hypotheses for QI to be useful, passive target detection may be better in practice
- **Quantum Metrology with POVMs**

Quantum Imaging

- Ghost imaging: Shih/Shapiro/Erkmen
- Sub-Rayleigh quantum lithography/imaging: Boto *et al.*, PRL **85**, 2733 (2000); Tsang, PRL **102**, 253601 (2009); Giovannetti *et al.*, PRA **79**, 013827 (2009)
- Experiments: D'Angelo *et al.*, PRL **87**, 013602 (2001); Guerrieri *et al.*, PRL **105**, 163602 (2010); Shin *et al.*, PRL **107**, 083603 (2011)
- Classical computational sub-Rayleigh imaging: STORM/PALM [Zhuang, Nature Photon. **3**, 365 (2009)]; STED (Hell), etc.
- Computational imaging for astronomy [Fienup]
- Not much rigorous work in quantum imaging that uses estimation/detection theory

Quantum Camera Design

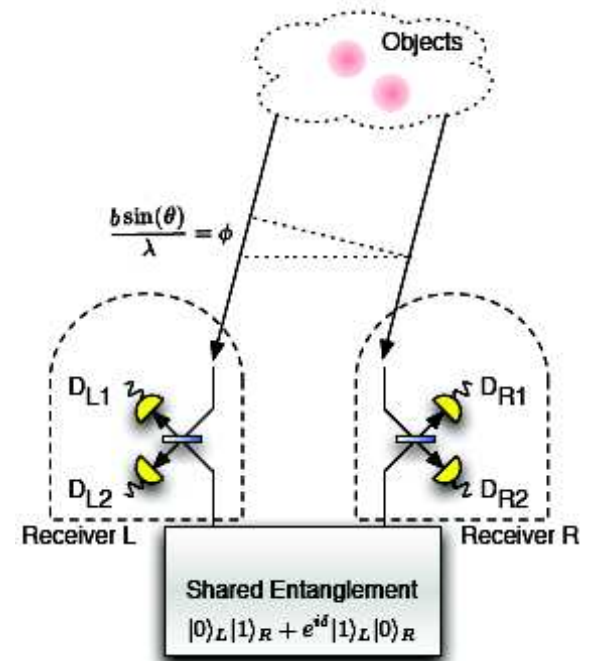
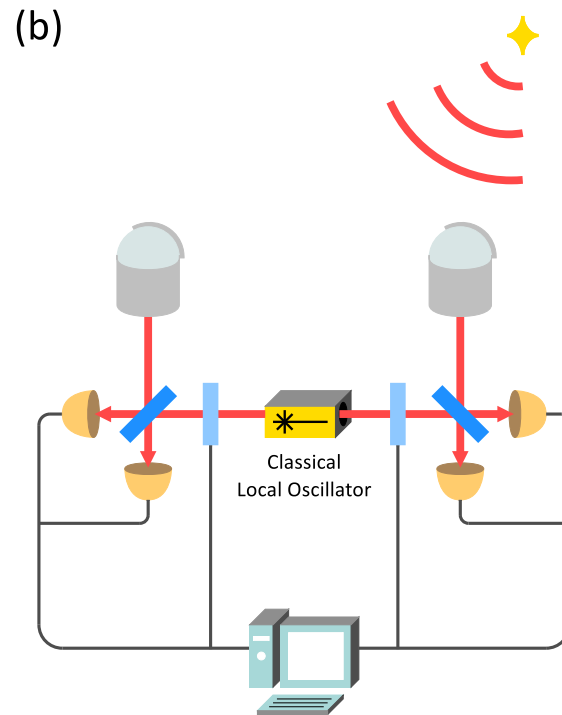
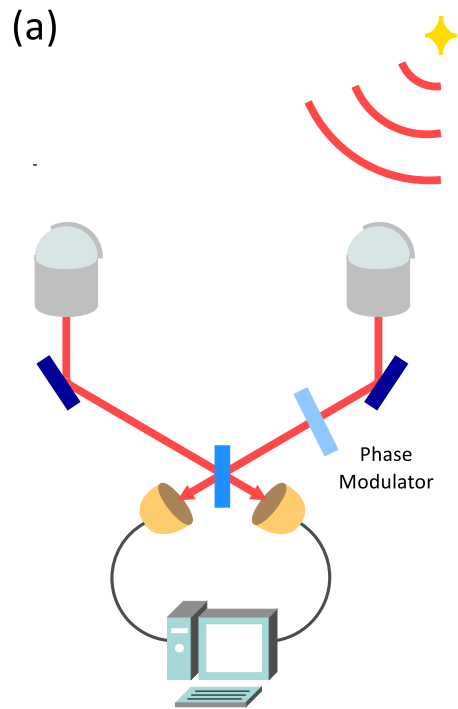


- start with multi-spatial-mode $\rho_x(\mathbf{r})$ (e.g., multimode thermal or coherent state)
- Record with imaging system/CCD/interferometer/digital holography (model by POVM $E[y(\mathbf{r})]$)

$$P[y(\mathbf{r})|x(\mathbf{r}')] = \text{tr} \{ E[y(\mathbf{r})] \Phi_{\text{aperture}} \Phi_{\text{diffract}} \rho_x(\mathbf{r}') \} \quad (6)$$

- Quantum bounds: multiparamter QCRB, etc.
- Do conventional imaging systems saturate these bounds?
- How to implement optimal POVM?

Stellar Interferometry



Gottesman, Jennewein, and Croke, e-print arXiv:1107.2939.

● Estimation of coherence:

$$\Gamma_{ab} = \langle b^\dagger a \rangle,$$

$$g^{(1)} = \frac{\langle b^\dagger a \rangle}{\sqrt{\langle b^\dagger b \rangle \langle a^\dagger a \rangle}} \quad (\text{normalized}). \quad (7)$$

Old-School Quantum Optics

- P representation:

$$\rho = \int d^2\alpha d^2\beta \Phi(\alpha, \beta) |\alpha, \beta\rangle \langle \alpha, \beta|. \quad (8)$$

- $\Phi(\alpha, \beta)$ is a two-mode zero-mean Gaussian for thermal light, i.e. **no entanglement**
- weak thermal light $\epsilon \equiv \langle a^\dagger a \rangle = \langle b^\dagger b \rangle \ll 1$ in photon-number basis:

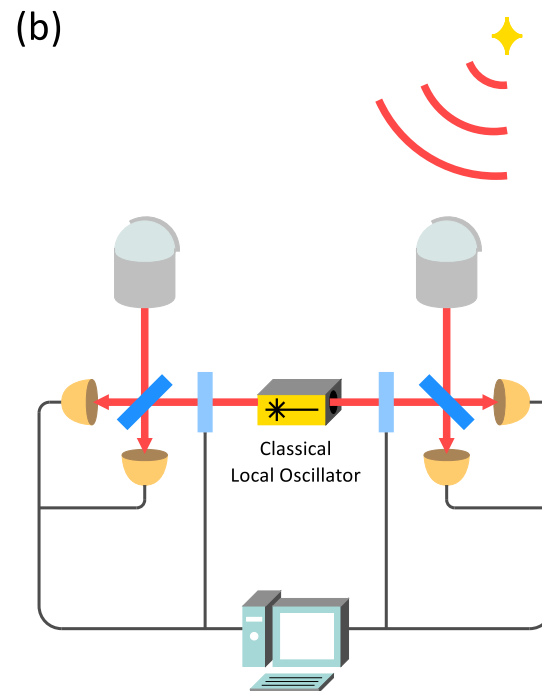
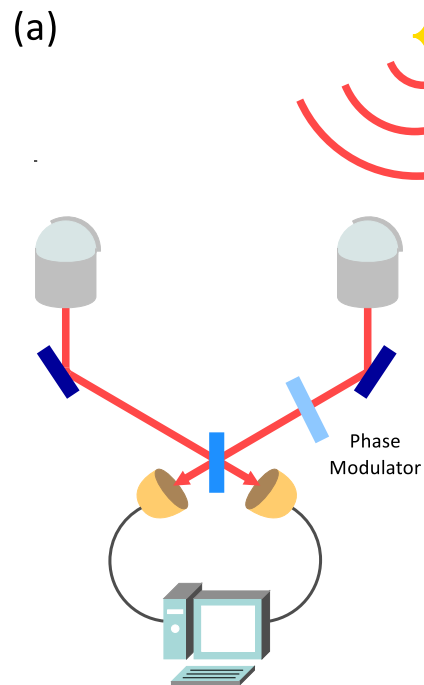
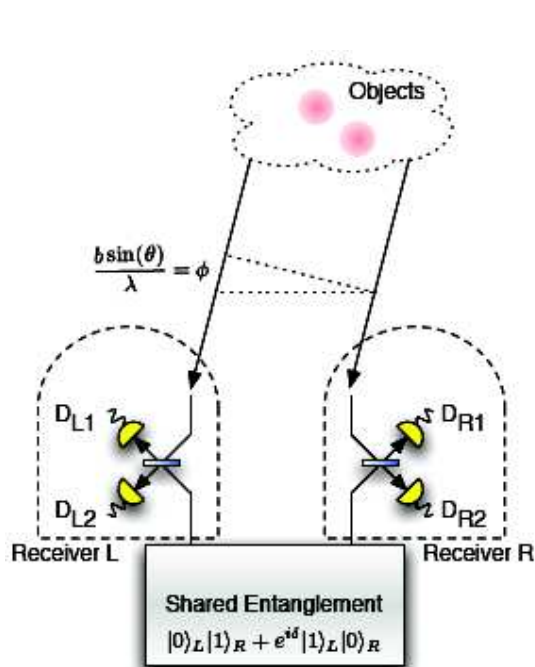
$$\begin{aligned} \rho = & (1 - \epsilon) |0, 0\rangle \langle 0, 0| + \frac{\epsilon}{2} [|0, 1\rangle \langle 1, 0| + |1, 0\rangle \langle 1, 0| + g^* |0, 1\rangle \langle 1, 0| + g |1, 0\rangle \langle 0, 1|] \\ & + O(\epsilon^2), \end{aligned} \quad (9)$$

$$P(y|g) = \text{tr} [E(y)\rho]. \quad (10)$$

- Classical Fisher information for $g = g_1 + ig_2$:

$$F_{jk} = \left\langle \frac{\partial}{\partial g_j} \ln P \frac{\partial}{\partial g_k} \ln P \right\rangle, \quad \Sigma \geq \frac{1}{M} F^{-1} \quad (11)$$

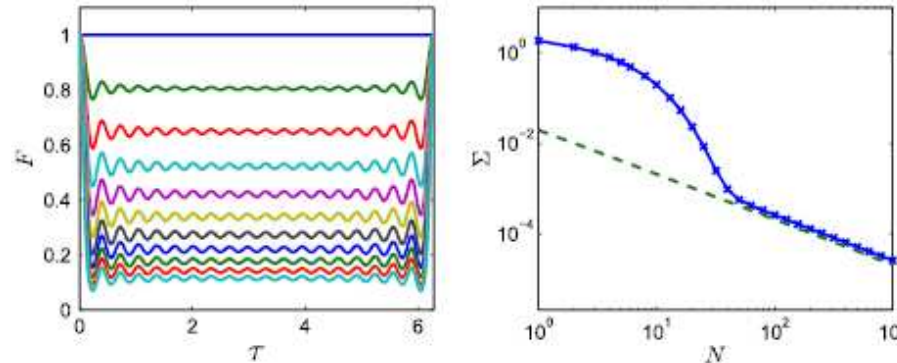
Bound for Local Measurements



- Nonlocal measurements (direct detection, shared-entanglement): $\|F\| \sim \epsilon$.
- A necessary condition for local (LOCC) measurement is **the PPT condition applied to the POVM** [Terhal *et al.*, PRL 86, 5807 (2001)]. Then $\|F\| \leq \epsilon^2 + O(\epsilon^3)$.
- Generalizable to repeated LOCC measurements
- **Quantum nonlocality** in measurement of nature, even if the state has no entanglement.
- M. Tsang, PRL **107**, 270402 (2011).

Misc.

- Quantum Ziv-Zakai bounds [Tsang, PRL **108**, 230401 (2012)]



- Continuous quantum hypothesis testing (for tests of physics using continuous quantum measurements) [Tsang, PRL **108**, 170502 (2012)]
- Cavity quantum microwave photonics [Tsang, PRA **81**, 063837 (2010); **84**, 043845 (2011)]

