
Recent & Historical Geochronology Instrument Developments @ JPL

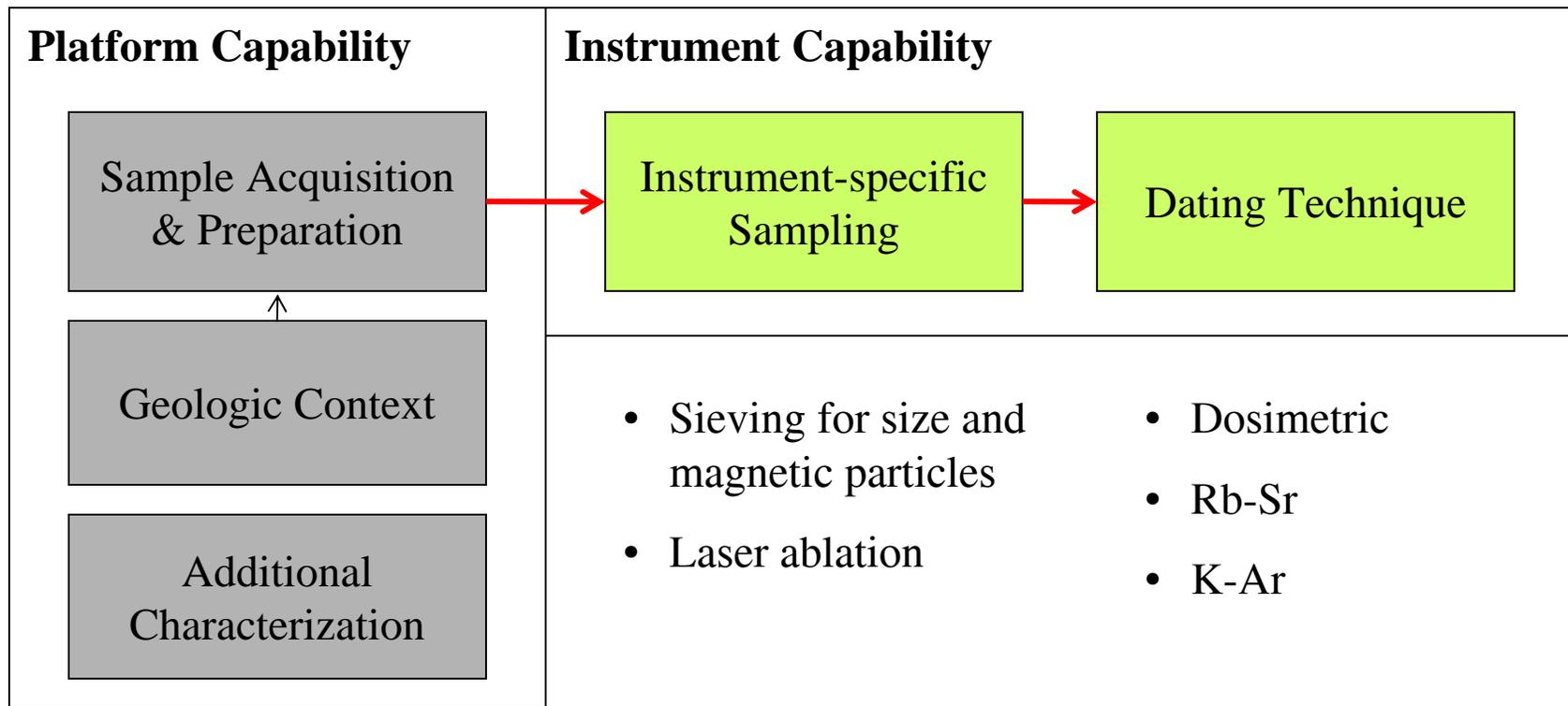
Paula Grunthaner

Special thanks to J. P. Kirby

Jet Propulsion Laboratory, California Institute of Technology

Topics

- What are the challenges for an *in situ* system?
- What approaches have been (or are being) tried?
 - Intent is to provide a sense of the development status of geochronology instrumentation
- What challenges do these developments face going forward to flight?



- Chronological techniques not applicable to all rock types
- Sample must be free of post-crystallization disturbances
- Dust contamination
- As a minimum, expose fresh surface
- Desirable, pre-selection of grains

In Situ Geochronometer Developments

- **Luminescence/ESR technique**
 - JPL, Ulmer Systems, Oklahoma State U, McMaster U, LPI
- **^{87}Rb – ^{87}Sr decay**
 - JPL, U Pittsburgh → Scott Anderson's talk
 - JPL, U. Wisconsin, Tangent Technologies, OI Analytical
- **^{40}K – ^{40}Ar decay & cosmogenic nucleotide buildup (^3He , $^{20,21,22}\text{Ne}$, $^{36,38}\text{Ar}$)**
 - UA, JPL, LANL → Tim Swindle's talk

Luminescence and Electron-spin Resonance Dating

S. Kim (JPL), S. McKeever (Oklahoma State U), W. Rink (McMaster U),
S. Clifford, LPI, A. Yen (JPL)

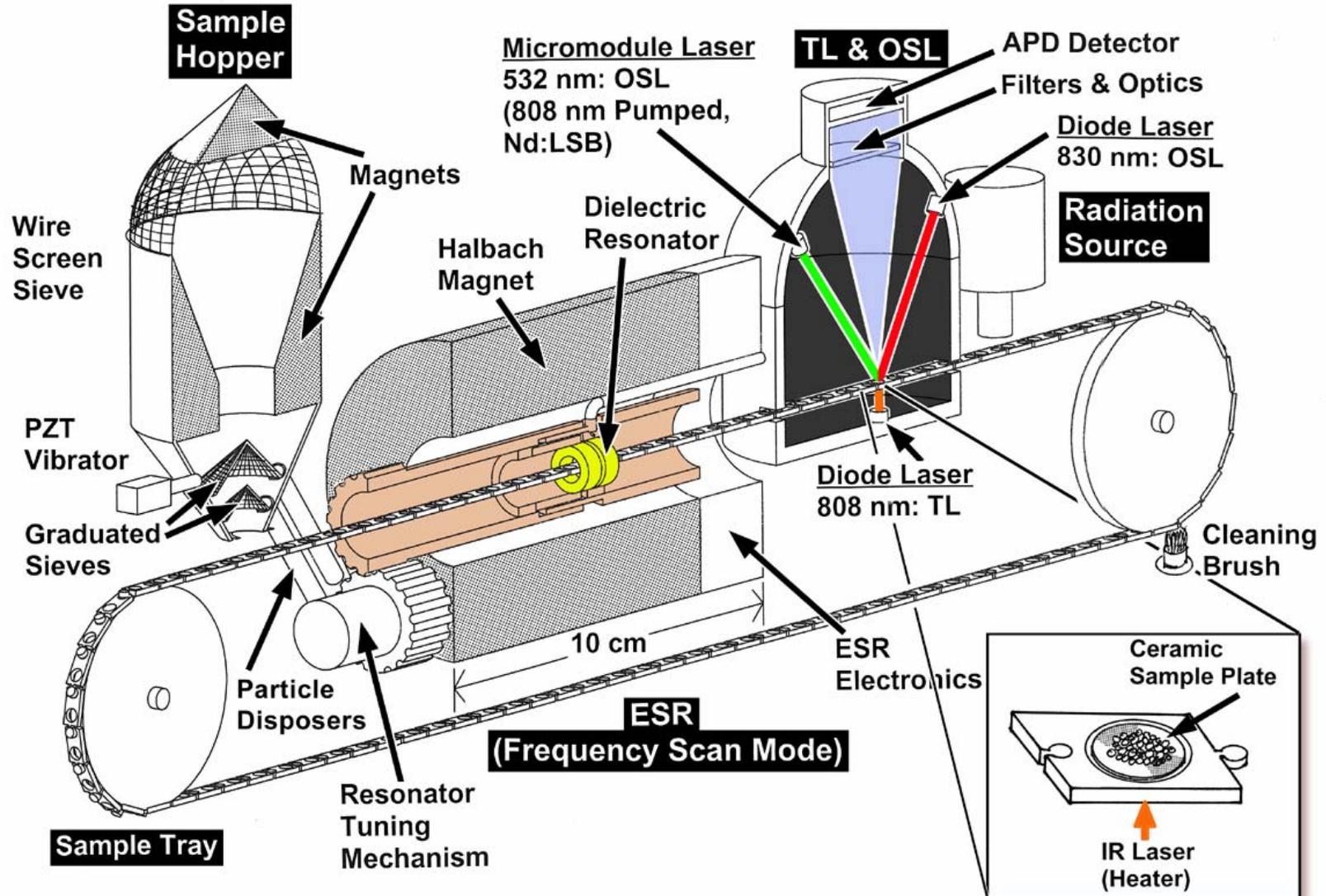
- Dosimetric approach that integrates TL, OSL, and ESR
- All 3 used for dating time since some dynamic geological resetting event
 - Time since last exposure to sunlight (solar reset) and thus burial ages of soils in sediments
 - Time since last thermal event such as volcanic heating, mechanical stresses
 - Measures the accumulation of absorbed radiation dose within mineral grains as $f(\text{time})$
 - Ionizing radiation accumulates trapped charge in the mineral, which is depleted when exposed to solar radiation (resetting the clock)
 - Luminescence intensity \propto dose since last exposure to sunlight \propto depositional age
- Utility: dating eolian, fluvial, periglacial, impact, volcanic processes spanning the past 100 ka – 1 Ma)

Applicability of techniques, as practiced on Earth

	TL	OSL	EPR
Eolian/Fluvial deposits (age of burial, solar reset)	M	H	L
Volcanic Rocks (crystallization age)	H	H	L
Sedimentary Volatile Deposits (age of crystallization of carbonates, sulfates)	L	L	H
Dust Particles in Polar Ice Caps (age of incorporation into ice)	L	H	L
Material characterization	L	L	H

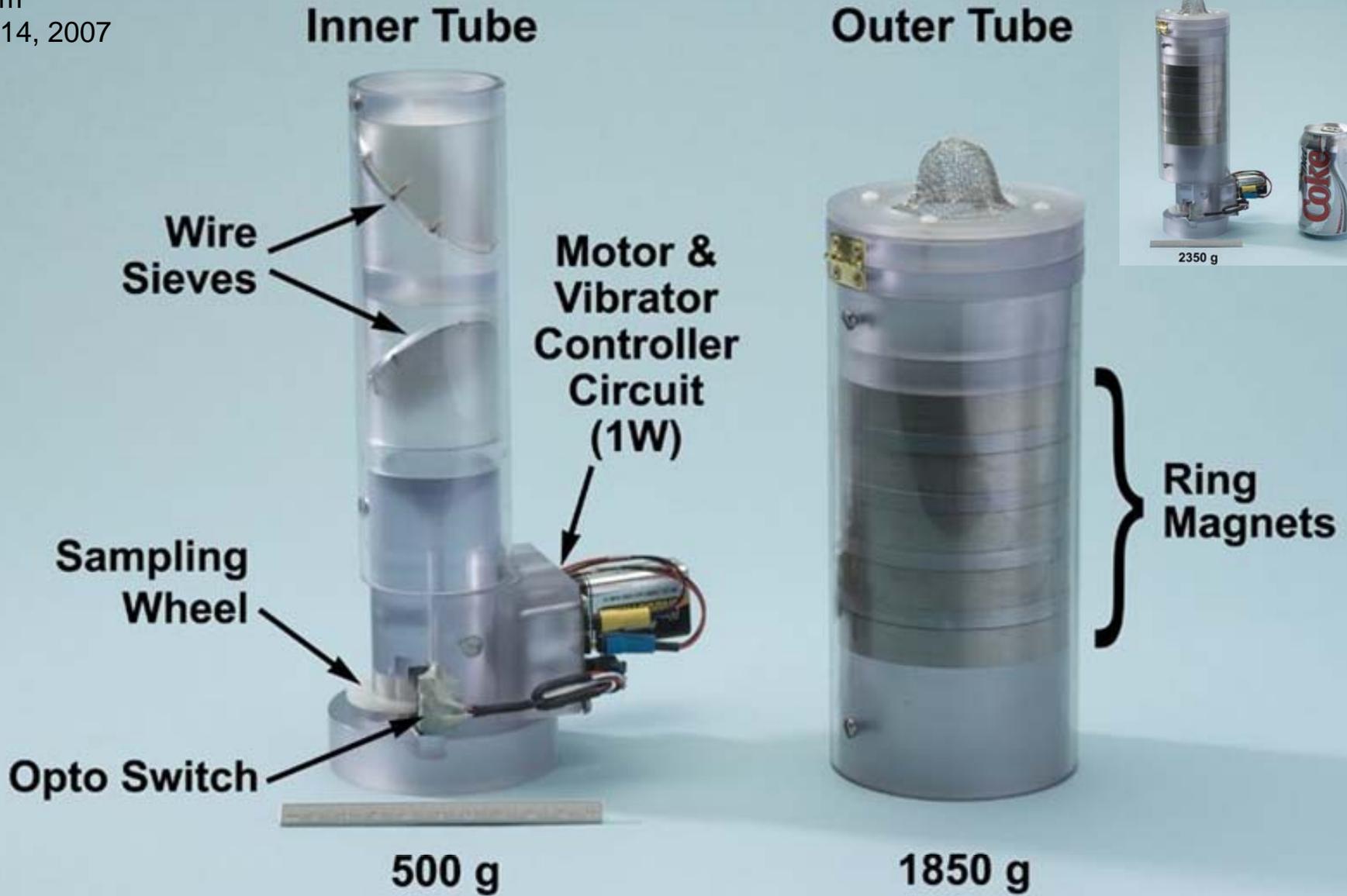
Applicability H: High M: Medium L: low

Luminescence and Electron-spin Resonance Dating Breadboard Concept



Luminescence and Electron-spin Resonance Dating Development Status—Sample Preparation Unit

S. Kim
May 14, 2007



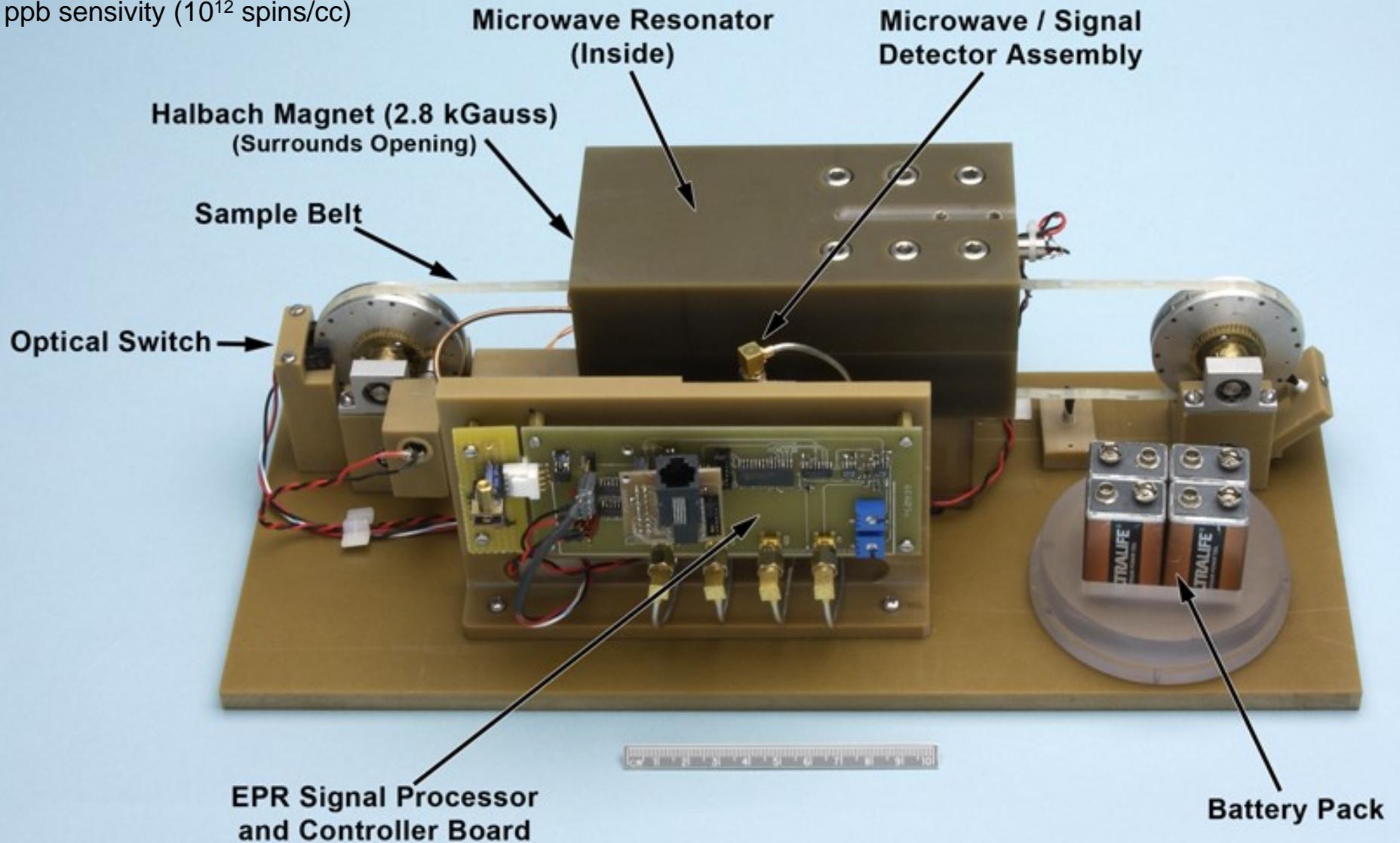
Luminescence and Electron-spin Resonance Dating Development Status—ESR Spectrometer subsystem

S. Kim

May 14, 2007

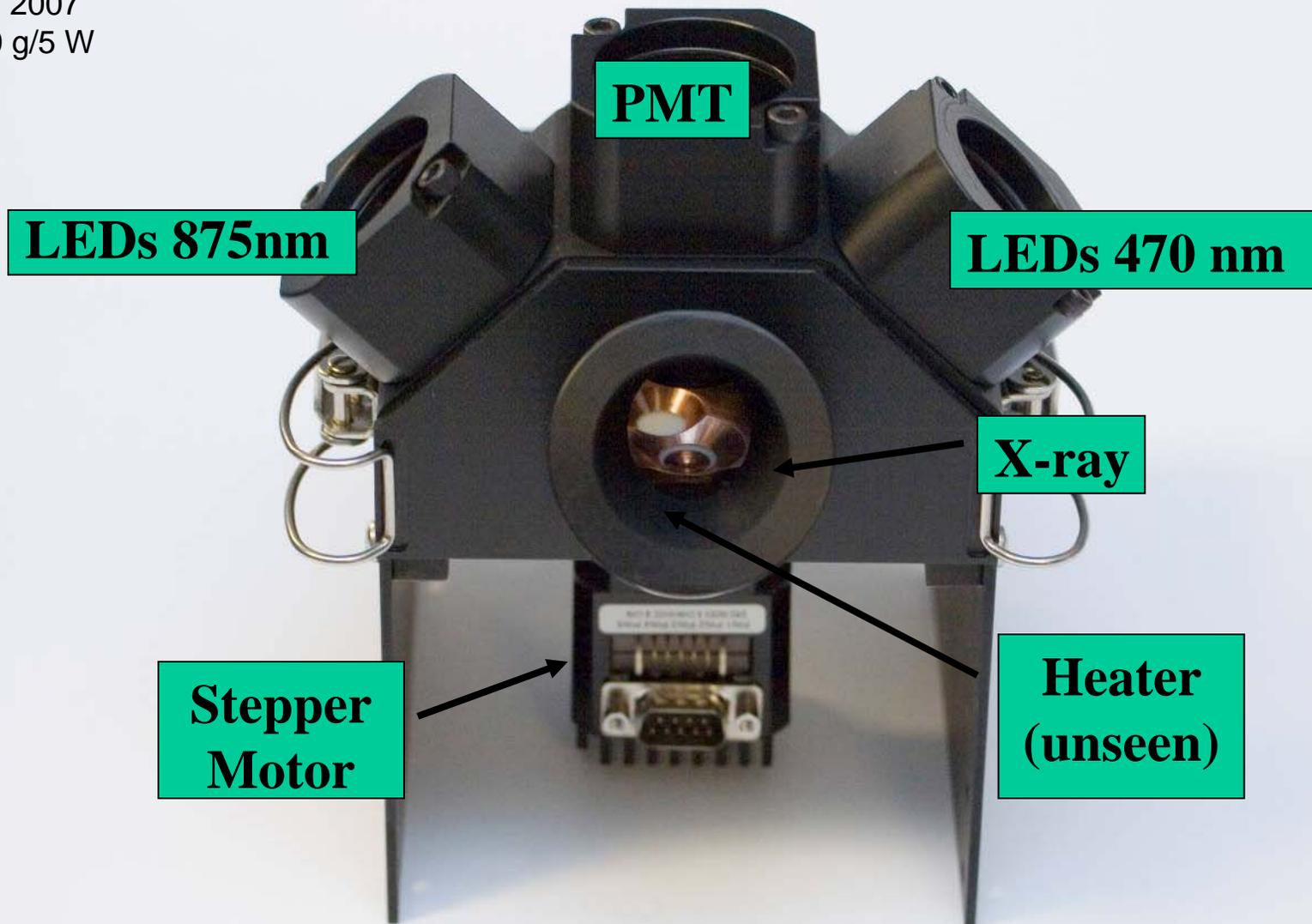
ppb sensitivity (10^{12} spins/cc)

(3W, 3.8 kg)



Luminescence and Electron-spin Resonance Dating Development Status— OSL and TL subsystem

S. Kim / Oklahoma State
May 14, 2007
ppb/400 g/5 W



Luminescence and Electron-spin Resonance Dating Challenges Going Forward

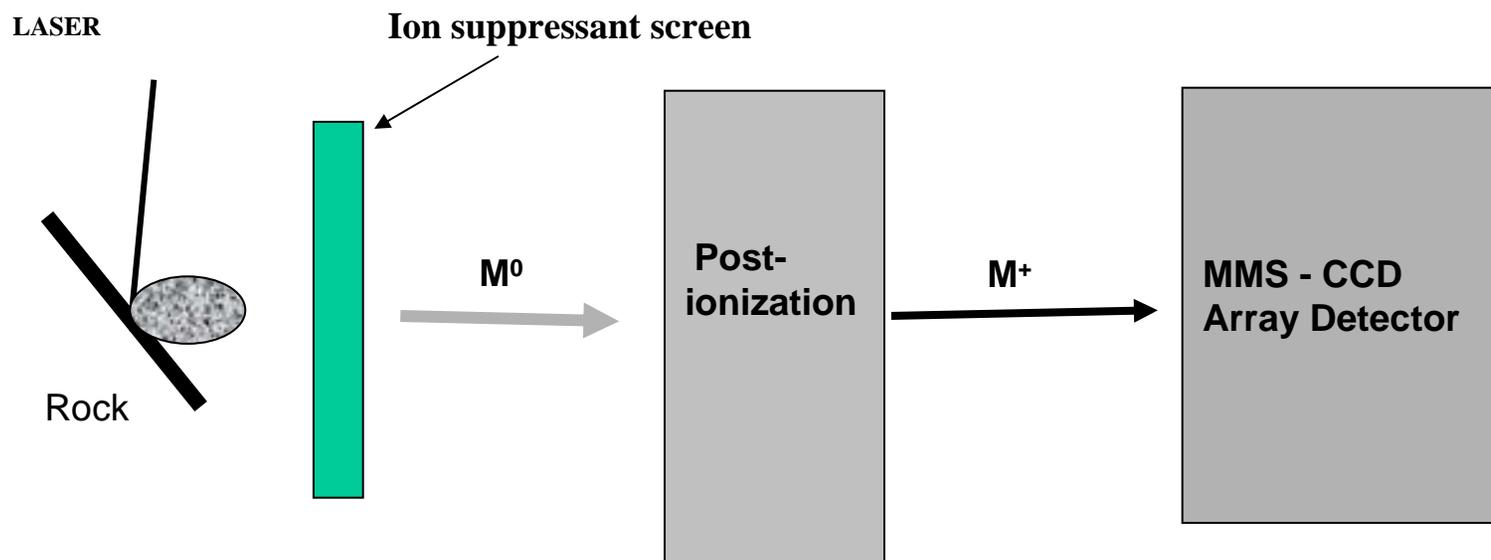
- No currently funded task
 - JPL PI submitted PIDDP on a different topic because geochronology not a priority in the NRA
- Science— Do we understand dosimetry on Mars well enough?
 - On Earth, use mineral separates; on Mars likely polymineralic
 - On Mars, lower natural dose rate but 1000x higher cosmic ray flux
 - Necessitates need to know burial rate
 - Gradual accumulation over time => variable dose rate
 - Efficiency of solar bleaching in low T of Mars

Rb-Sr Geochronology

- A few reminders about the technique
 - ^{87}Rb to ^{87}Sr by β^- decay
 - $(^{87}\text{Sr}/^{86}\text{Sr})_i = 0$ not valid nor is it possible to estimate. So ratio is calculated by measuring 2 or more minerals in same rock with spread in Rb/Sr ratios
 - Measurement masses are 85, 86, 87. Use fact that $^{87}\text{Rb}/^{85}\text{Rb}$ is constant in all SS materials to derive desired D_r/D_s and P/D_s for isochrons

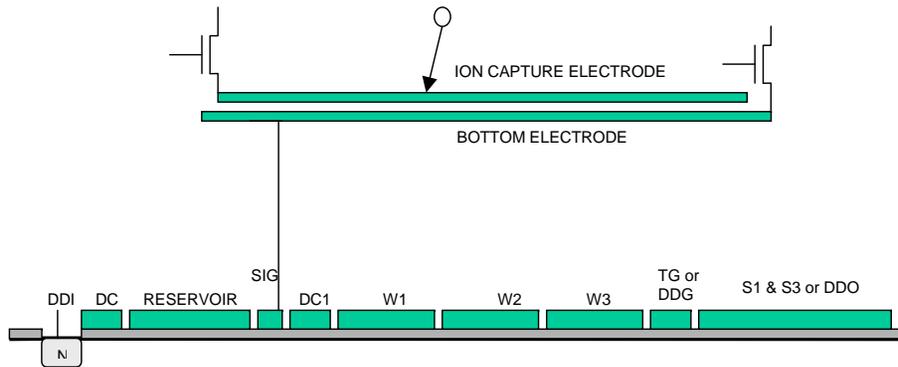
Laser Ablated – Electron Impact Ionization – MS for Geochronology (Rb-Sr)

M. Sinha (JPL), B. Beard (U WI), M. Wadsworth (Tangent), OI Analytica

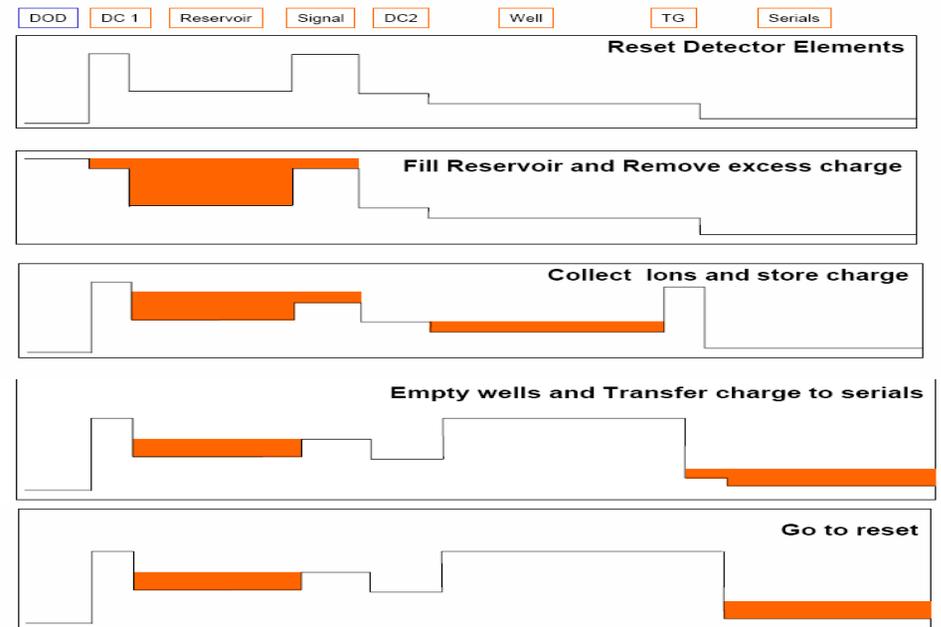


- Laser-ablation and sampling of neutrals only
 - Potential for cleaner mineral separates & better understanding of internal isochrons
 - 10^2 – 10^6 more neutrals than ions; may better represent elemental abundance
- Electron impact ionization of neutrals creates 100–1000 μ s ion pulse
- Non-scanning magnetic sector MS with custom focal plane array detector integrates entire ion pulse simultaneously
- Estimated precision of age dating: 95% confidence of ± 150 Ma for a 500 Ma rock

Direct Ion Detection with a Modified CCD Focal Plane Array



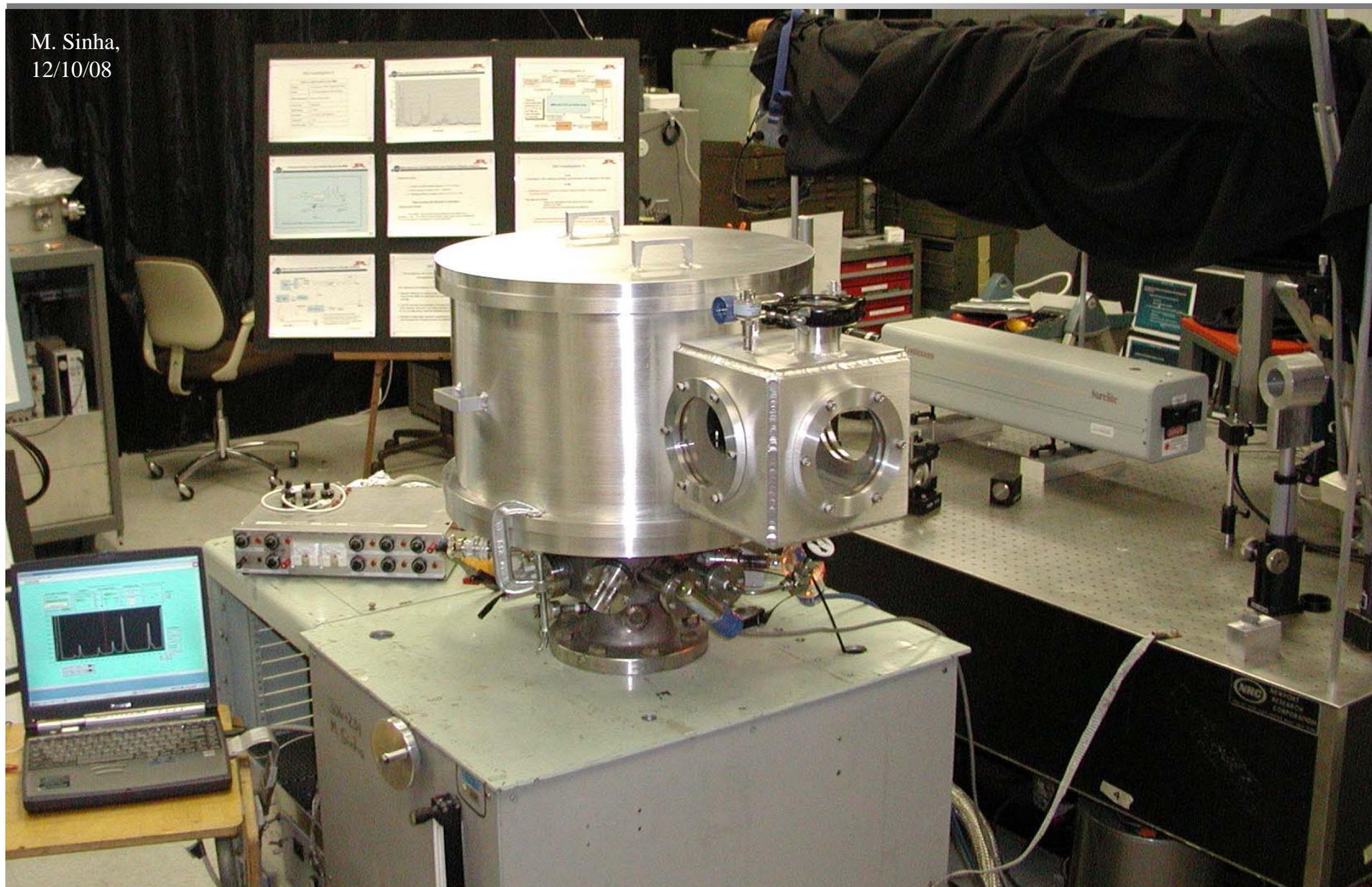
- Linear array of CCD pixels
- Photodiode replaced by a capacitive sensing element that serves as ion detector
- Capacitive element is coupled to CCD shift register and creates a packet of signal charge proportional to the charge on the capacitor
- 2nd generation CCD improves sensitivity, reduces noise, increases dynamic range and resolution



Laser Ablated – Electron Impact Ionization – MS for Geochronology

Sinha, et al

M. Sinha,
12/10/08

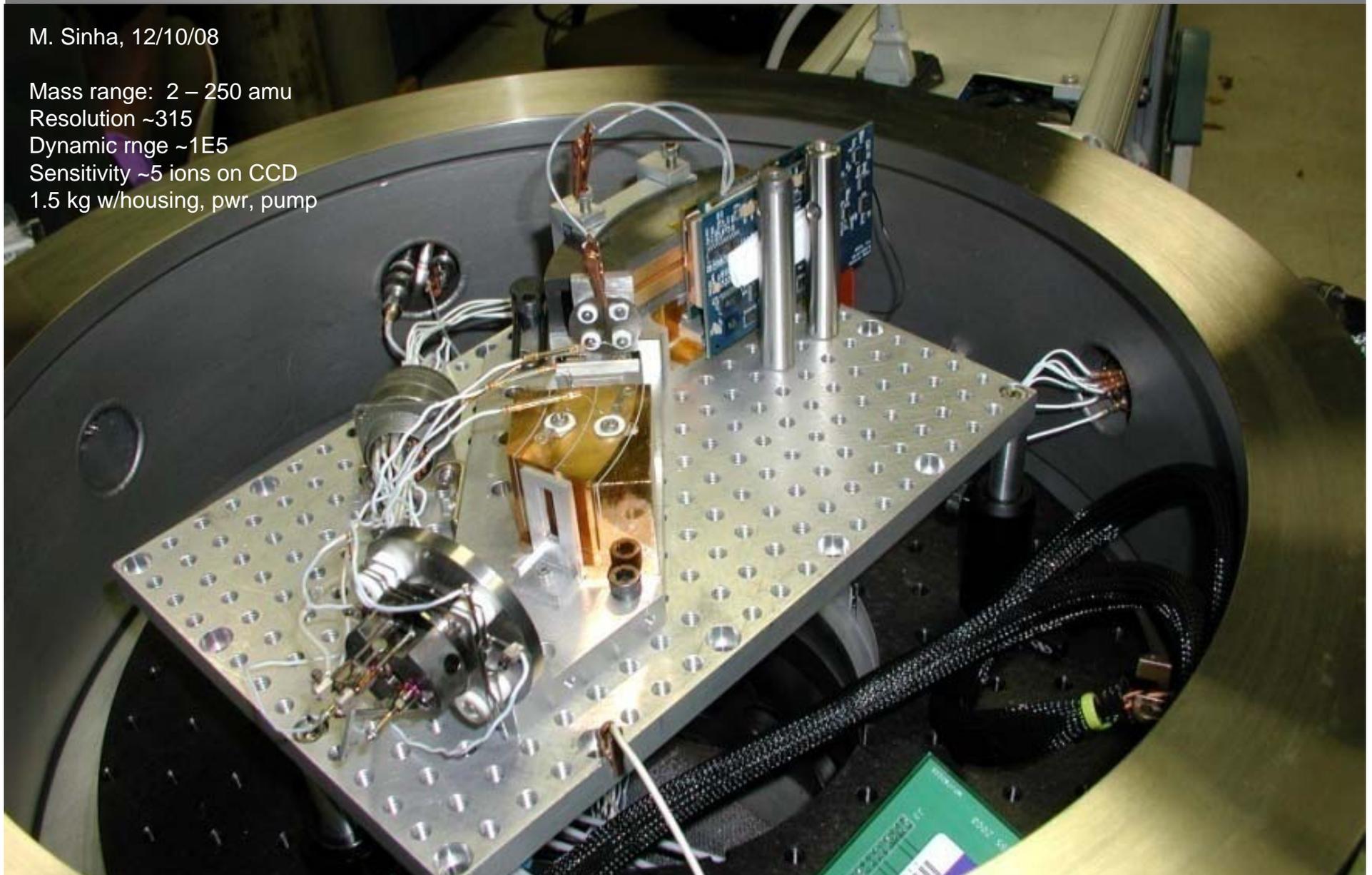


Laser Ablated – Electron Impact Ionization – MS for Geochronology

Sinha, et al

M. Sinha, 12/10/08

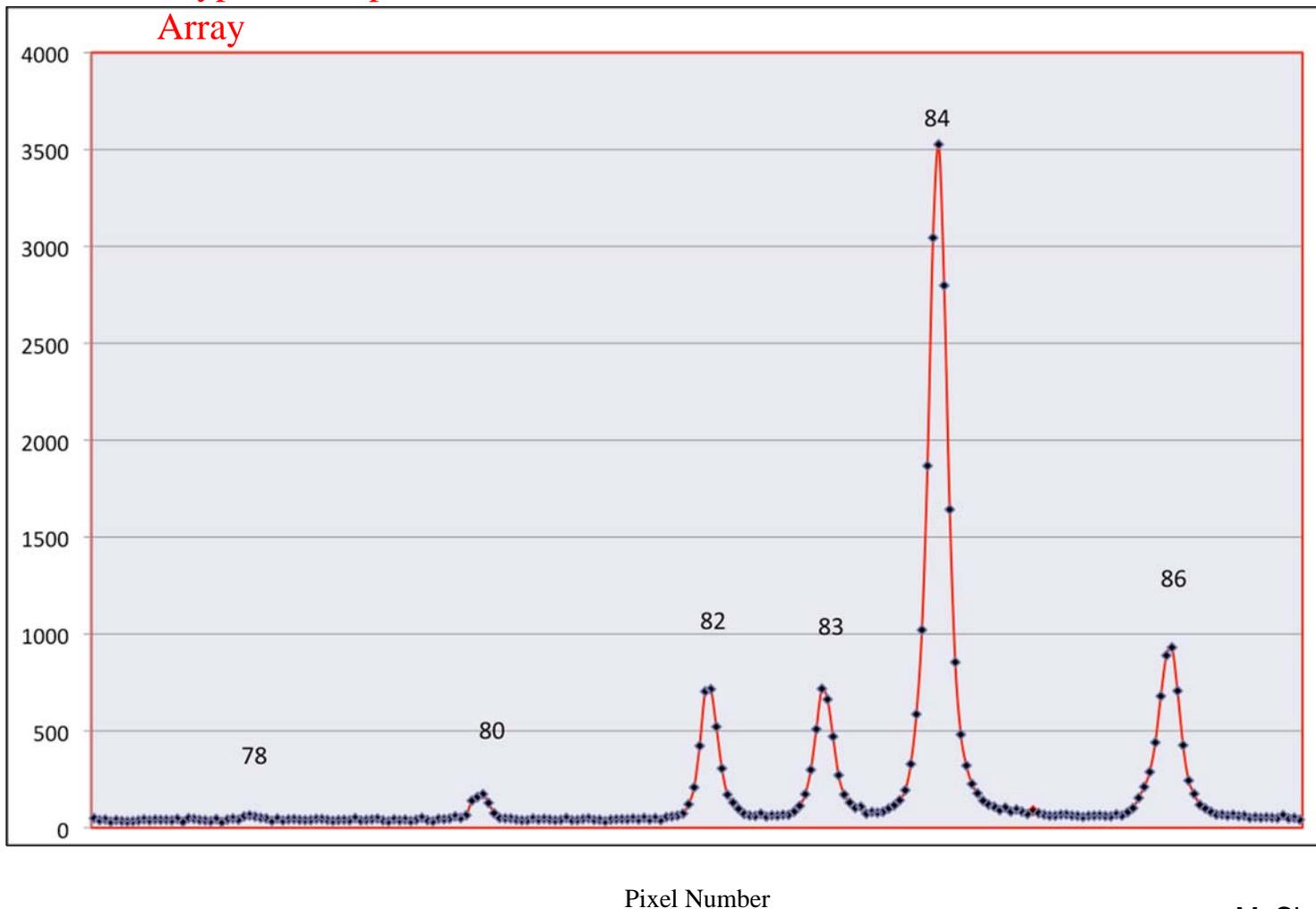
Mass range: 2 – 250 amu
Resolution ~315
Dynamic range ~1E5
Sensitivity ~5 ions on CCD
1.5 kg w/housing, pwr, pump



Laser Ablated – Electron Impact Ionization – MS for Geochronology

Sinha, et al

Krypton Isotope Measurement with CCD Direct-Ion-Detector Array



Error Estimates for Rb-Sr ratios

	Plagioclase	Clinopyroxene	Orthopyroxene	Olivine	
Sr ppm	100	10	3	1.5	
$^{87}\text{Sr}/^{86}\text{Sr}$ and % error	$0.70537 \pm 0.0793 \%$	$0.70990 \pm 0.1770 \%$	$0.71134 \pm 0.5499 \%$	$0.70836 \pm 0.5869 \%$	yesterday $\pm 1\%$
$^{87}\text{Rb}/^{86}\text{Sr}$ and % error	$0.0513 \pm 1.277 \%$	$0.687 \pm 0.349 \%$	$0.8901 \pm 0.5869 \%$	$0.4717 \pm 1.203 \%$	$\pm 10\%$

- Assumes 1×10^6 atoms/laser shot, 10% of atoms delivered to ion source, 1 in 1000 atoms ionized, 20% of ions arrive at detector, ions accumulated for 100 shots, all minerals had $(^{87}\text{Sr}/^{86}\text{Sr})^i$ of 0.705 and are 500 Ma old
- For a 4-mineral isochron using the above estimates, 95% confidence of age error ± 140 Ma for 500 Ma rock

Challenges Going Forward for Both Rb-Sr Systems

- Rb and Sr must be ionized in direct proportion to their abundance in the sample
- Laser ablation for sampling
- Potential for cleaner mineral separates & better understanding of internal isochrons
- Both techniques sample neutrals, not ions; 10^2 – 10^6 more neutrals than ions; may better represent elemental abundance
- But, laser ablation (w/ICPMS) has shown significant elemental fractionation for elements of different volatility. Neutral generation is thermal.
 - Rb (688 °C) and Sr (1382 °C) have very different melting points
 - The extent of elemental fractionation must be characterized for different minerals
- Ionization efficiencies must be understood. In this respect, electron impact is more commonly used and may be better understood than resonant ionization (but, heck, I'm not sure)
- Resonant ionization is highly sensitive for ionization and highly selective. Electron impact is not selective and can, therefore, be used to characterize other neutral species.
- Both approaches must develop differentially pumped systems to keep sample outside (or bring the sample inside and pump down each time)

Summary

- There are several age dating systems operating at TRL 4 with potential to meet performance requirements
- None of these systems have yet demonstrated they can deliver the performance demanded
 - Elemental sampling and/or ionization bias effects need to be examined
 - H/W needs miniaturization & subsystems
- KISS should identify the tall tent poles that may not be funded under existing PIDDP/MIDP/etc funds
- To paraphrase Gregg, development needs to proceed *tout de suite*