

# Exoplanets, SETI, and Technosignatures

A short course for the KISS workshop

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(for Jason Wright)

May 20, 2019

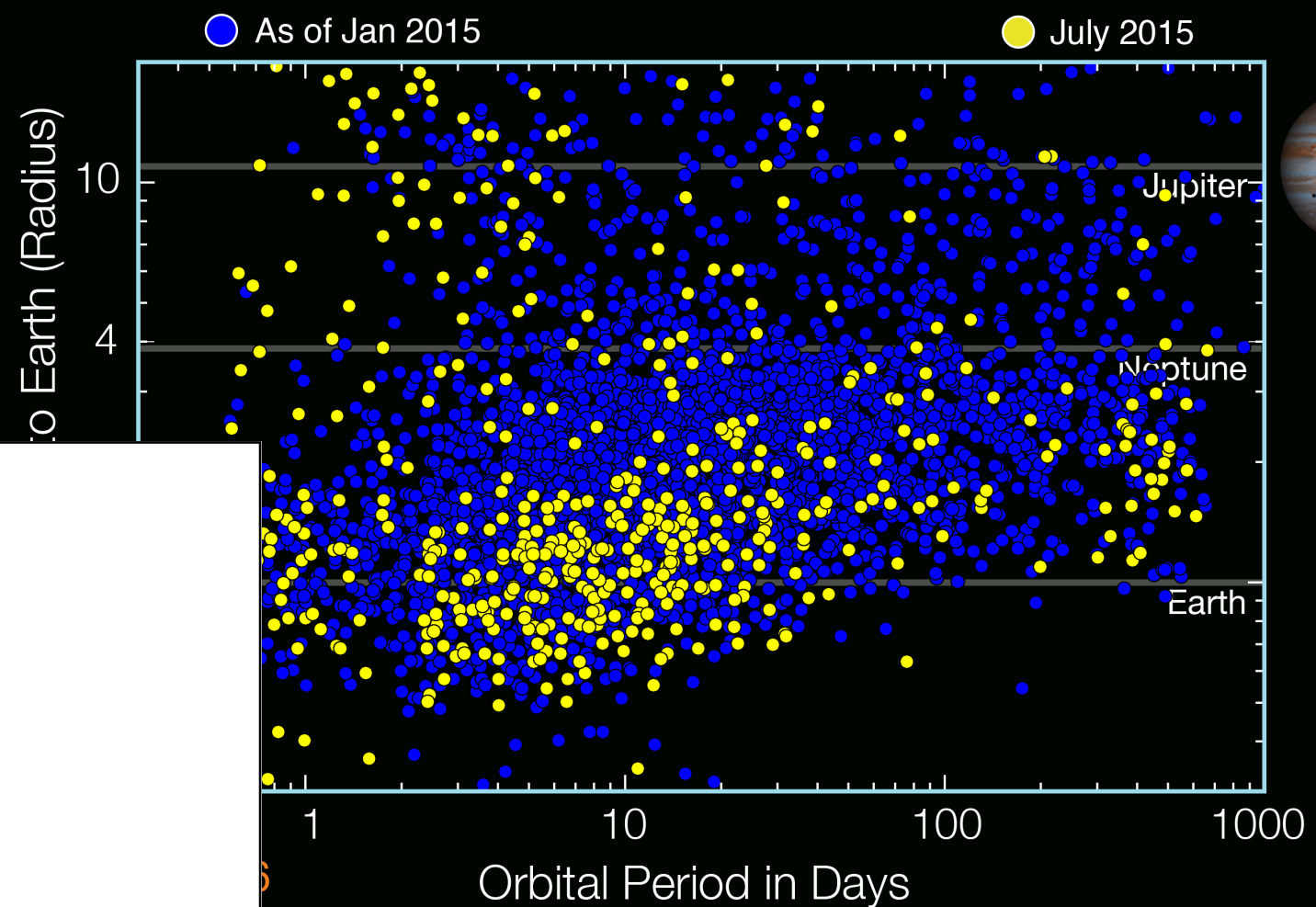
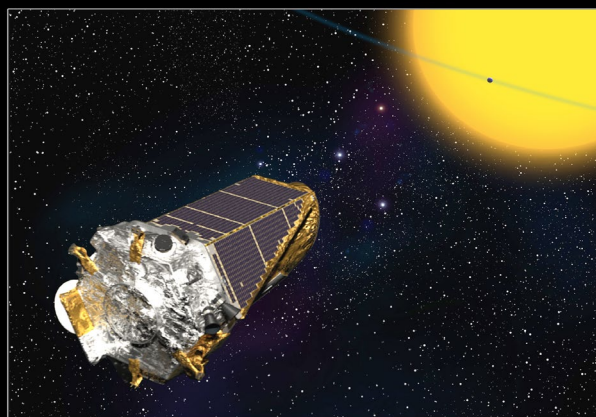
# Exoplanets are Ubiquitous

Number of planets per star:  $\sim 1$  (bigger than Earth, inside 1 AU)

Fraction of Sun-like stars with Earth-size planet in Habitable Zone: 20% (\*)

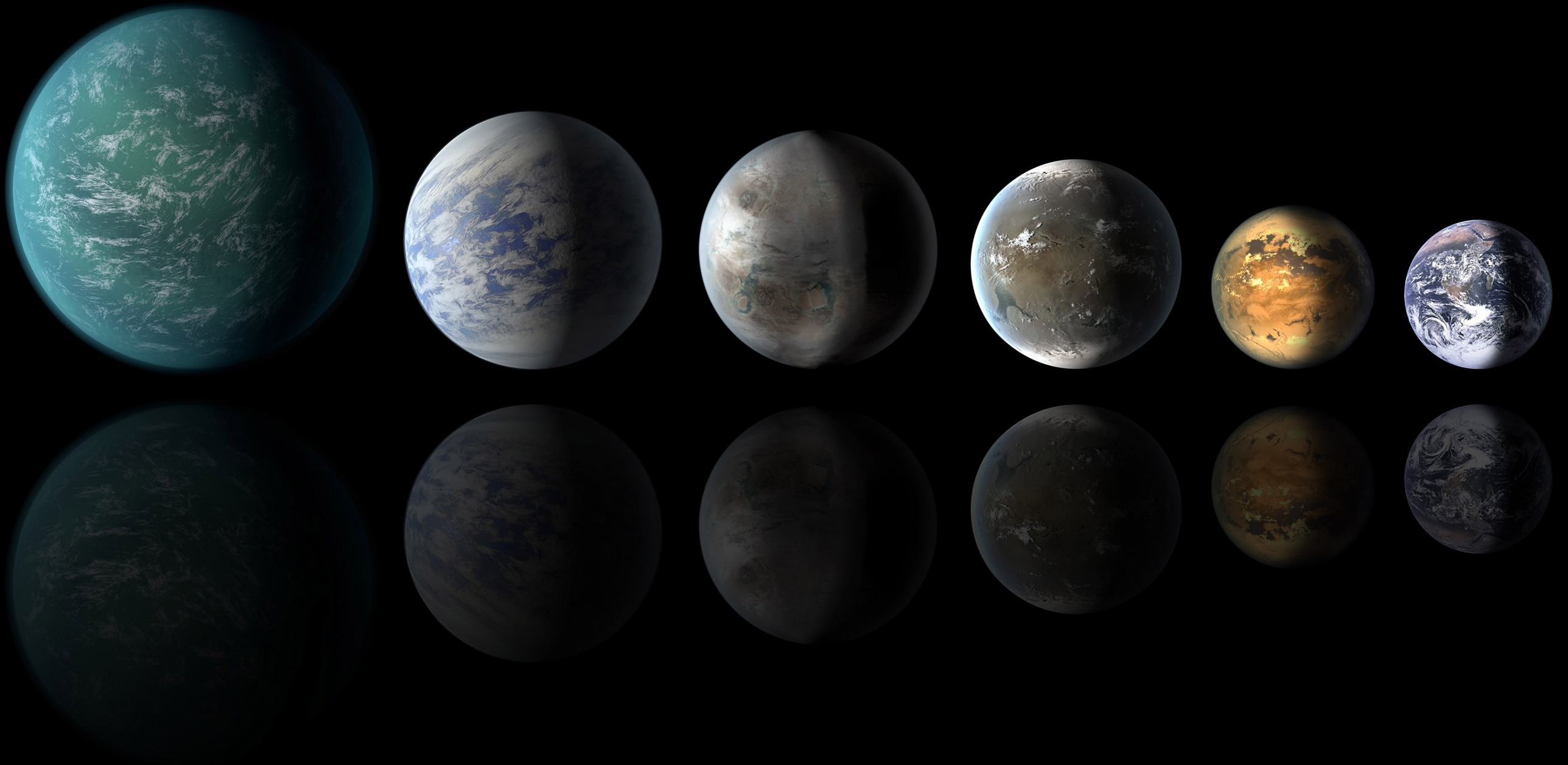
\* many caveats

Kepler Space Telescope



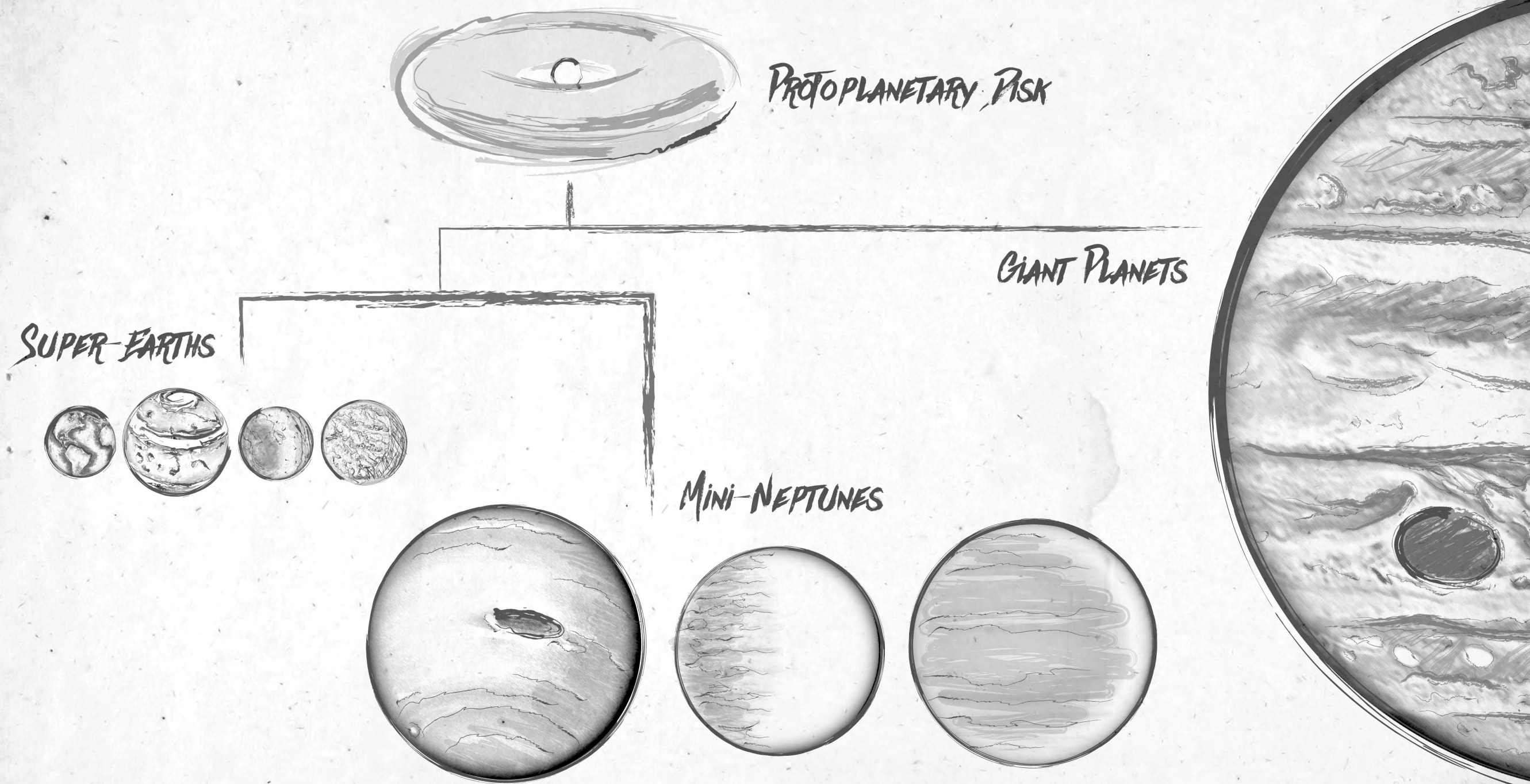
Images: NASA

# Exoplanets are Diverse





# Some Order in the Diversity





# Solar System not a Universal Template



# What makes a planet 'Earth-like'?

Size

Mass

Density

Composition

Atmosphere



Water

Simple Life

Complex Life

Coffee, Sushi, ...

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# SETI

“Search for Extraterrestrial Intelligence”

# Jargon

- See ad hoc committee on SETI nomenclature report (Wright, Sheikh, Almár, Denning, Dick, Tarter, 2018)
- Also:
  - Almár (2008, *Acta Astronautica*, 68, 351)
  - Denning (2008, NASA-SP-2009-4802 Ch. 3 pp.63–124)
  - Wright (2018, arXiv:1803.06972)

# “SETI”

Originally a radio program at NASA

Now the name of the field generally

CETI: “Communication...”

SETT: “...technology”

SETEE: “...electrical engineers”

Let's not quibble.



# “Technosignatures”

Apparently coined by Jill Tarter in 2006

Makes analogy to biosignatures explicit

Inclusive of all signs of technology: radio, laser, waste heat, whatever.

“Produces fewer antibodies” at NASA (per Michael New)

Search for technosignatures = SETI

# Loaded terms

“Civilization”

“Colonization”

“Alien”

“Alien race”

If you use these, be aware that they have meanings you may not intend

Mean what you write and write what you mean.

# Searches for Alien Life in Astrobiology

## Biosignatures

## Technosignatures (SETI)

### Communication

### Artifact

### "Nature-plus"

### METI

#### Exoplanets:

Atmospheric gases  
Surface reflection spectra

#### Solar System:

Microfossils  
Molecular Biomarkers

#### NIROSETI:

pulsed lasers  
continuous lasers

#### Radio:

carrier wave  
broadband

#### Other carriers

(X rays, v's,  
GW's, etc.)

#### Waste heat:

Dyson spheres  
Type III galaxies

#### Radio:

radar  
propulsion

#### Solar System:

Probes  
Structures

#### Exoplanets:

Pollution  
Waste heat

#### Transits:

Arnold  
megastructures

Disappearing  
stars

"Tickling"  
Cepheids

Ġ red spirals

Przybylski's Star

Organized pulsar  
arrays

Strong radio  
transmission

Embedded  
messages in leaked  
emission

Voyager records  
Pioneer plaques



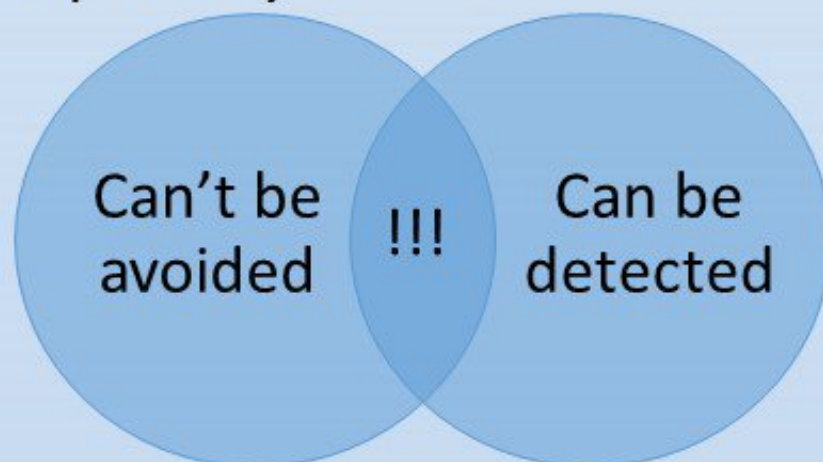
# Technosignature Axes of Merit

With so many potential technosignatures to look for, how do we compare them?

OR

Given finite resources, what do we prioritize?

Inspired by...



Practical

Search can only be done in the far future

Search is costly

Search has no ancillary benefits

Search can be done now

Search is cheap

Search has many ancillary benefits

Scientific

Short-lived

Long-lived

Ambiguous

Unambiguous

High Extrapolation from Earth 2000 Tech

Low Extrapolation from Earth 2000 Tech

Contrived/Specific

Inevitable

Non-detectable

Detectable

Information-Poor

Information-Rich

Best Ideas

S. Sheikh, with input by D. Kipping, A. Frank, S. Walker

# Dyson's First Law of SETI Investigations

“Every search for alien civilizations should be planned to give interesting results even when no aliens are discovered.”





# Communication SETI

## The birth of Radio SETI

1960 — Cocconi & Morrison suggest interstellar communication via radio waves

### SEARCHING FOR INTERSTELLAR COMMUNICATIONS

By GIUSEPPE COCCONI\* and PHILIP MORRISON†

Cornell University, Ithaca, New York

NO theories yet exist which enable a reliable estimate of the probabilities of (1) planet formation; (2) origin of life; (3) evolution of societies possessing advanced scientific capabilities. In the absence of such theories, our environment suggests that stars of the main sequence with a lifetime of many billions of years can possess planets, that of a small set of such planets two (Earth and very probably Mars) support life, that life on one such planet includes a society recently capable of considerable scientific investigation. The lifetime of such societies is not known; but it seems unwarranted to deny that among such societies some might maintain themselves for times very long compared to the time of human history, perhaps for times comparable with geological time. It follows, then, that near some star rather like the Sun there are civilizations with scientific interests and with technical possibilities much greater than those now available to us.

\* Now on leave at CERN, Geneva.

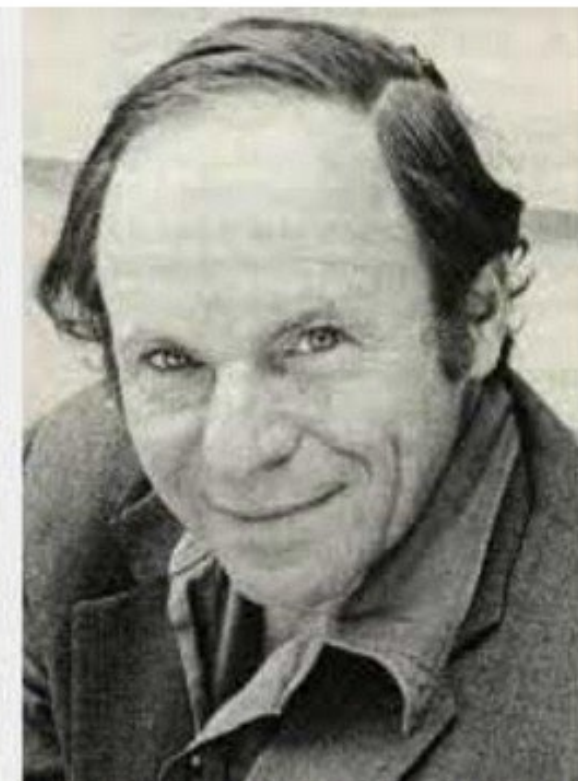
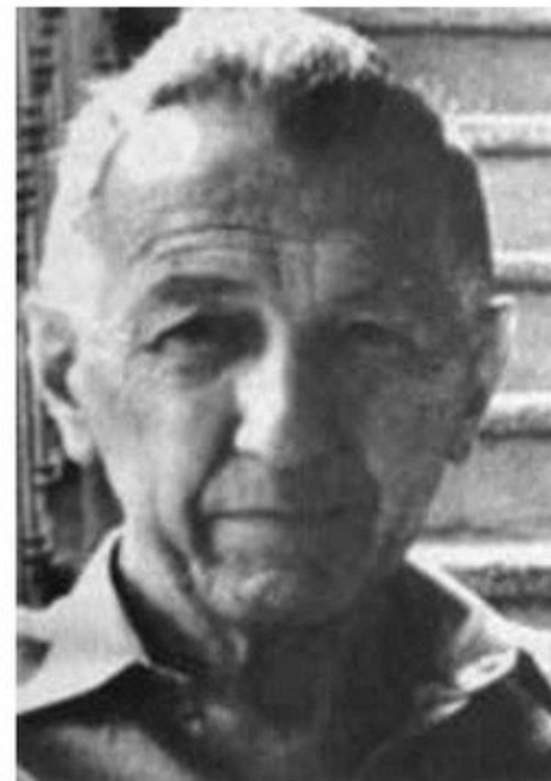
† Now on leave at the Imperial College of Science and Technology, London, S.W.7.

To the beings of such a society, our Sun must appear as a likely site for the evolution of a new society. It is highly probable that for a long time they will have been expecting the development of science near the Sun. We: they established a channel would one day become known look forward patiently to the Sun which would make society has entered the channel. What sort of a channel would

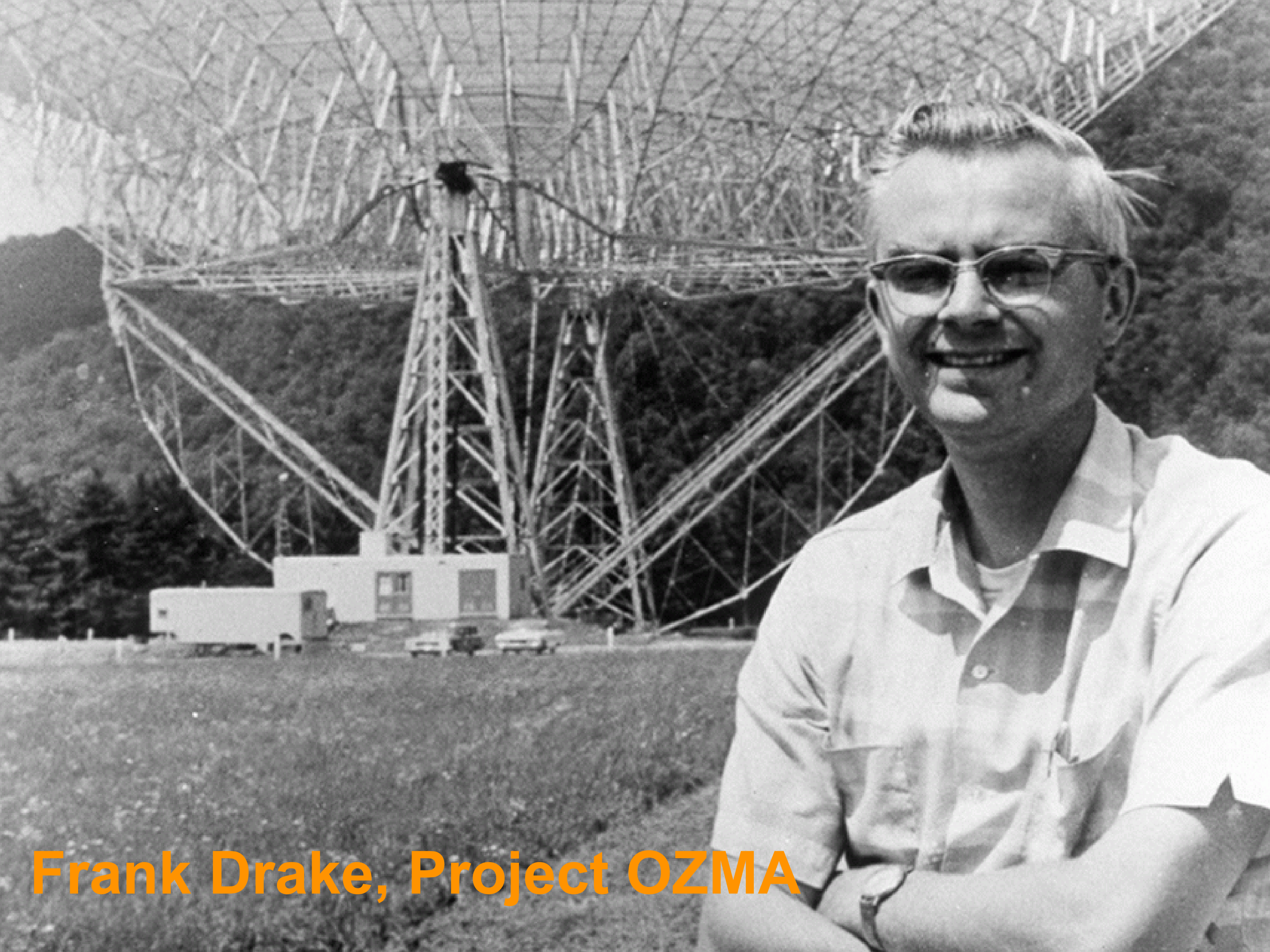
#### The Optimal

Interstellar communication plasma without dispersion is practical, so far as we magnetic waves.

Since the object of this is to find a newly evolved that the channel used with minimum burden of frequency







**Frank Drake, Project OZMA**



 **BREAKTHROUGH**  
**LISTEN**



**SETI**  
● **INSTITUTE**

# Not just narrowband carrier waves

- Radar
- Pulsed emission
- Spillover from point-to-point communications
- Byproducts of propulsion / high energy work
- Huge range of necessarily non-natural waveforms can be searched for “agnostically”



# Communication SETI

## The birth of Laser SETI

1961 — Schwartz & Townes point out lasers work well, too

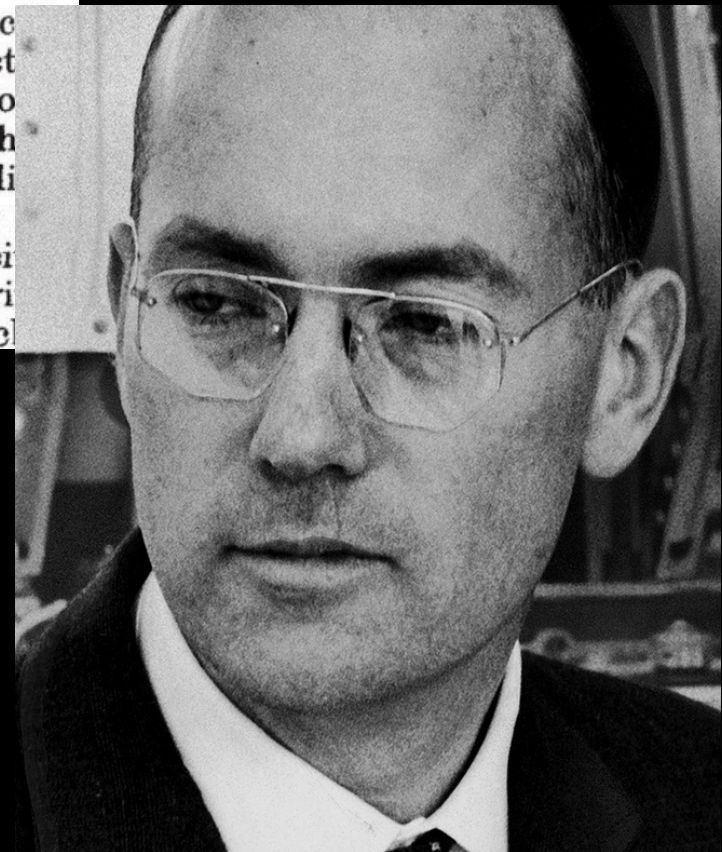
### INTERSTELLAR AND INTERPLANETARY COMMUNICATION BY OPTICAL MASERS

By DR. R. N. SCHWARTZ and PROF. C. H. TOWNES\*

Institute for Defense Analyses, Washington, D.C.

**L**ONG-RANGE communication by radio-waves is already well known, and the possibility of interstellar communication by radio-waves in the microwave region has been suggested in several interesting proposals<sup>1-3</sup> to search for signals from intelligent beings on planets associated with nearby stars. The supposition is that curiosity such as our own would motivate advanced civilizations associated with stars other than our Sun to make determined efforts to

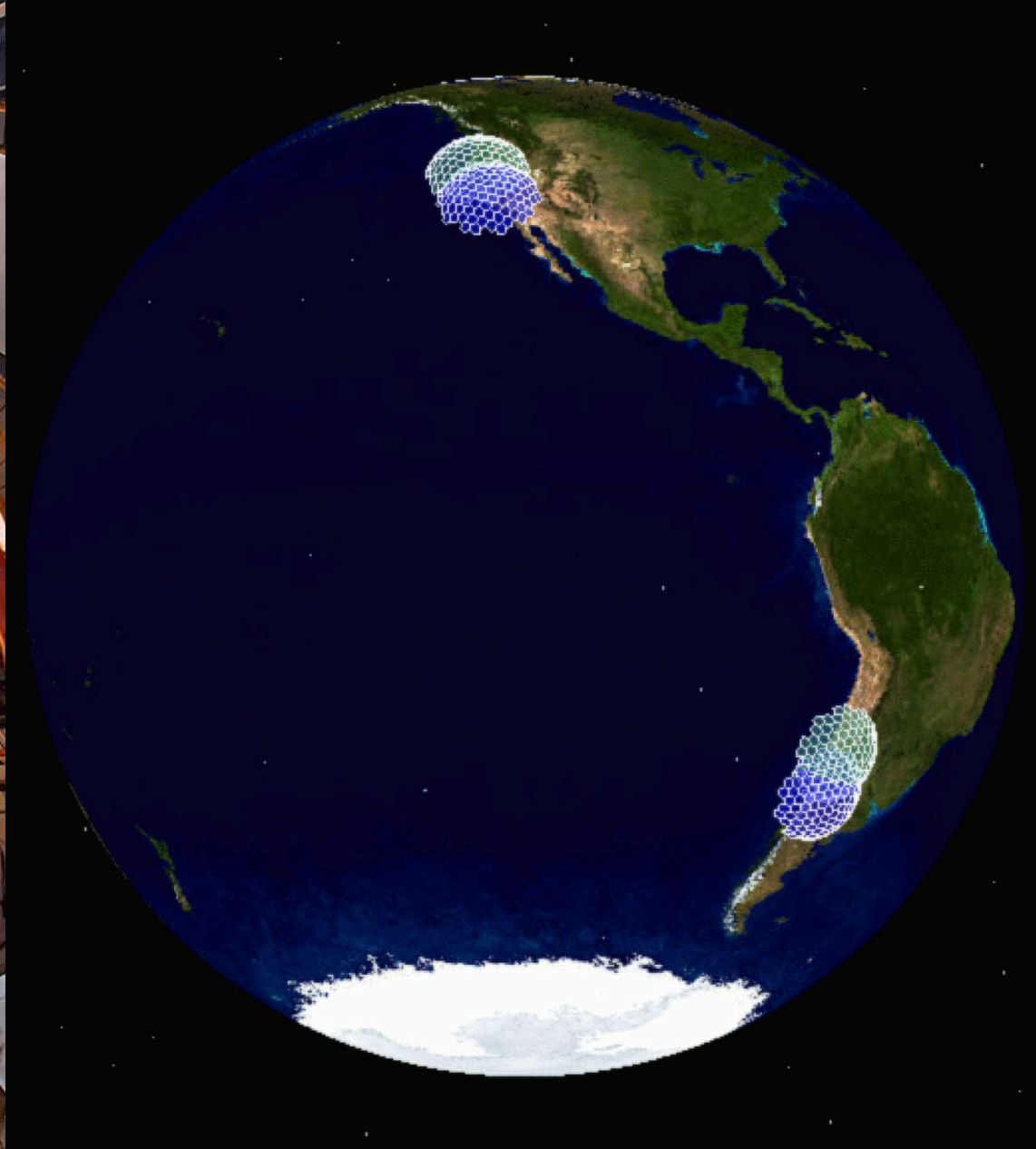
frequency-interval of about 10 kilocycles per sec. The latter case is much closer to theoretical expectations<sup>4</sup> for an ideal maser in so far as coherence is concerned. There seems to be no general reason, other than the necessary dissipation of power, why solid state optical masers cannot operate continuously at high power and with a short-term monochromaticity close to theoretical expectations, or hence with frequency-widths very much less than 1 megacycle.



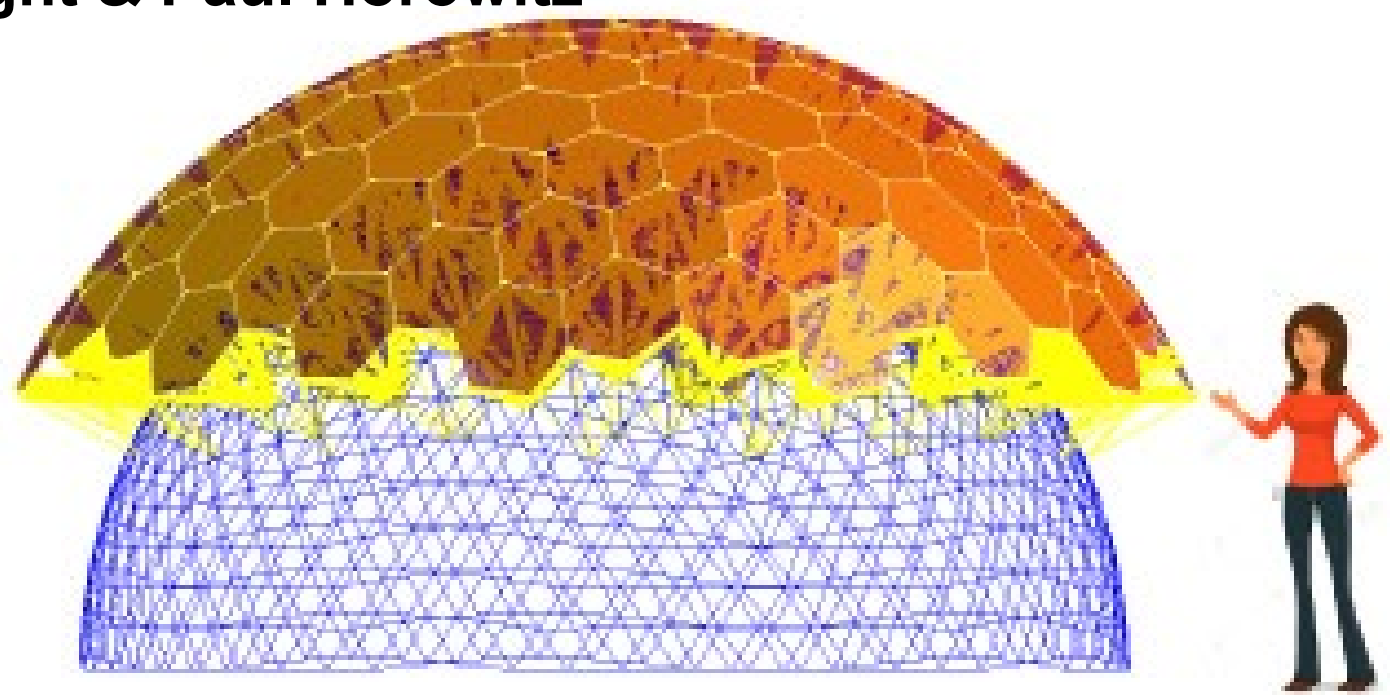
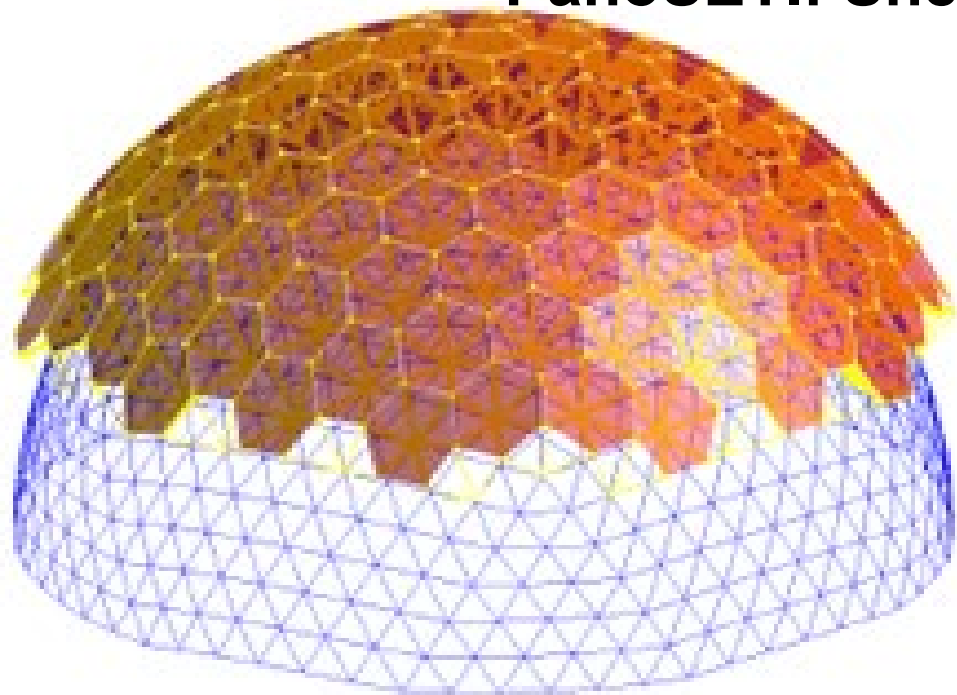
# Laser advantages

- Pulsed lasers provide very high bandwidth
- Optical transmissions can have very high gains
- IR lasers have great dust penetration
- Can be concentrated in time or frequency domains



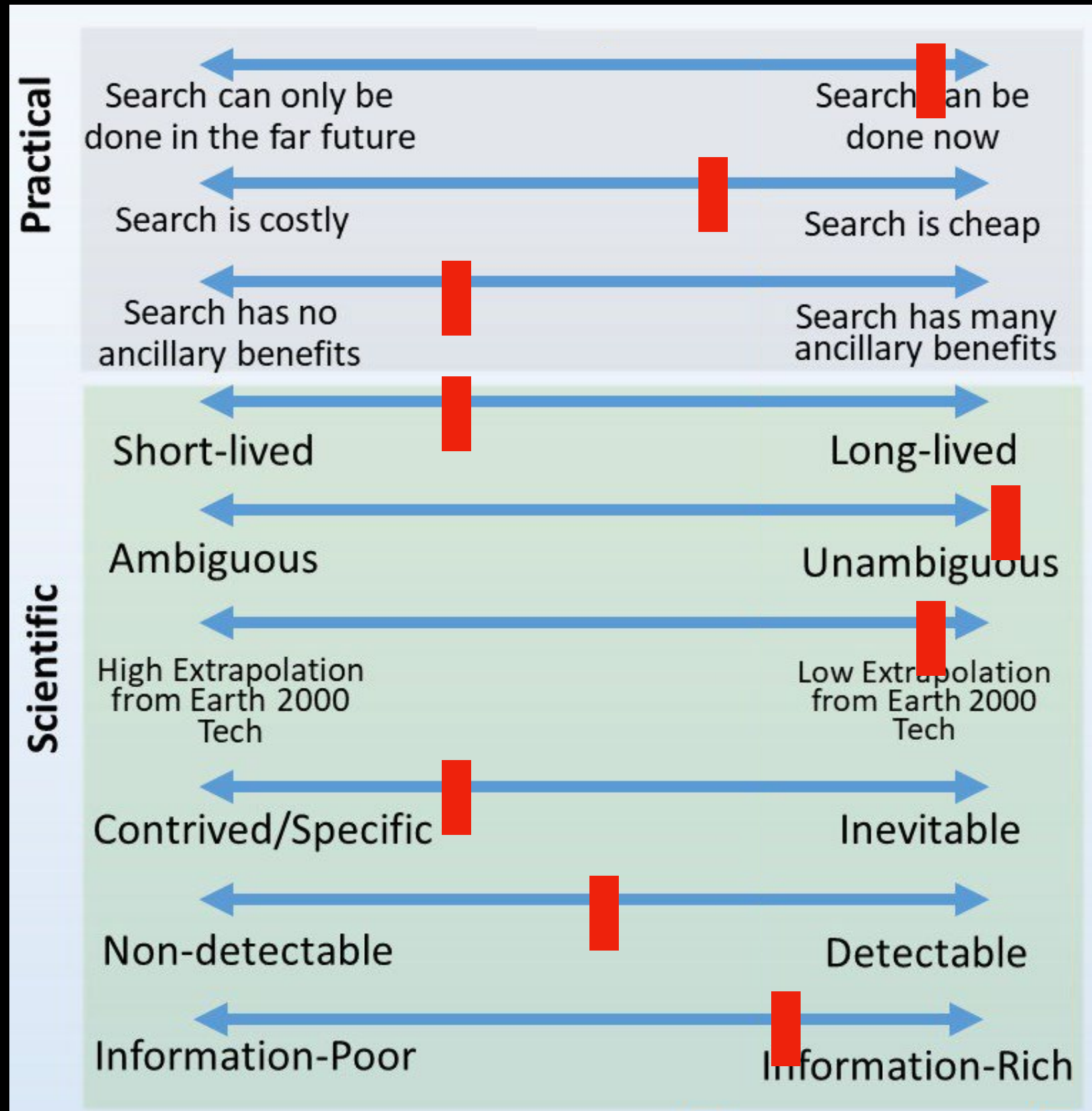


**PanoSETI: Shelley Wright & Paul Horowitz**





# Radio/Laser SETI



# Artifact SETI



Dyson (1960)

## Search for Artificial Stellar Sources of Infrared Radiation

*Abstract.* If extraterrestrial intelligent beings exist and have reached a high level of technical development, one by-product of their energy metabolism is likely to be the large-scale conversion of starlight into far-infrared radiation. It is proposed that a search for sources of infrared radiation should accompany the recently initiated search for interstellar radio communications.

Cocconi and Morrison (1) have called attention to the importance and feasibility of listening for radio signals trans-

here” really just any orbital m  
collects or generates energy

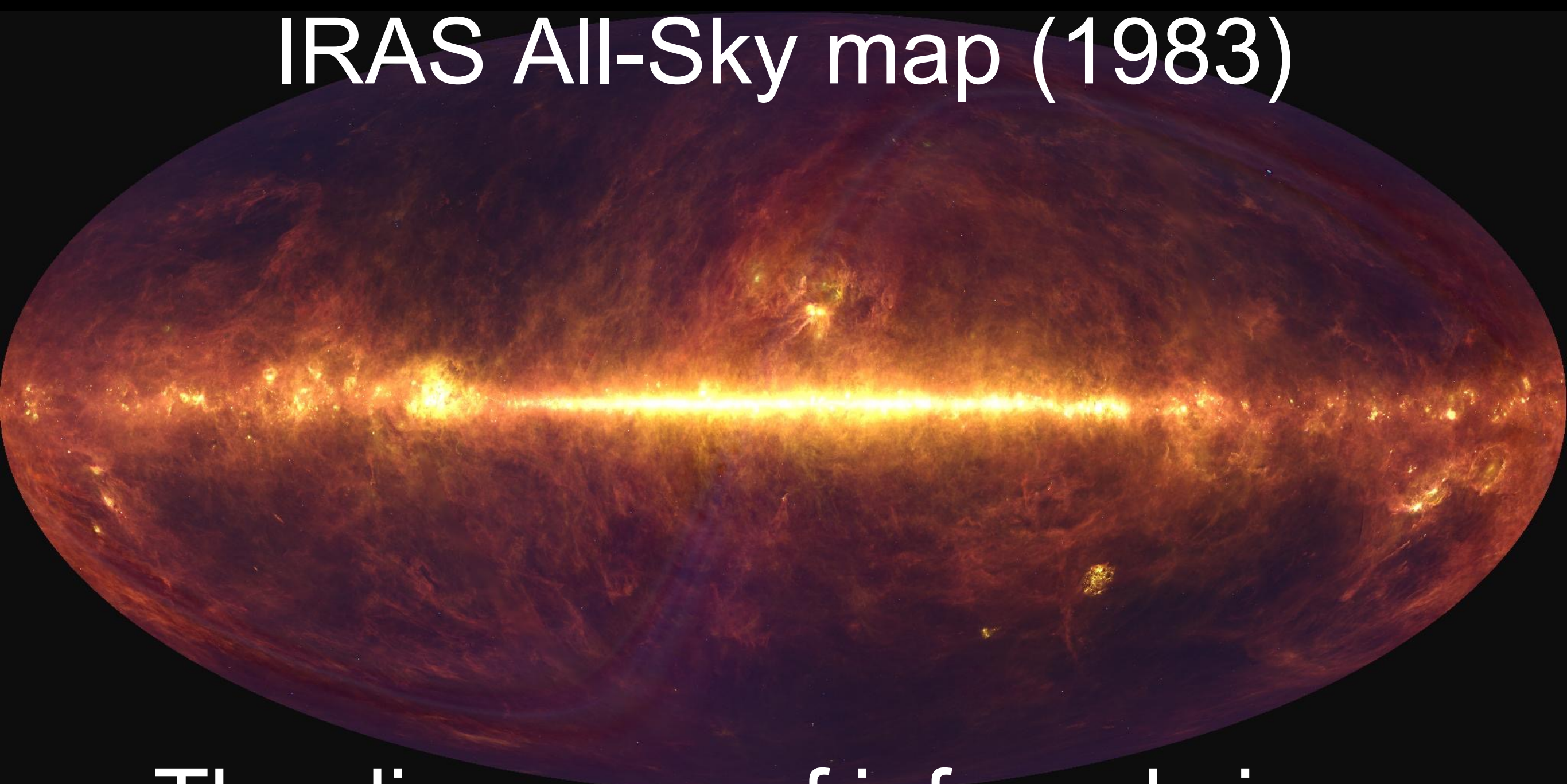
Emission of low-entropy “waste” heat in IR is required by thermodynamics

Sensitive to *almost all* technological energy use in the system

Primary confounder is dust which has similar MIR properties



# IRAS All-Sky map (1983)

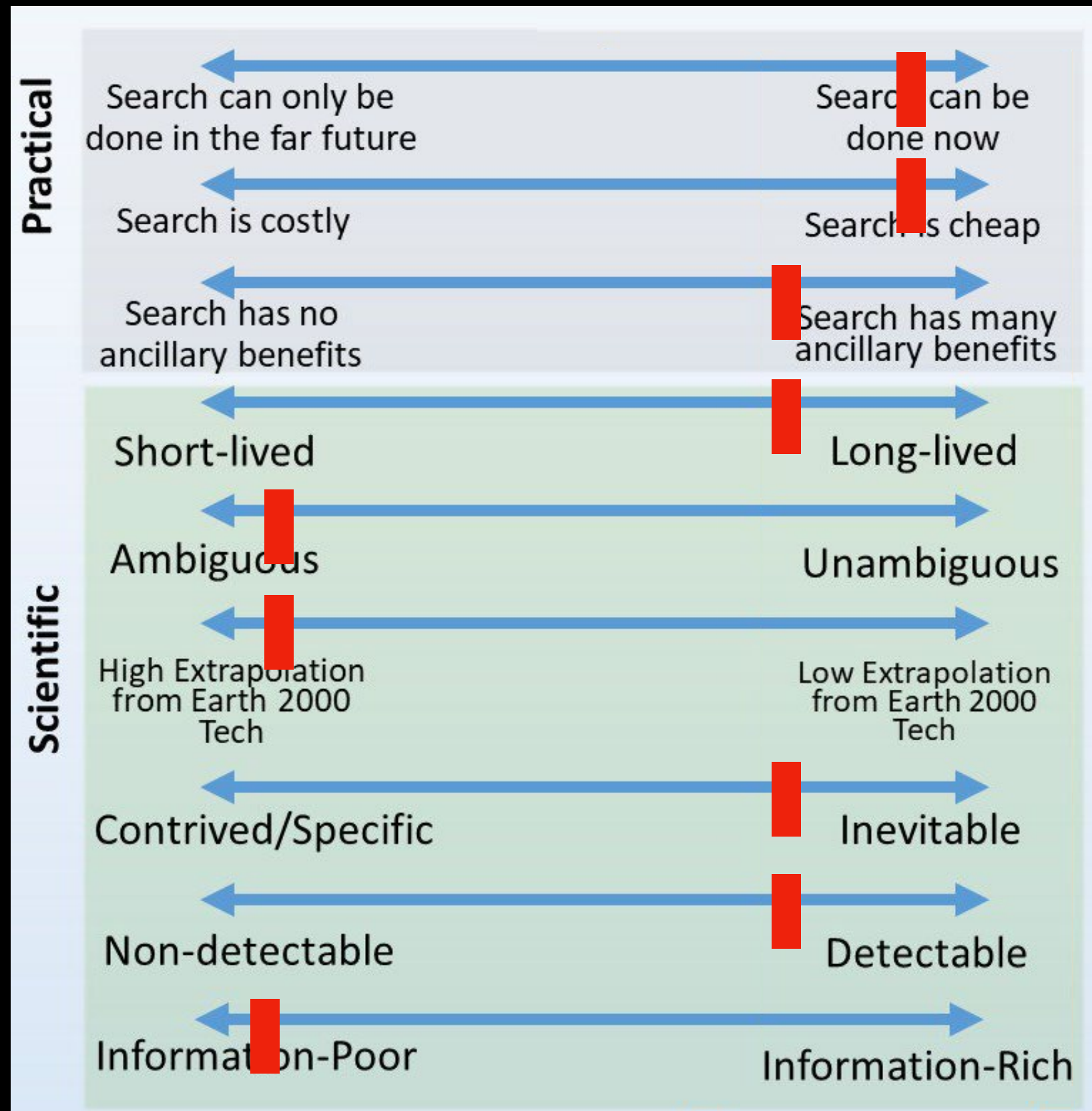


The discovery of infrared cirrus complicated Dyson sphere searches.

# Some work

- Jugaku & Nishimura (2004): Nearby stars with *IRAS*
- Carrigan (2009): All-sky *IRAS* non-detection
- Annis (1999): Pan-galactic Dyson spheres via low optical surface brightness (Kardashev type III)
- Griffith et al (2015):  $\hat{G}$  extragalactic search upper limits with *WISE*
- Arnold (2005): *Kepler* as megastructure hunter (see also Boyajian's Star)

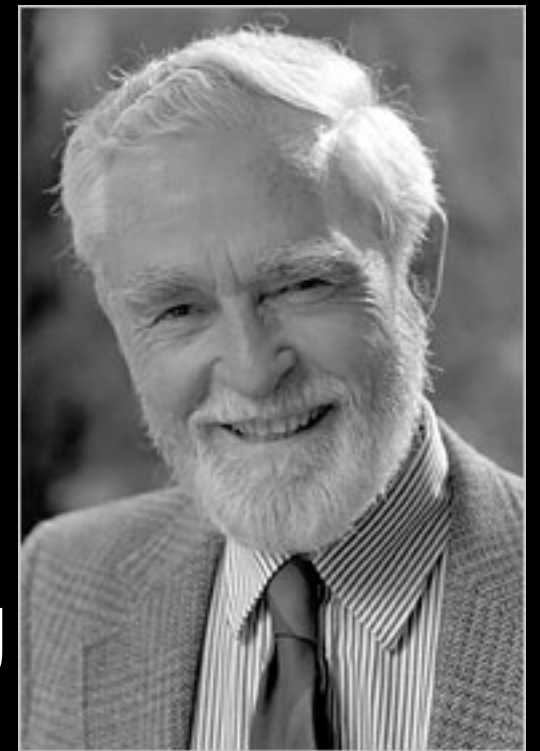
# Waste Heat / Transit SETI





# Solar System SETI

Bracewell (1960)



## COMMUNICATIONS FROM SUPERIOR GALACTIC COMMUNITIES

By PROF. R. N. BRACEWELL

Radioscience Laboratory, Stanford University, California

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ha

SINCE Morrison and Cocconi<sup>1</sup> published the suggestion that there might be advanced societies elsewhere in the Galaxy, superior to ourselves in technological development, who are beaming transmissions at us on a frequency of 1,420 Mc./s., Drake<sup>2</sup> has described equipment under construction to look

planetary systems. Beyond their immediate neighbourhood, it might be feasible for them to spray some number of suitable stars, say, one thousand, with modest probes. Each probe would be sent into a circular orbit about one of the thousand stars, at a distance within the habitable zone of temperature

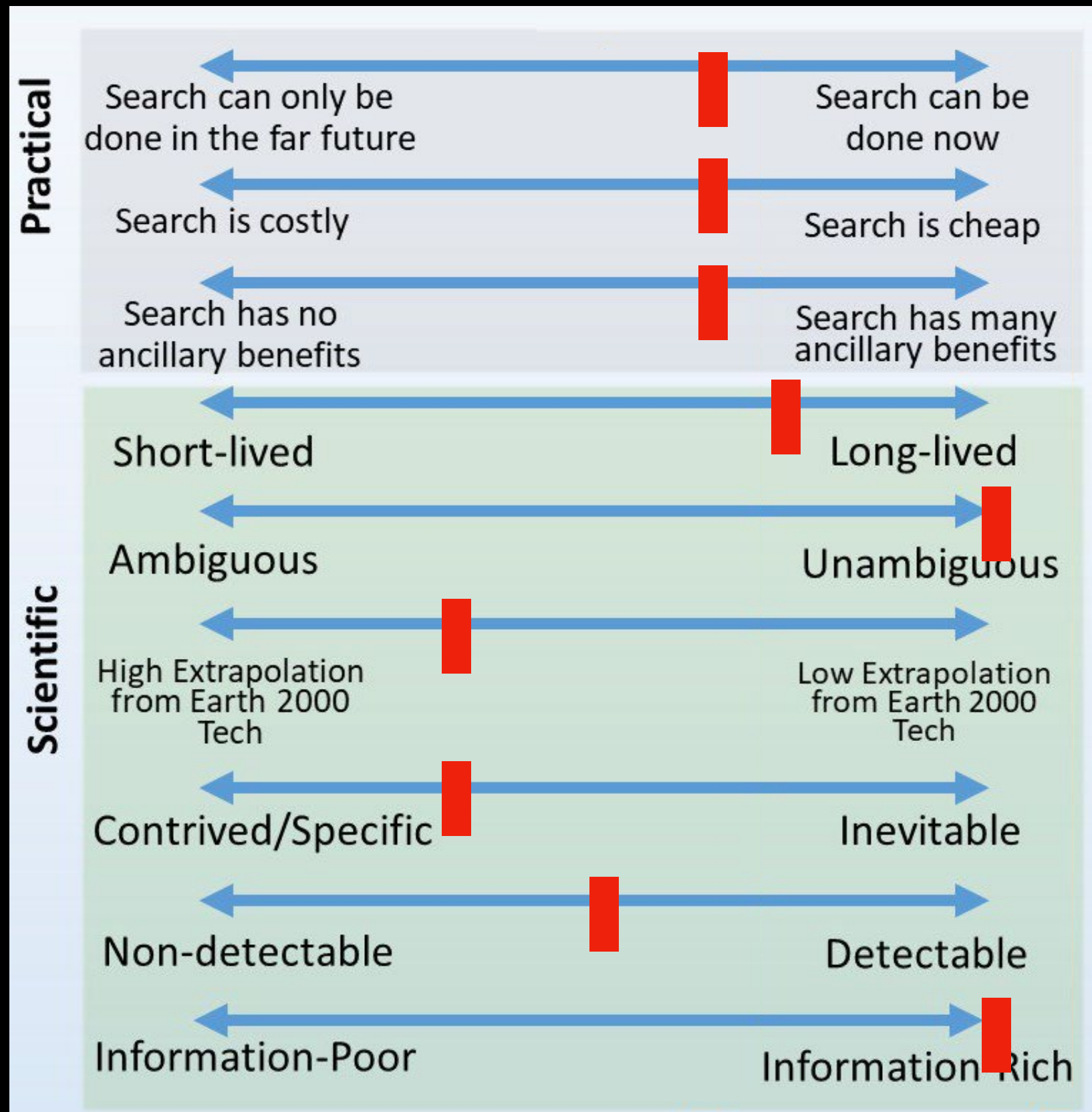
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Probes might be sent for many purposes (contact, exploration) and might “pass through” or “lurk” indefinitely

# Solar System SETI

- Intelligent Martians were once legitimate scientific hypothesis
- Mariner showed no cities: now idea that technology *ever* existed there is taboo
- Claim that there are proveably zero probes in the Solar System *now* has been used to argue ETIs must not exist
- In reality, completeness is very low (Haqq-Misra & Kopparapu 2012)
- Moon has best high-resolution imaging for searches (Davies & Wagner 2013)
- Free-floating artifacts (Freitas 1980, 1983)
- 'Oumuamua (Bialy & Loeb 2018)

# Solar System SETI





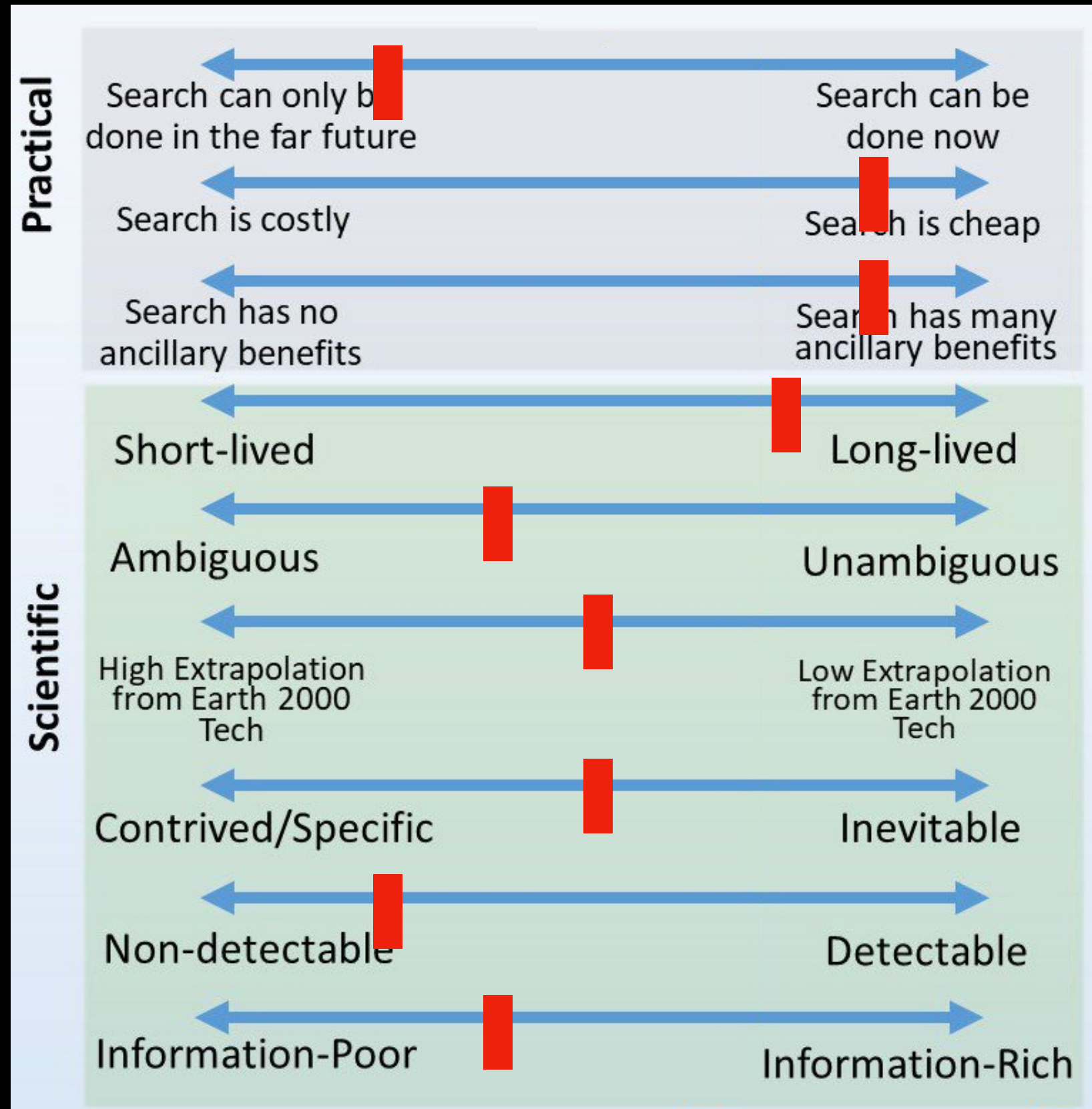
# Exoplanet atmospheric pollution

Very little work on this so far

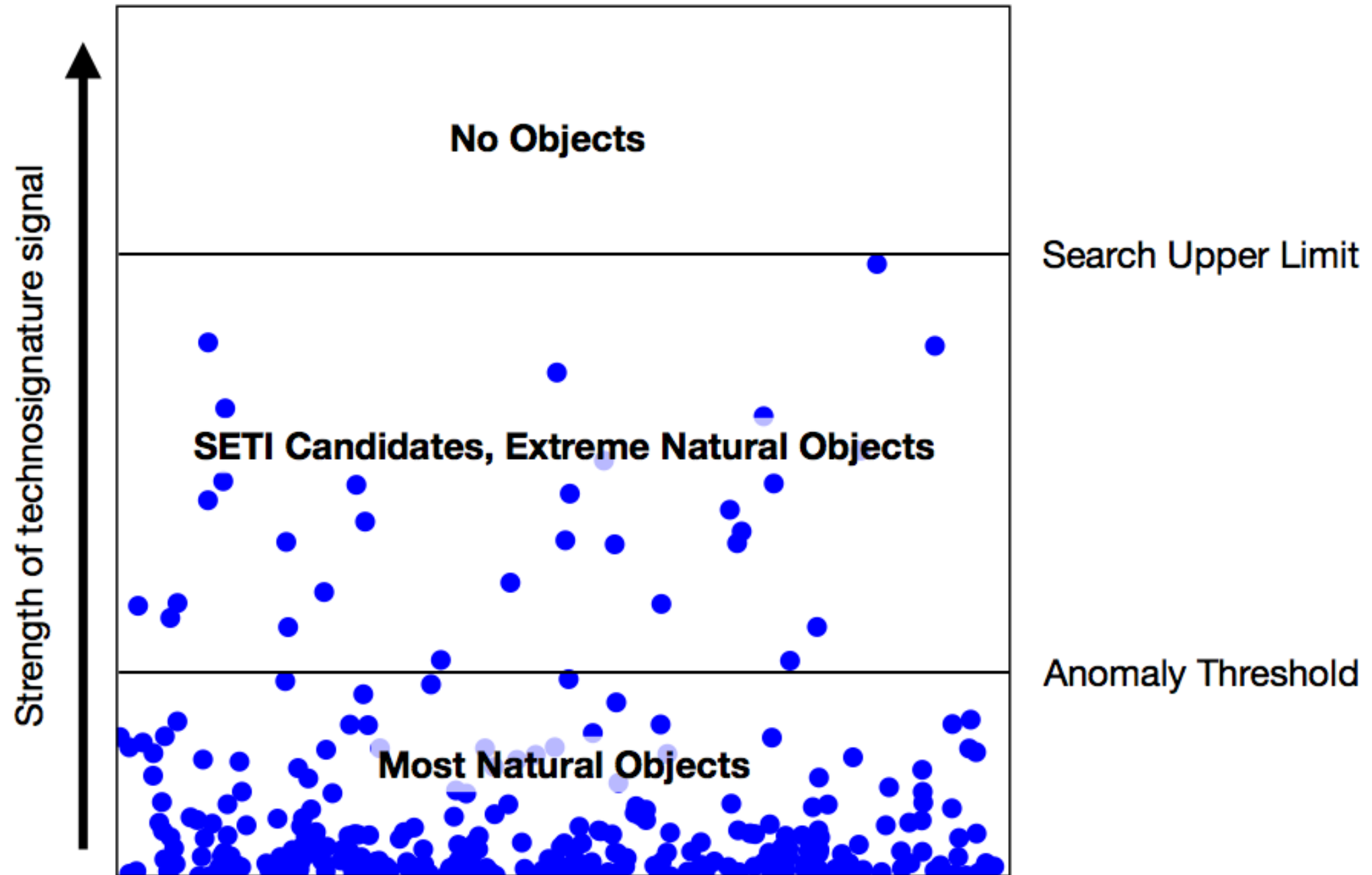
Natural synergy with hunts for biosignatures

- Lin, Gonzalez Abad, & Loeb (2014)
- Lingam & Loeb (2017)

# Exoplanet Pollution SETI



# Upper limits in SETI



# History of Database Mining for Technosignatures

| Database | Searching for?                          | Authors                      | Methods                                | Explicitly SETI? |
|----------|---|------------------------------|--|------------------|
| IRAS     | Dyson Spheres                           | Jugaku & Nishimura, Carrigan | Least-squares fitting to expected data | Yes              |
| WISE     | Dyson Spheres                           | Griffith/Wright et al.       | Least-squares fitting to expected data | Yes              |
| Kepler   | Dyson Spheres                           | Giles & Walkowicz            | Machine Learning (clustering)          | No               |
| SDSS     | Kardashev Type III                      | Baron & Poznanski            | Unsupervised Random Forest             | No               |
| [n/a]    | Shape Agnostic Artificial Radio Signals | Zhang et al.                 | Self-supervised Deep Learning          | Yes              |

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Przybylski's Star

Organized pulsar  
arrays

Strong radio  
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Embedded  
messages in leaked  
emission

Voyager records  
Pioneer plaques