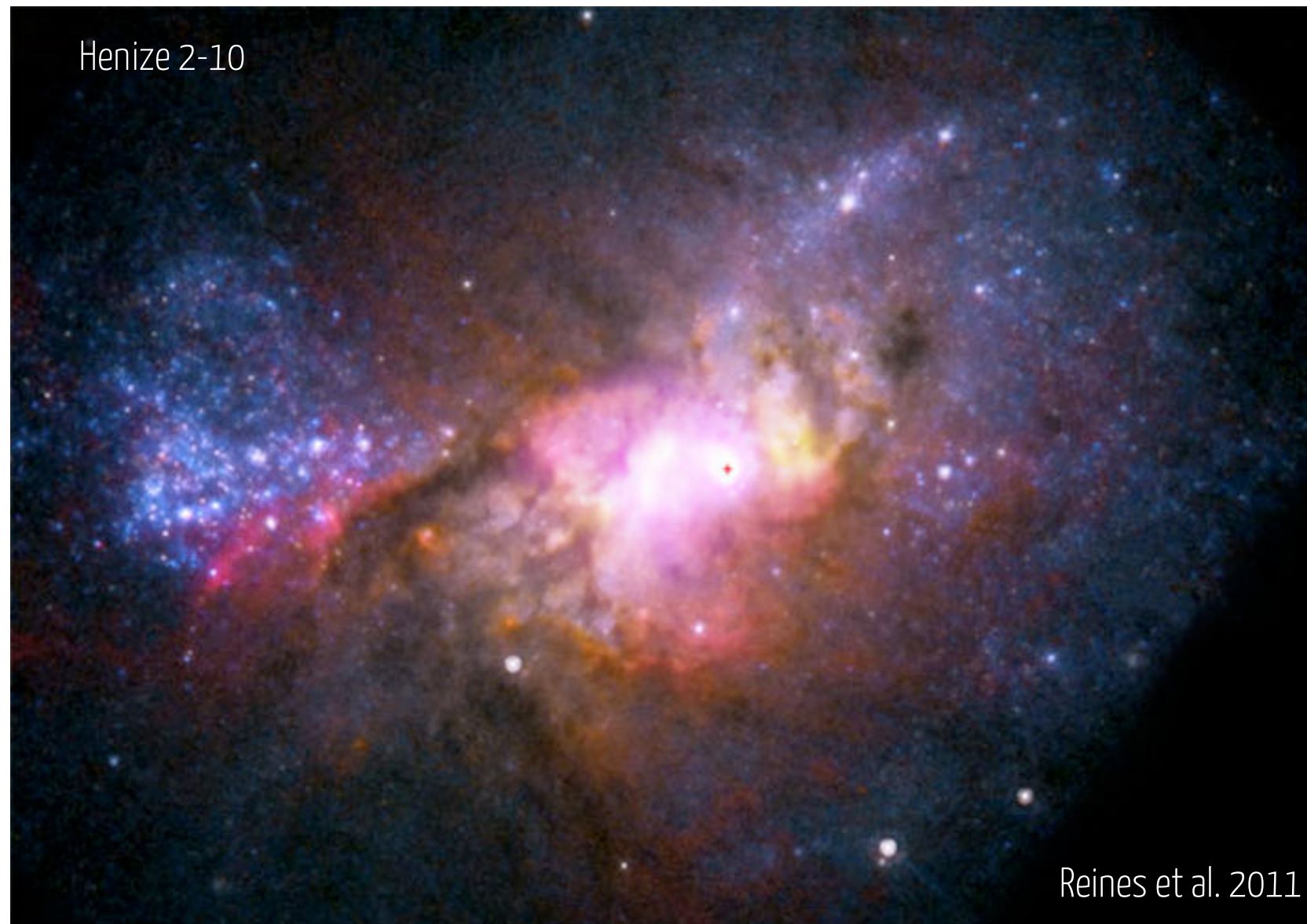
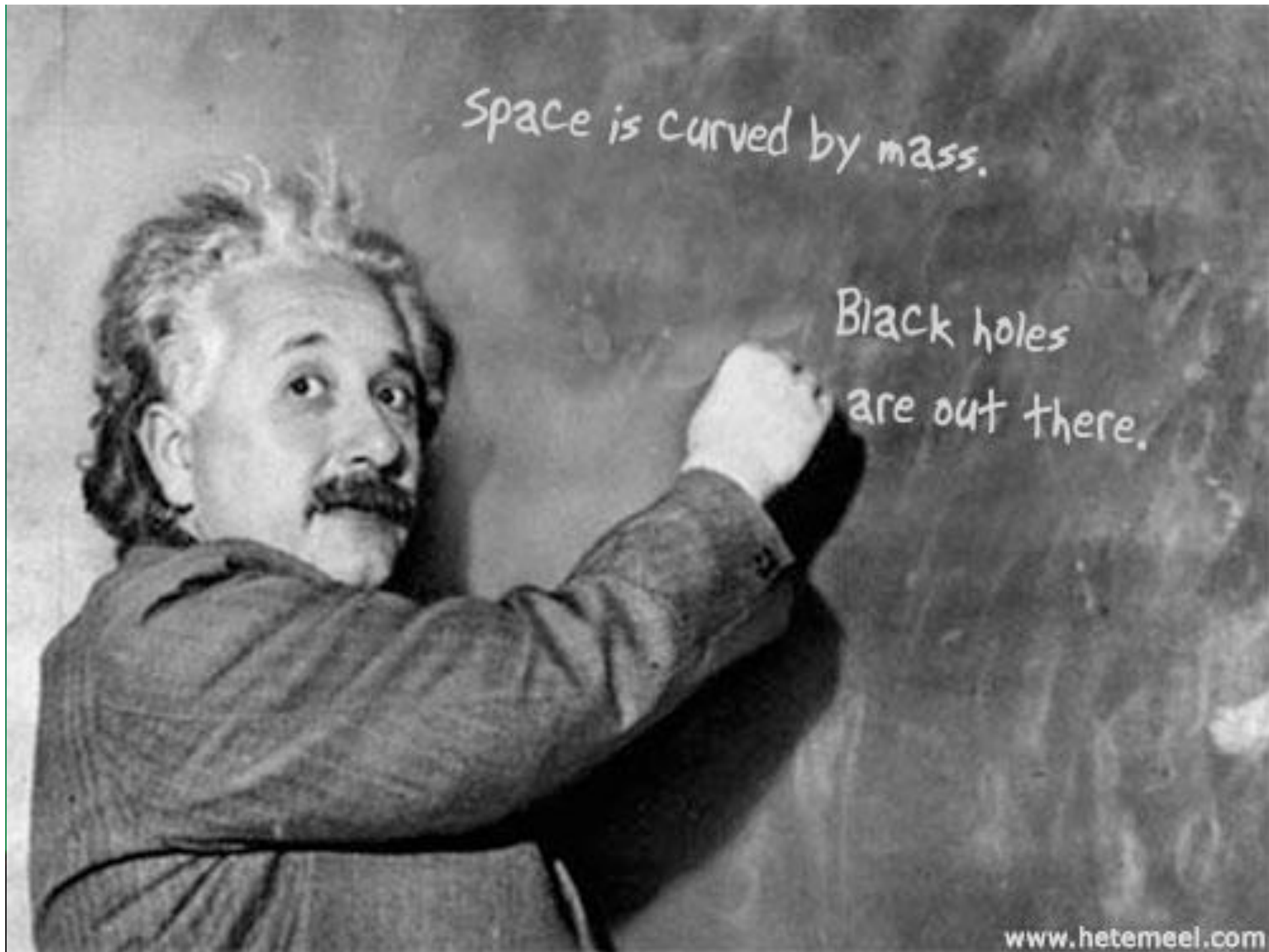


# Black Holes 101



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Ask a physicist: what is a black hole?





**A black hole is an infinite curvature in spacetime**



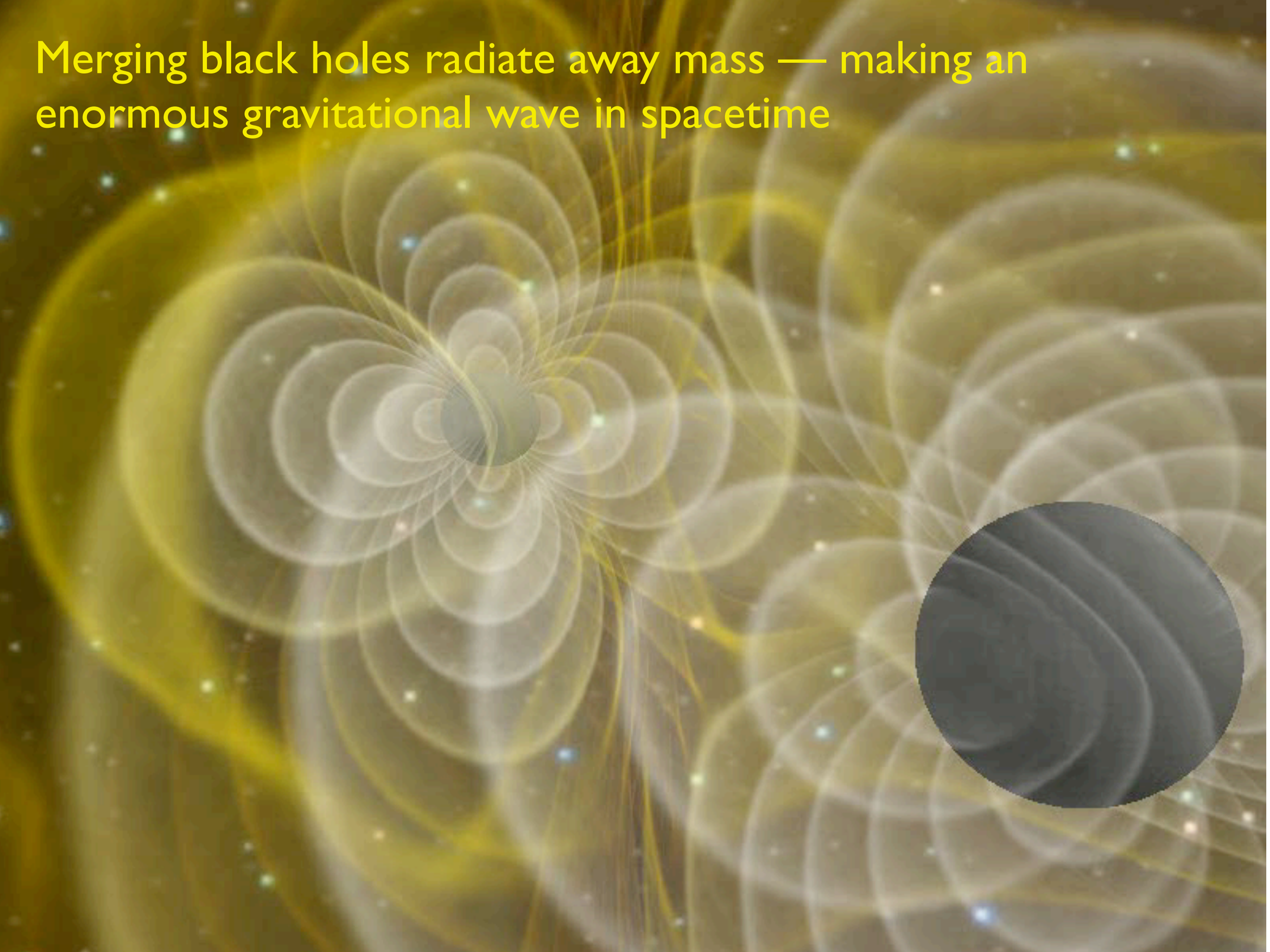




Text

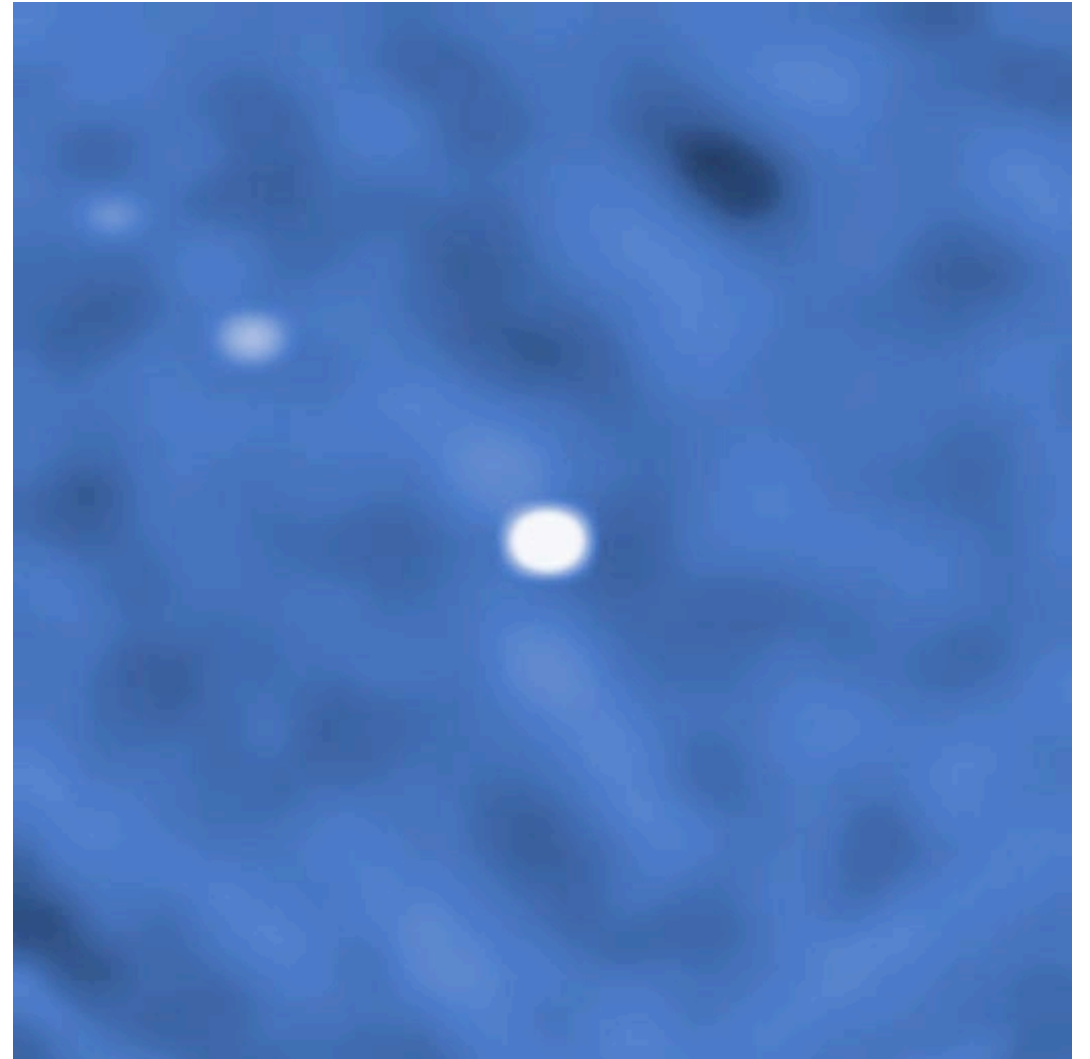


Merging black holes radiate away mass — making an enormous gravitational wave in spacetime





# Ask an astronomer: what is a black hole?



Gas that falls into a black hole heats up and emits high energy photons, is entrained by magnetic fields in a jet.



BLAZAR

RADIO LOUD  
QUASAR

RADIO  
GALAXY

# AGN UNIFICATION

GAS CLOUDS IN  
'NARROW  
LINE  
REGION'

SUPERMASSIVE BLACK HOLE

ACCRETION  
DISC

DUSTY GAS  
CALLED A "TORUS"

"BROAD LINE  
REGION"


SEYFERT II


SEYFERT I



Meg Urry

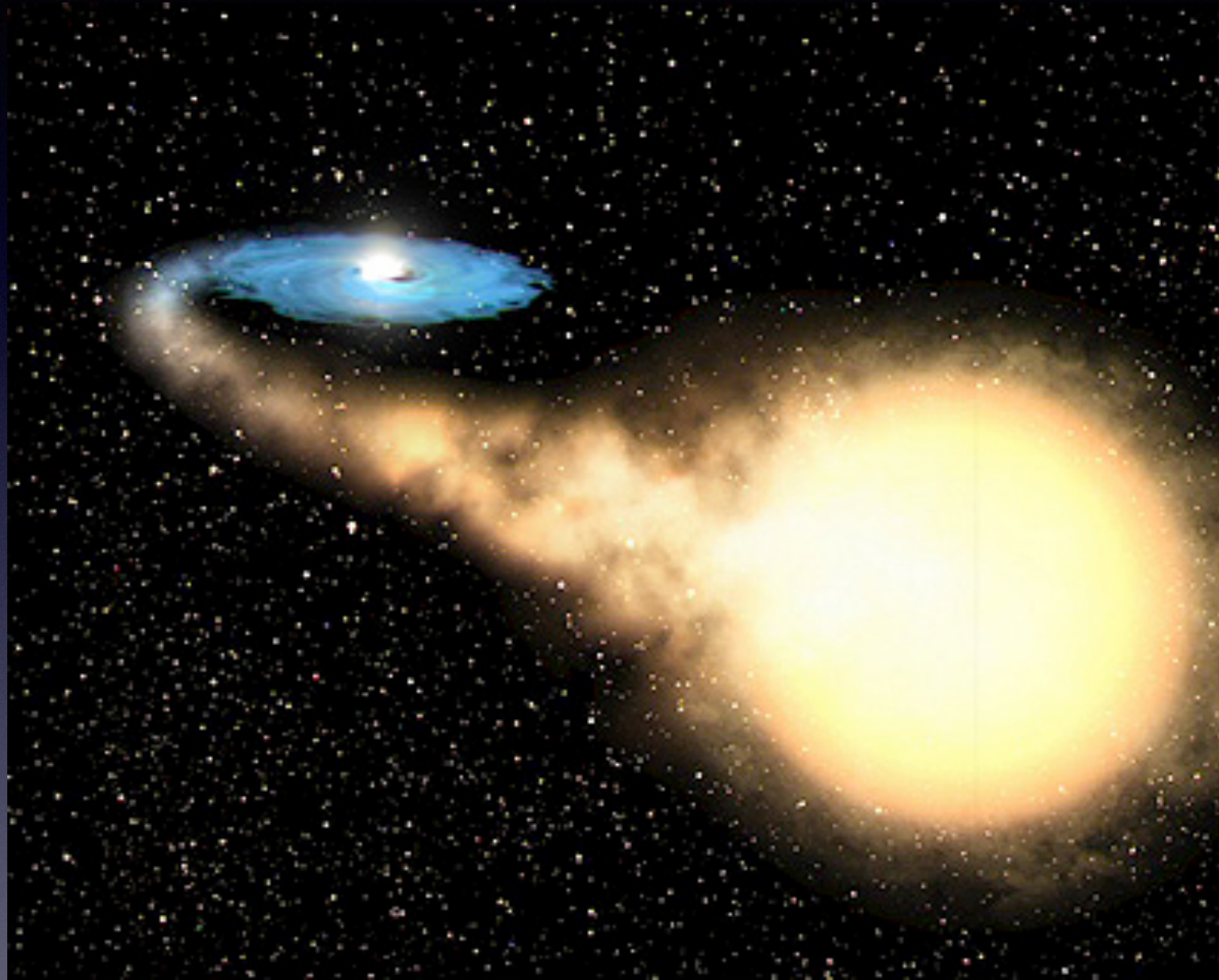
## KEY

 STUFF

 WHAT AN OBSERVER  
SEES



We know there are two types of black holes:

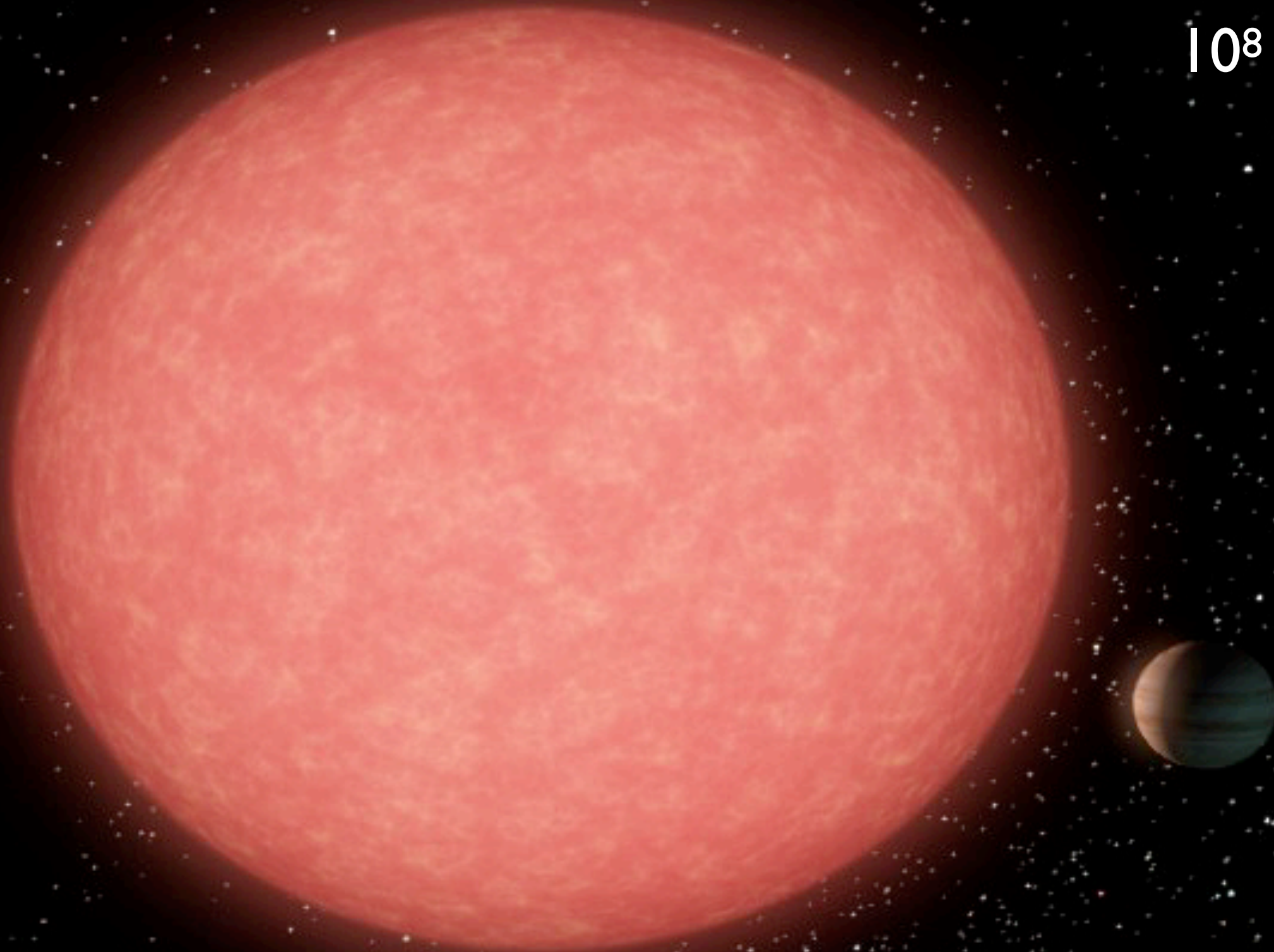


Stellar mass black holes made from massive dead stars



We know there are two types of black holes:

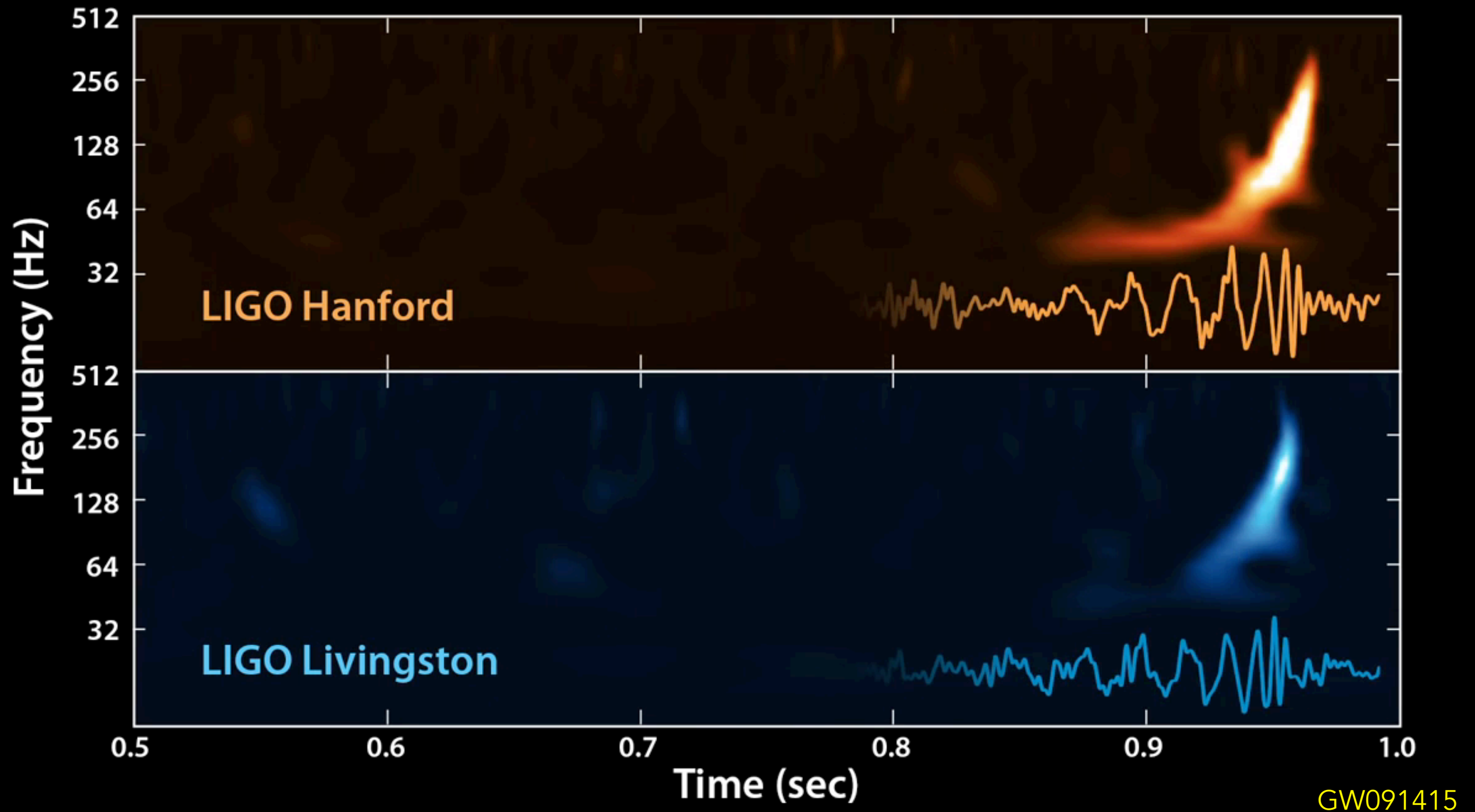
$10^8$  in our galaxy!



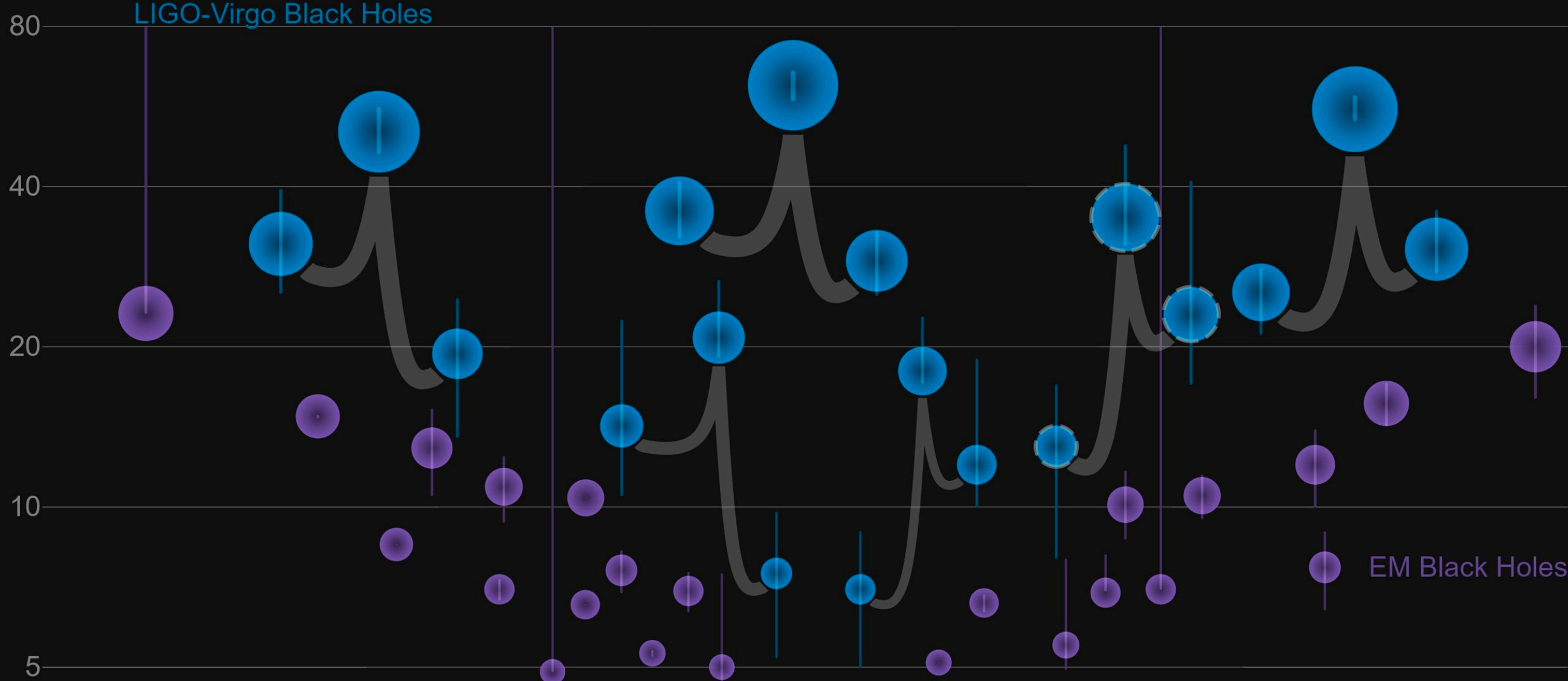
Stellar mass black holes made from massive dead stars



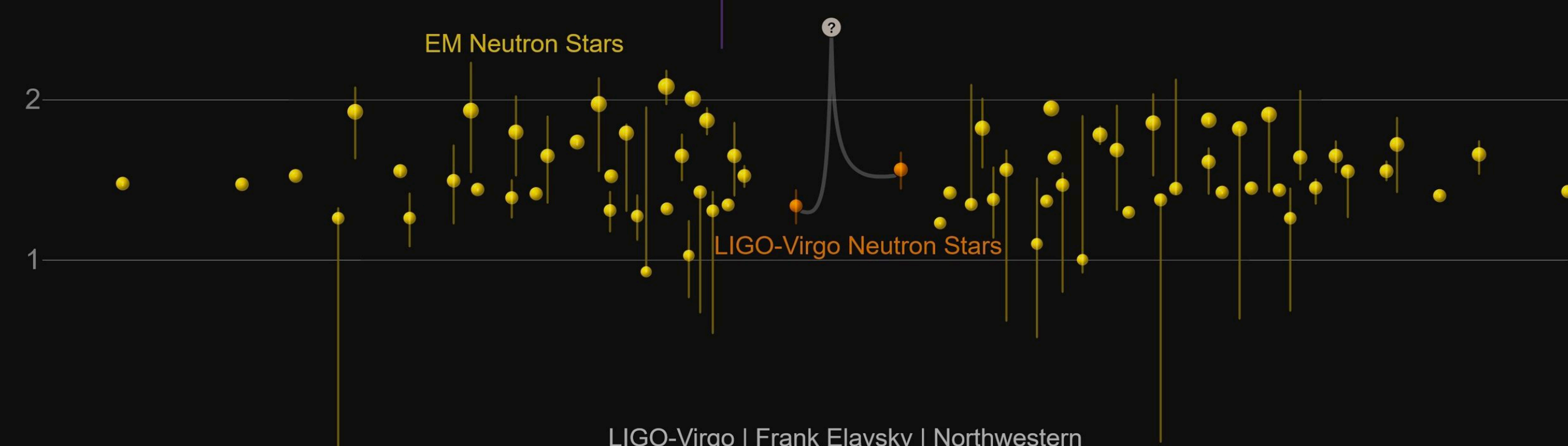
# The chirp heard around the world





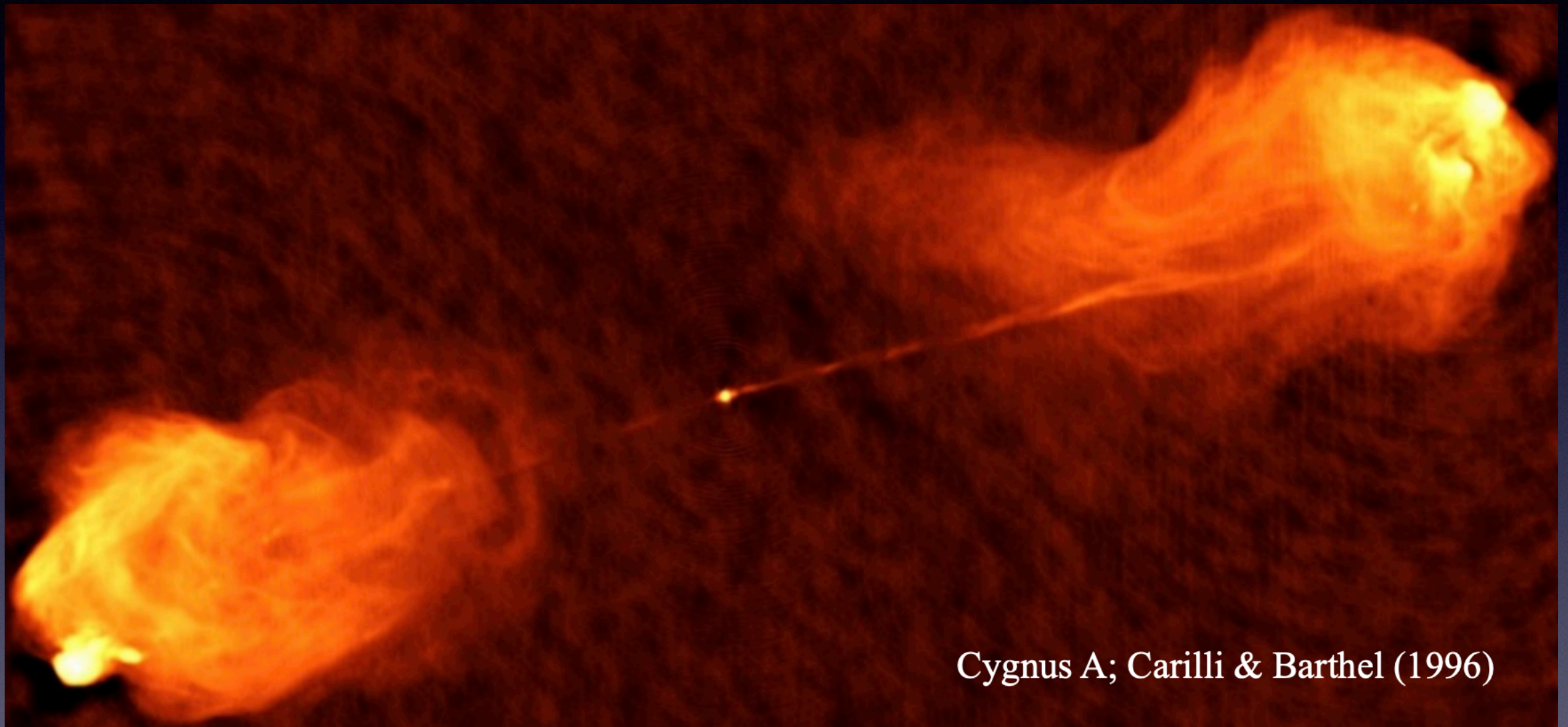


EM Black Holes





We know there are two types of black holes.



Cygnus A; Carilli & Barthel (1996)

Supermassive black holes (SMBHs) lurking at galaxy centers



**Are there other types of black holes?**



# Evidence of an intermediate mass black hole --- in the outskirts of a galaxy

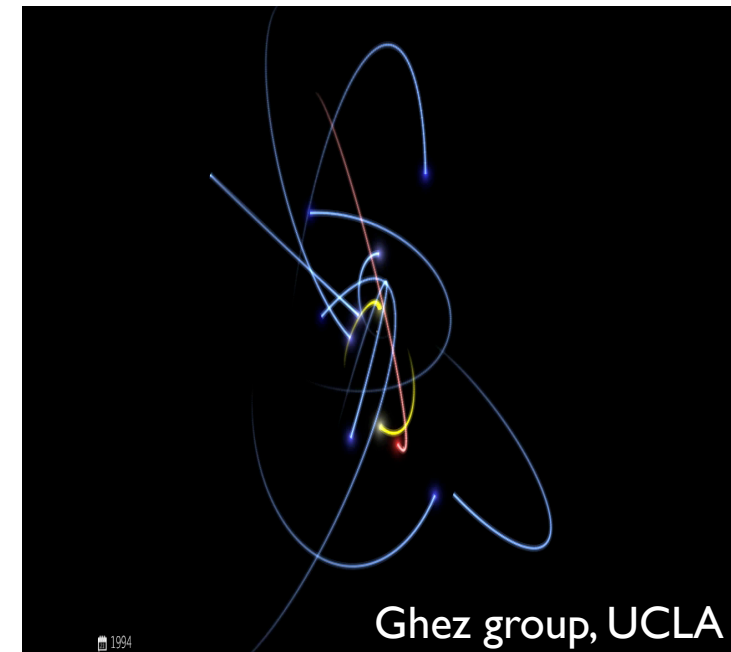
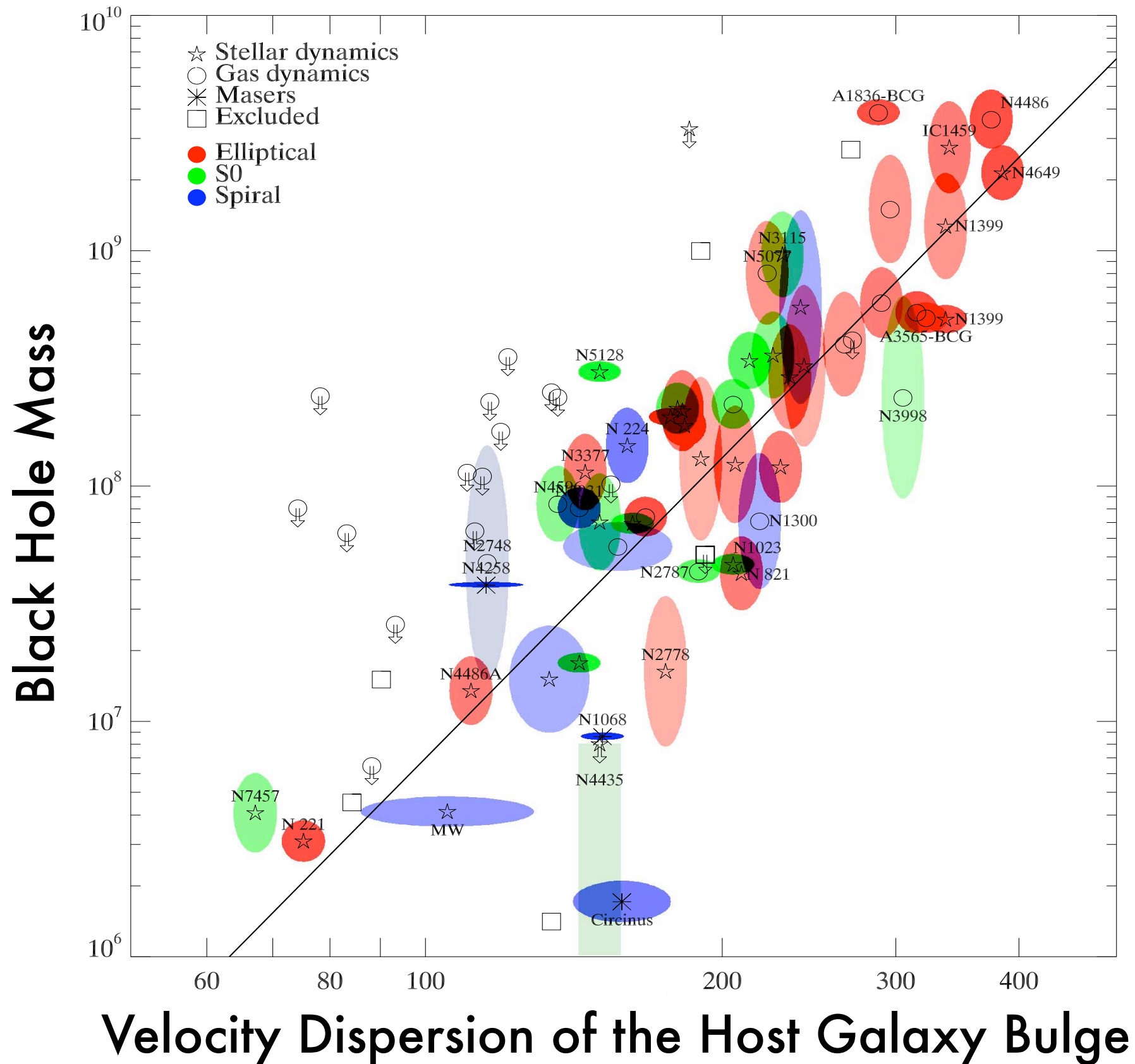
Farrell et al. 2009; 2012



*>500  $M_{\odot}$ , with stellar shroud!*



# A black hole and its galaxy co-evolve when they should be decoupled!



$$R_{\text{sch}} = 2 G M / c^2 = 0(10^{-6}) \text{ pc}$$

$$R_{\text{infl}} = G M / \sigma_0^2 = 0(10^0) \text{ pc}$$

$$R_e = 0(10^3) \text{ pc}$$

Gultekin et al 2009 -- see also Gebhardt et al 2000; Ferrarese & Merritt 2000:  
McConnell+Ma 2013, and work is on-going...

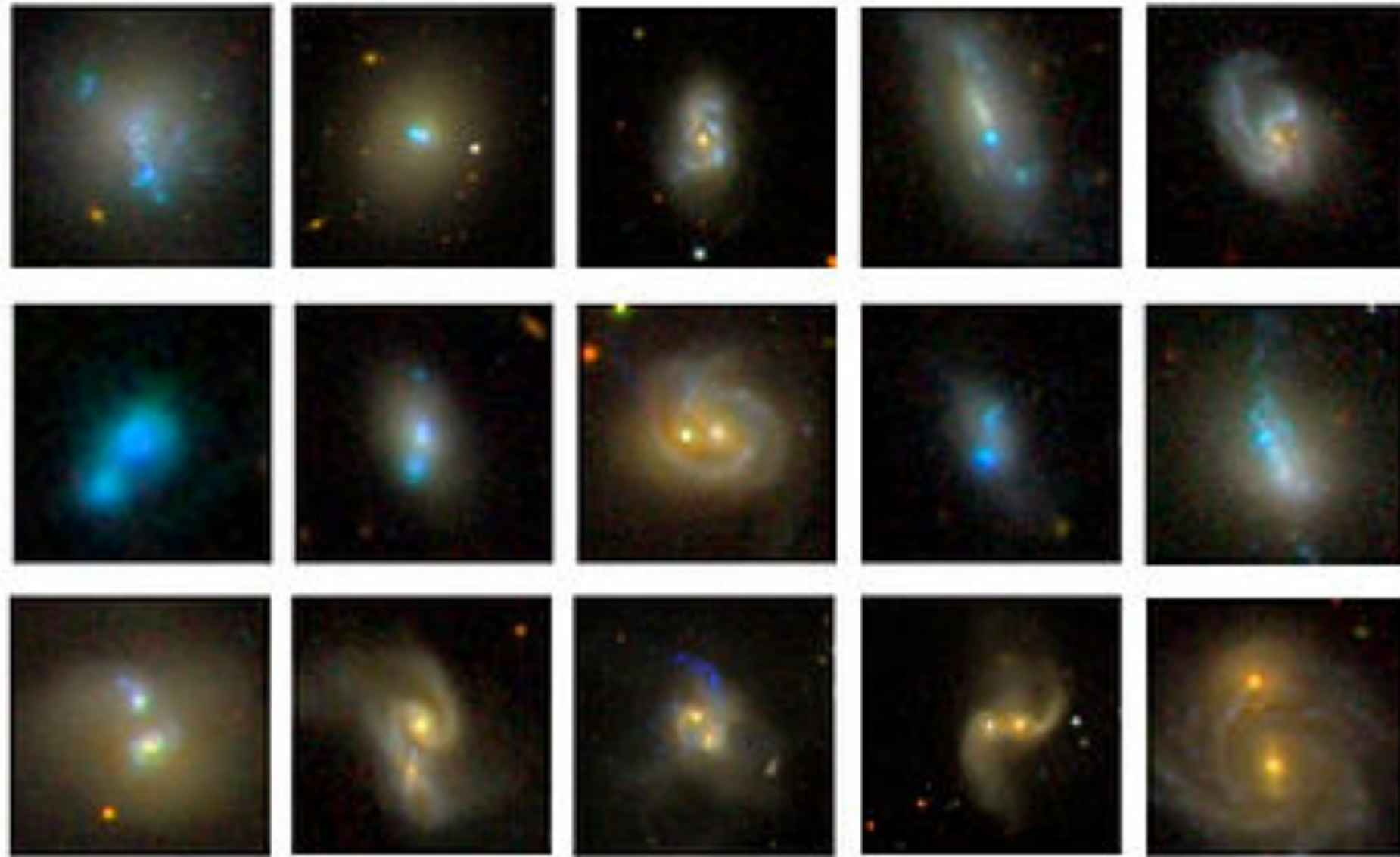


Prevailing paradigm:  
galaxy mergers fuel black holes + black hole feedback regulates galaxy star formation





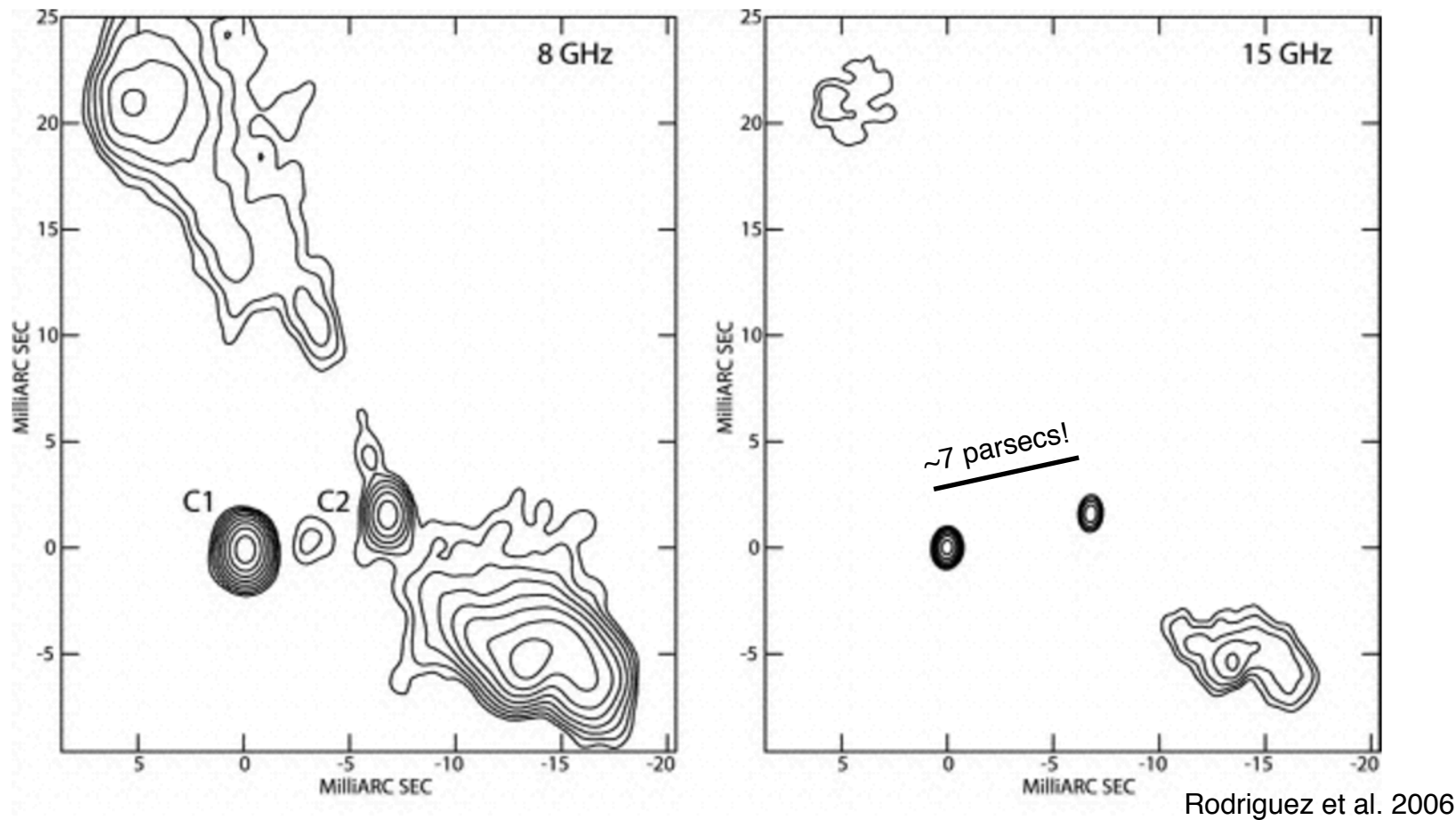
While there are certainly dual AGN,



Mezuca et al. 2014



...there are (arguably) no known binary black holes



Radio very long baseline interferometric search of  
~3100 AGN, only 1 found to be consistent with a  
binary black hole

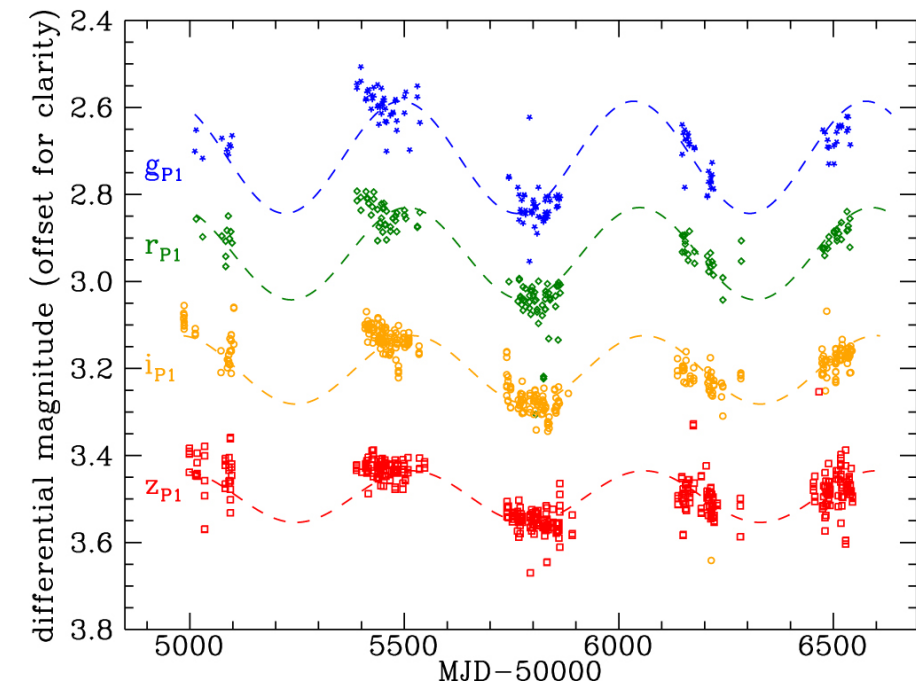
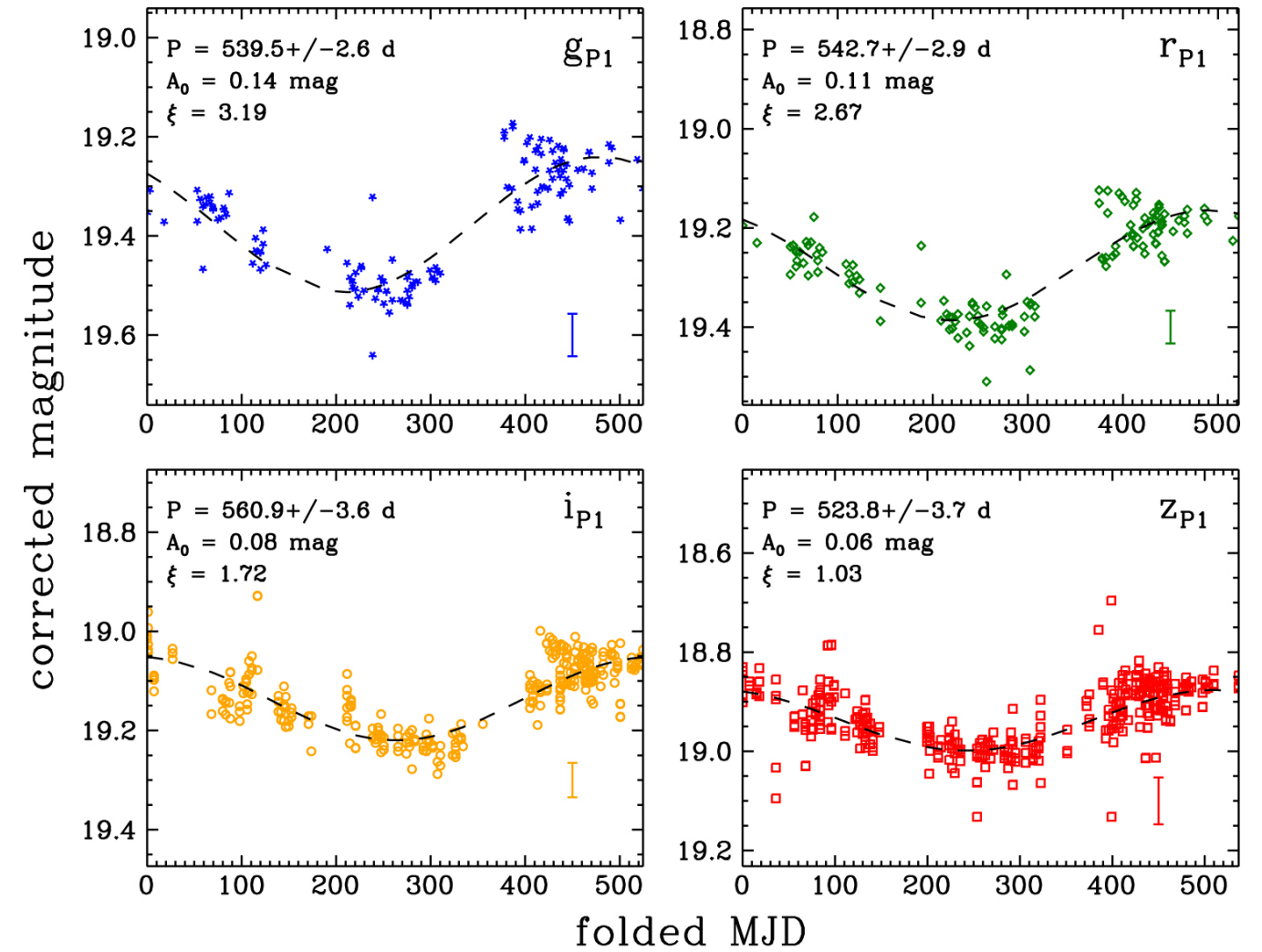
Burke-Spolaor 2011

Stay tuned! Time-domain astronomy will help here...



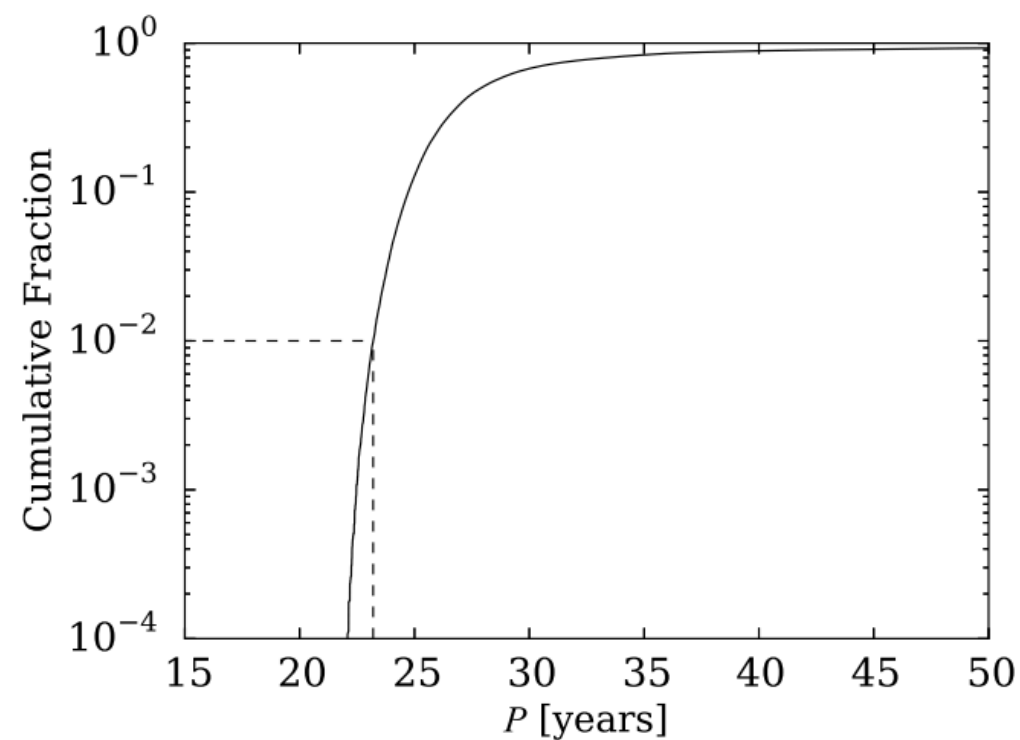
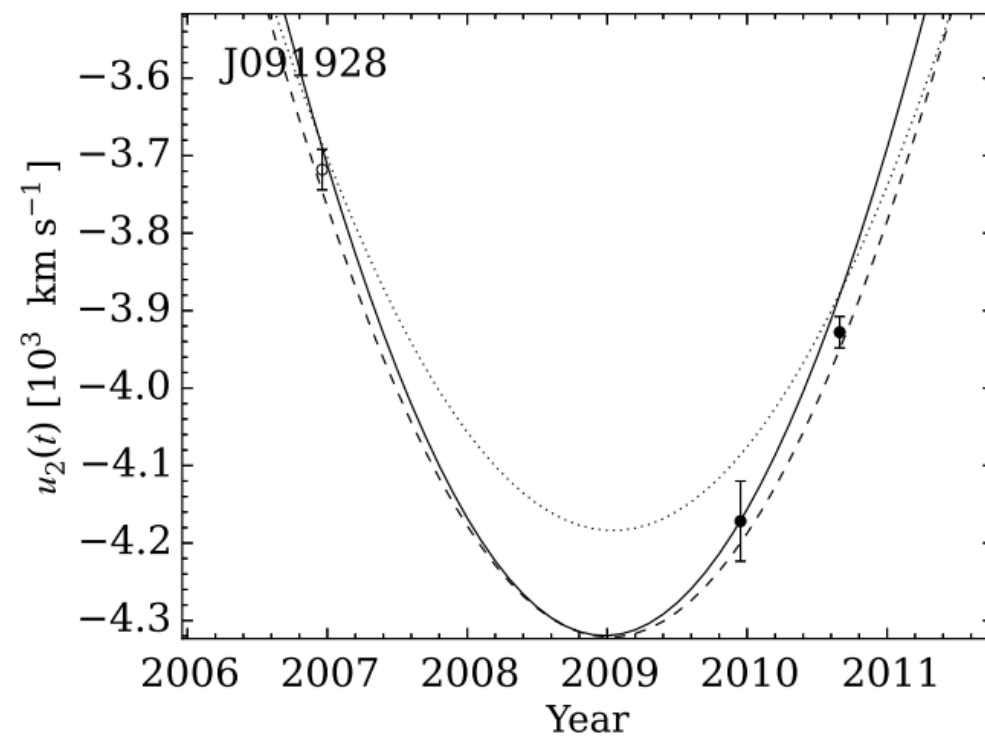
# Pan-Starrs PS0 J334.2028+01.4075

Periodicity caused by  $542 \pm 15$  day orbit of a  $10^{10}$  solar mass binary at  $0.05 < q < 0.25$  @  $z=2.06$  — separation of  $\sim 10 R_s$ !!





Looking for the radial motion of the spectral lines in quasars, there are ~3 good BBH candidates



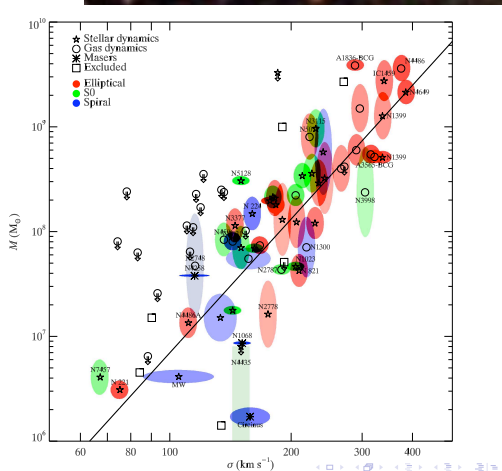
Stay tuned: By 2028,  
LSST should find  $\sim 10^4$   
BBH candidates

# Rule-breaker: Unassuming galaxy with 17 billion solar mass black hole!



van den Bosch et al. 2012

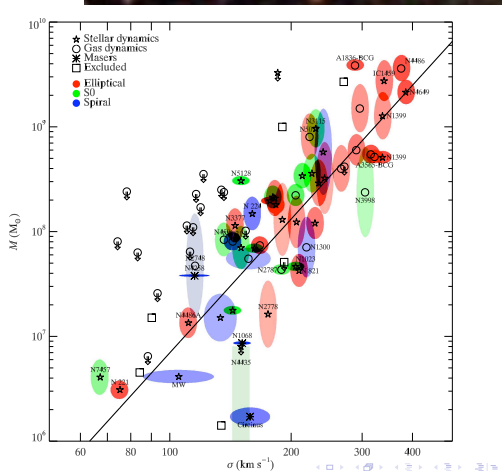
Perseus cluster





# Rule-breaker: Unassuming galaxy with ~~17~~ billion solar mass black hole!

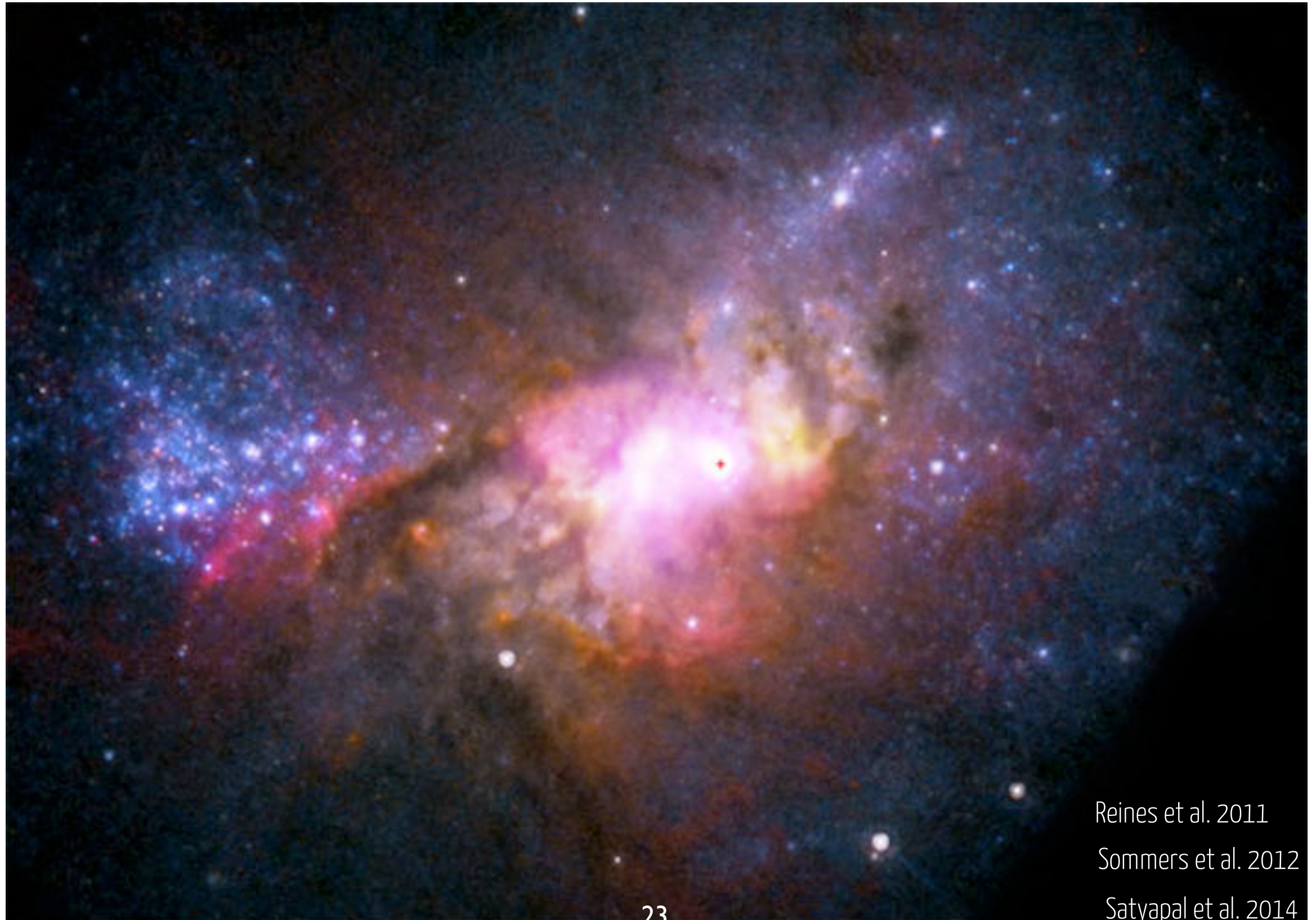
Stay tuned! MASSIVE Sur





# Heinze 2-10 is dwarf with a million solar mass black hole

and there are SMBHs in bulgeless galaxies, too!



Reines et al. 2011

Sommers et al. 2012

Satyapal et al. 2014



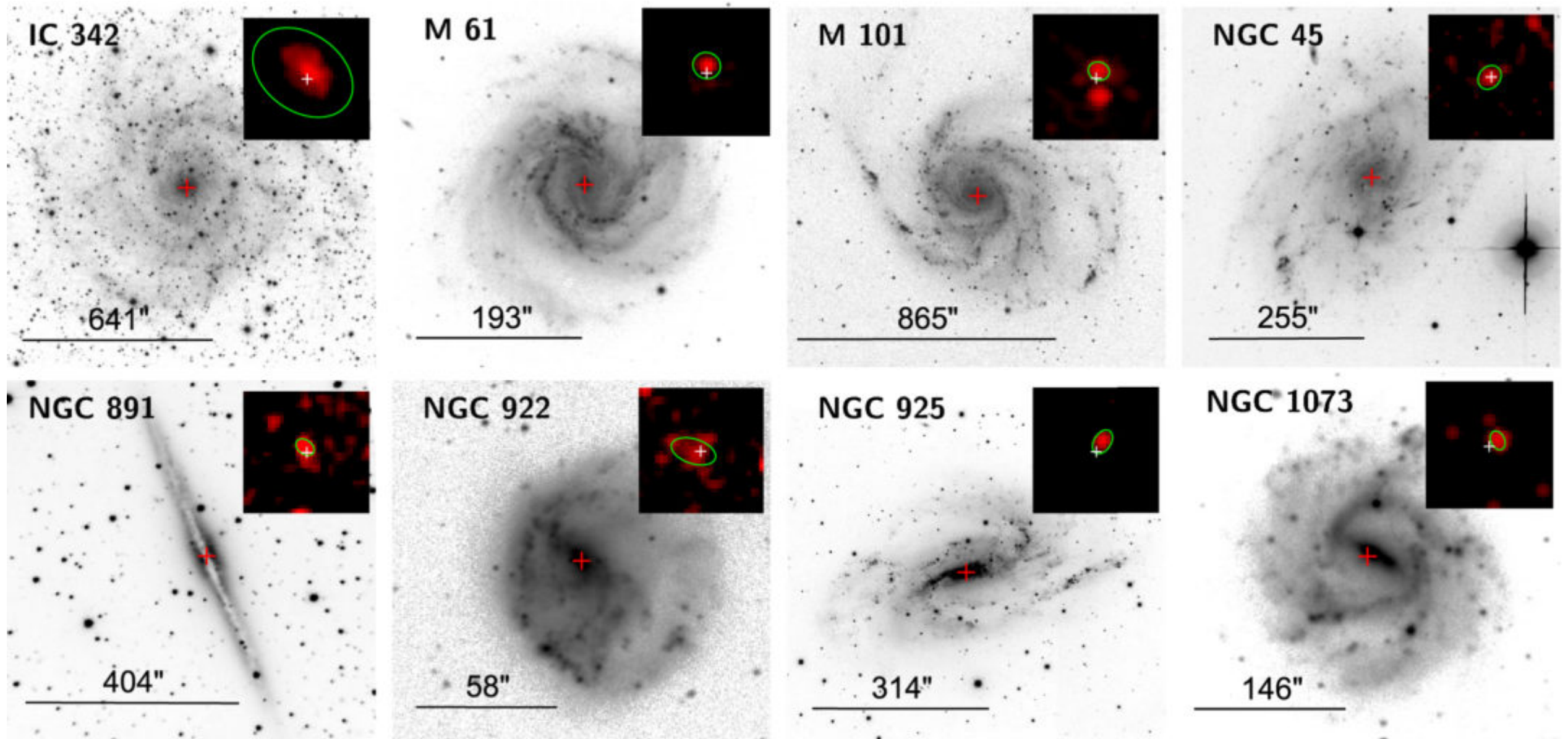
...and in low surface brightness galaxies, like Malin 1...



Warning: viral masses — assume line  
width maps to velocity for Keplerian  
motion



# Chandra reveals new SMBHs with $<10^6$ solar masses in disk galaxies



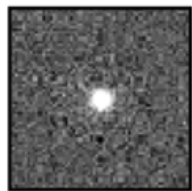
She et al 2017 — 21% of disk galaxies host SMBHs like these.



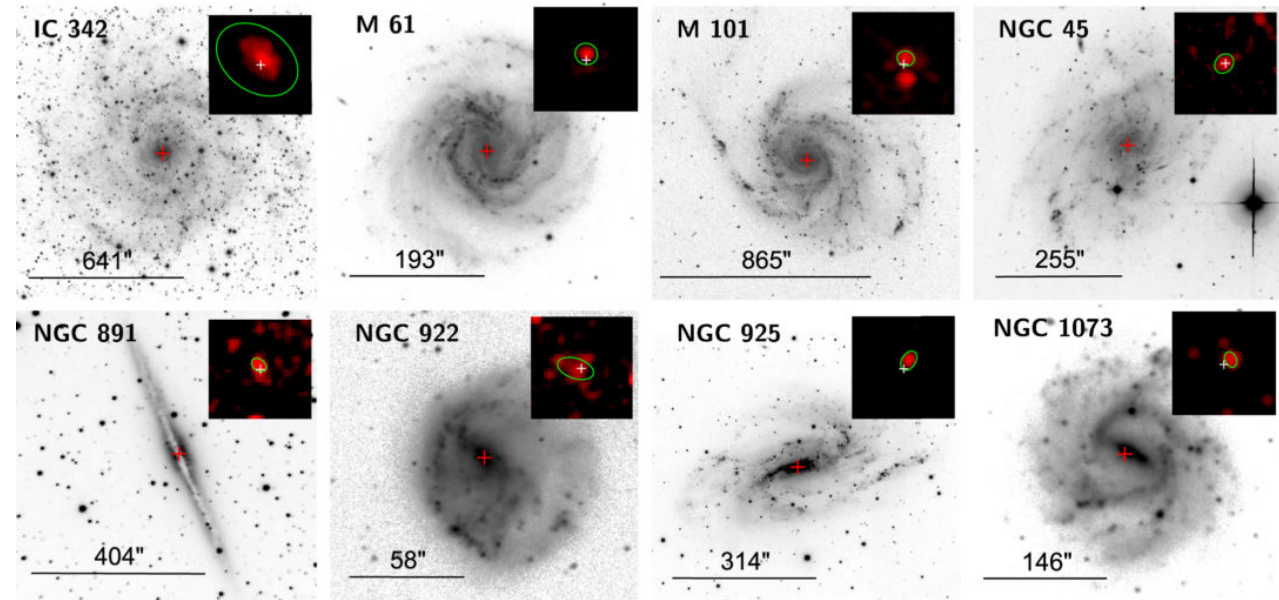
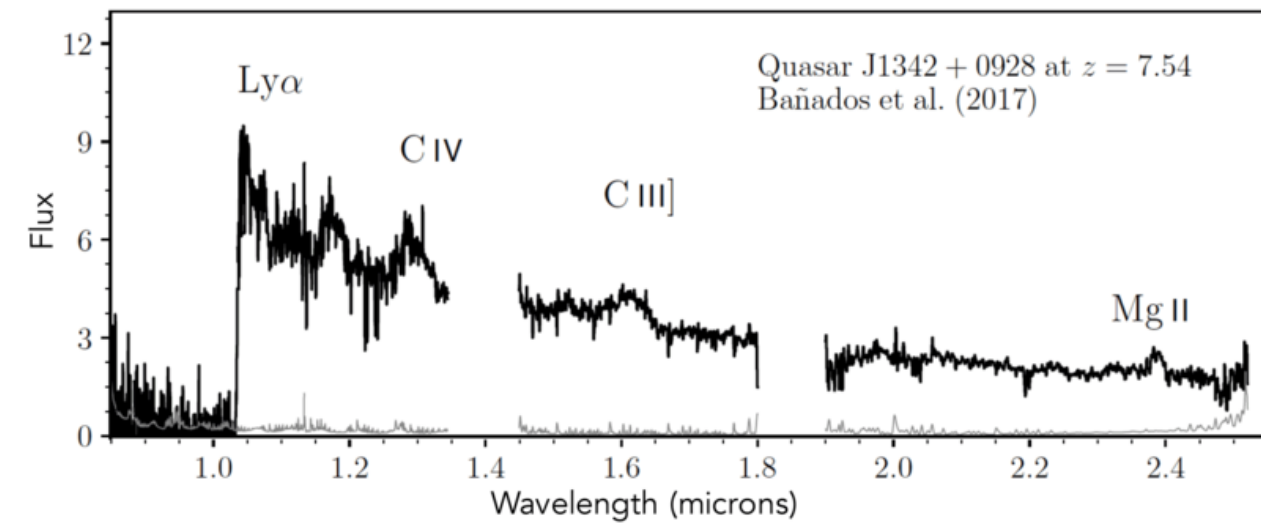
DECaLS z-band



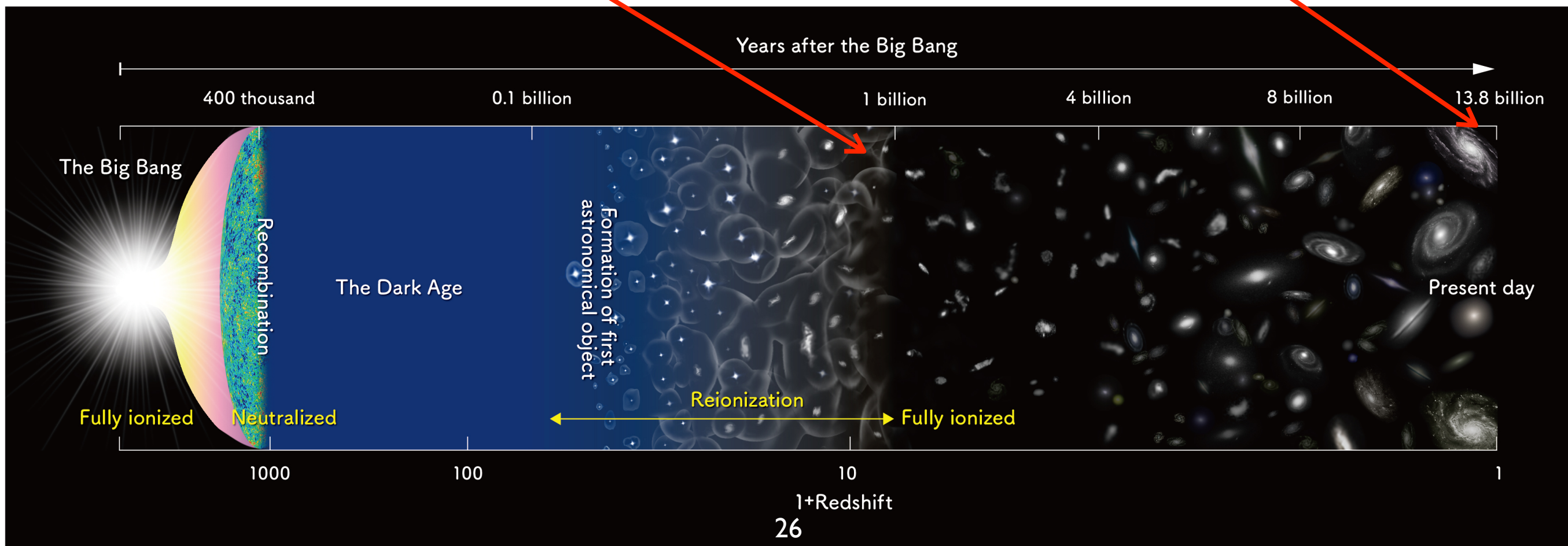
Magellan J-band



$\sim 10^6 - 10^9 M_{\odot}$  Black Holes



$8 \times 10^8 M_{\odot}$  Black Hole!

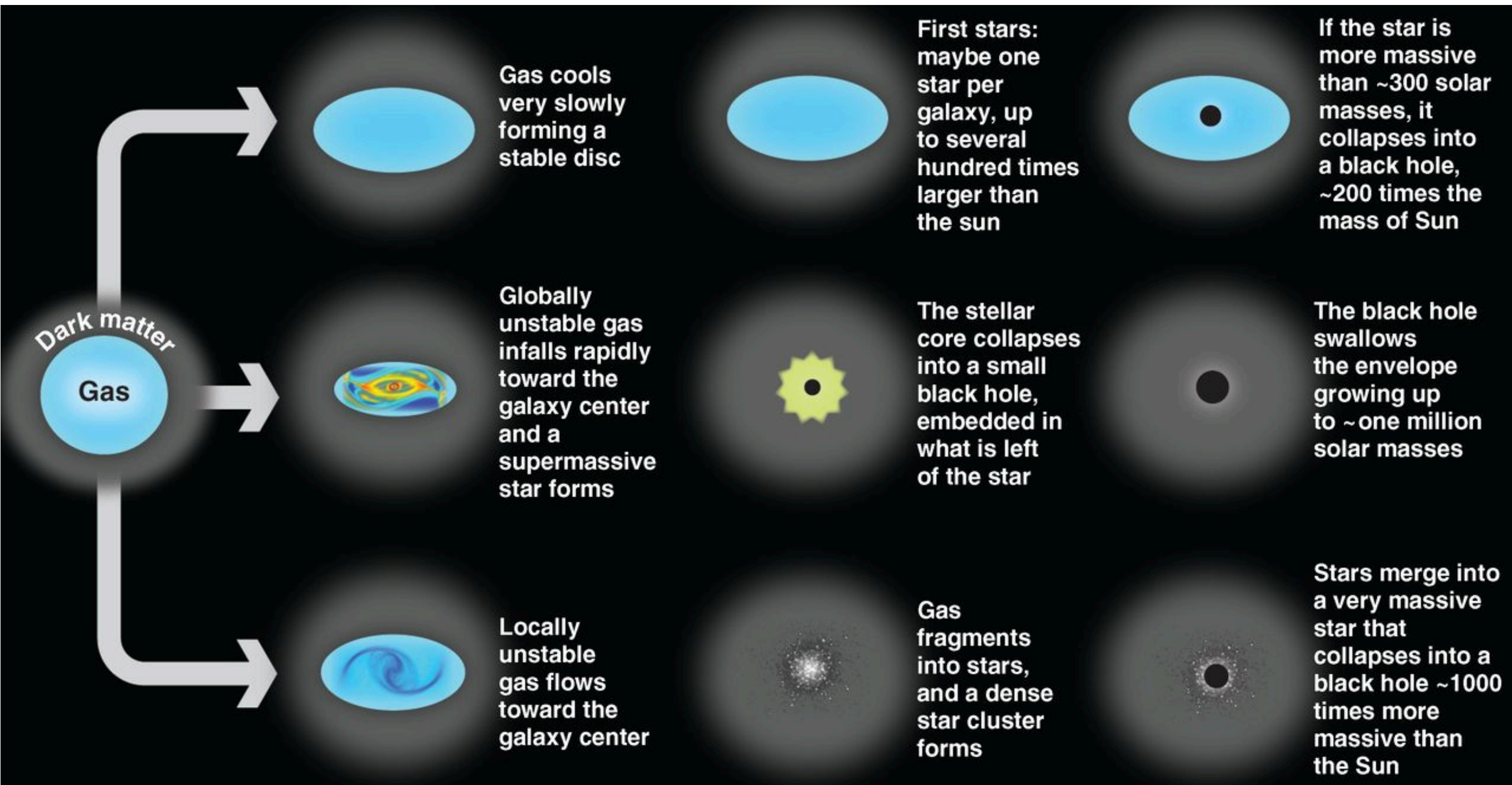




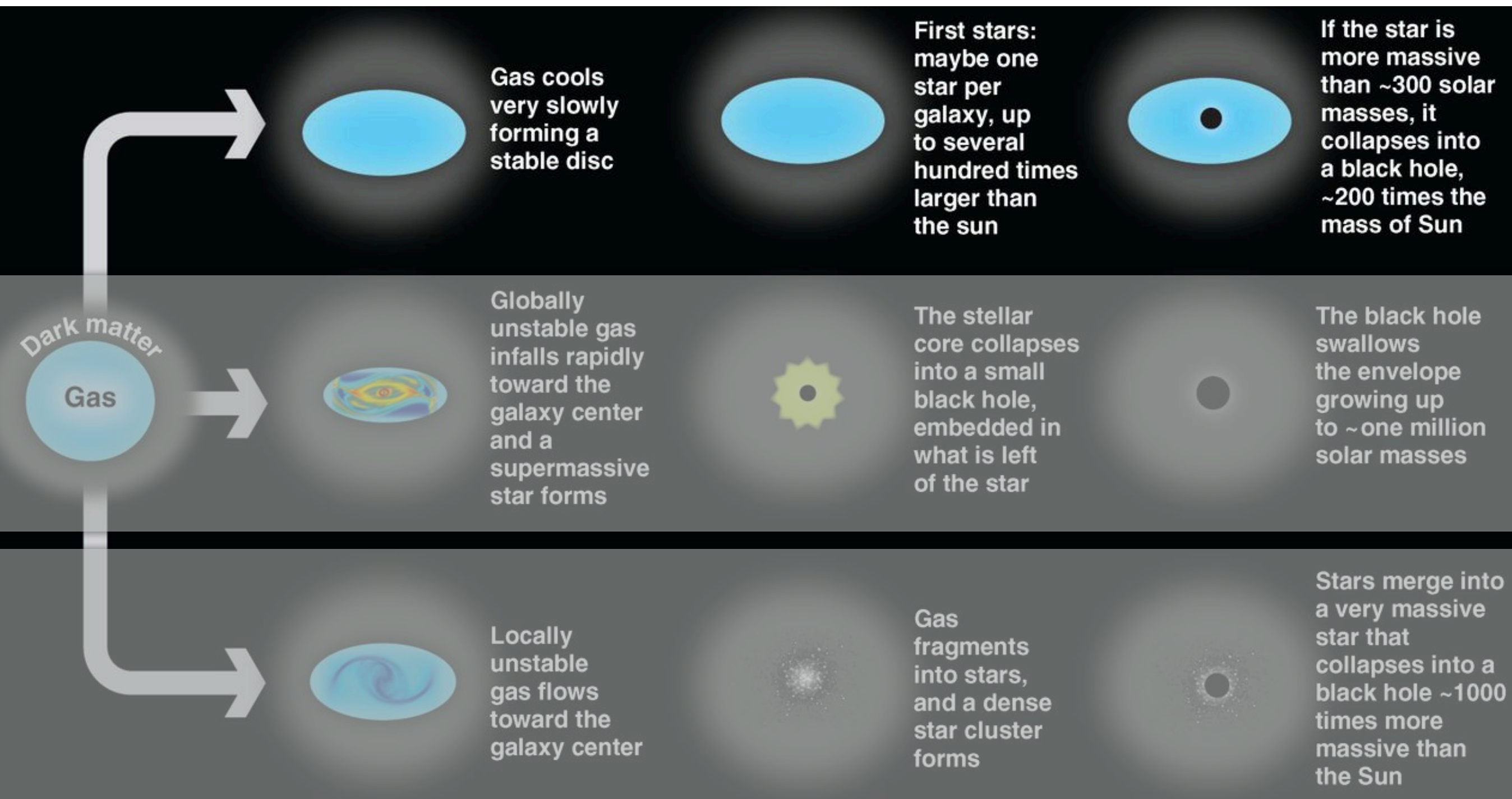
**How are supermassive black holes born  
and how can gigantic ones be in place so  
soon after the Big Bang?**



# Forming a black hole: let me count (some of) the ways



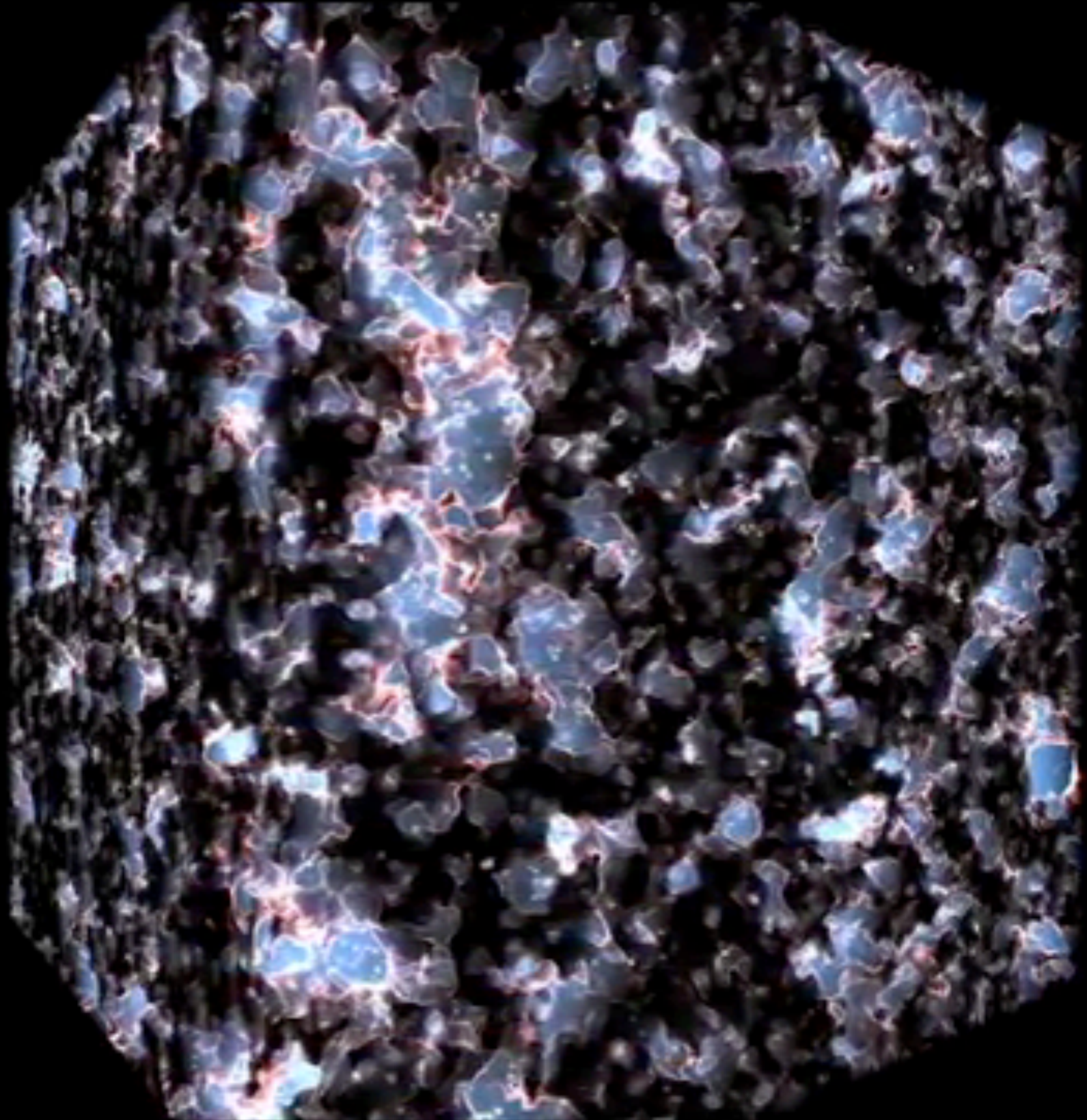
# One channel: Light seeds from the first generation of stars



From excellent Volonteri review

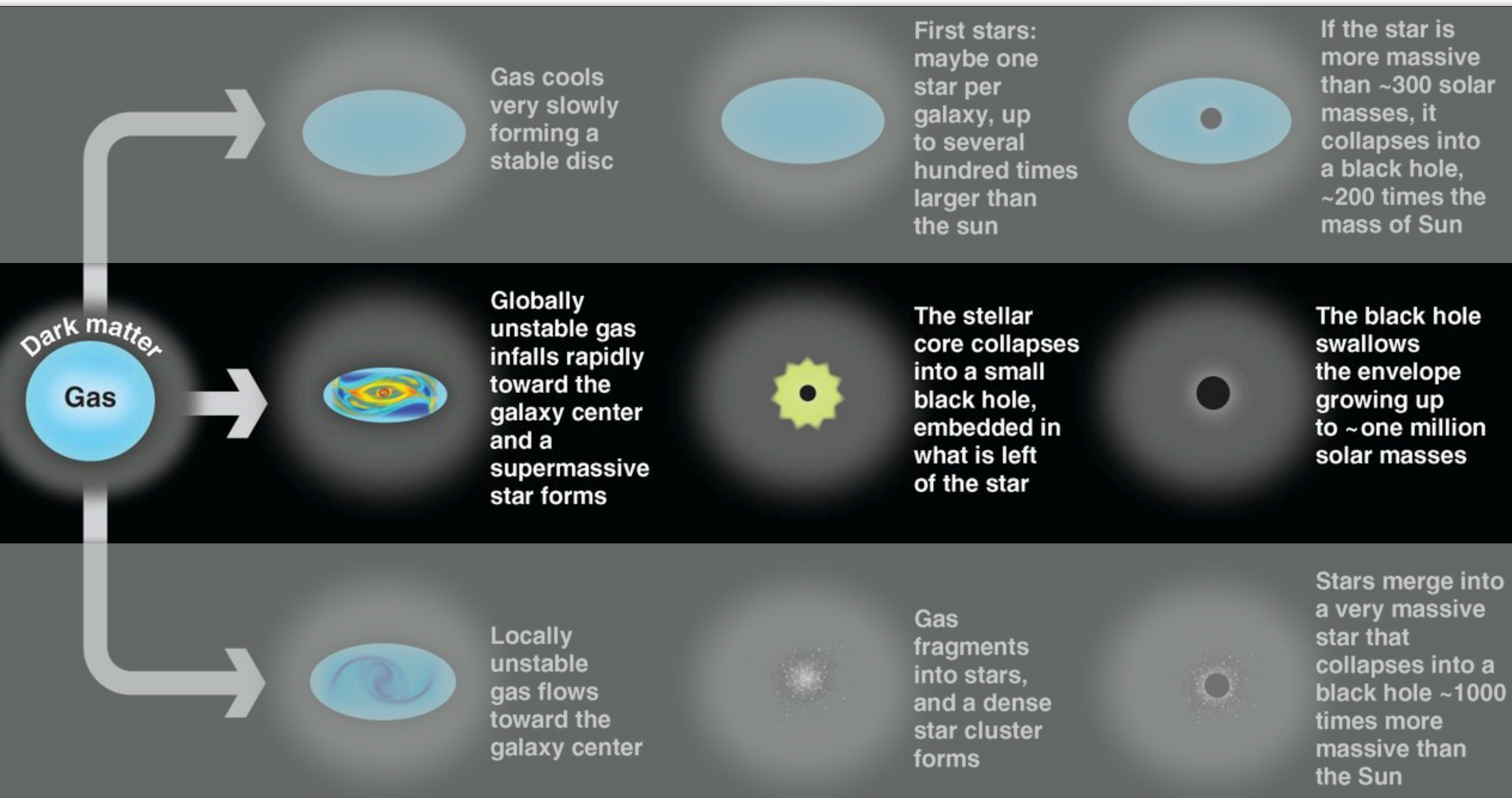


These first stars help heat and reionize the universe. The remnants are  $10\text{-}100M_{\odot}$  seed black holes



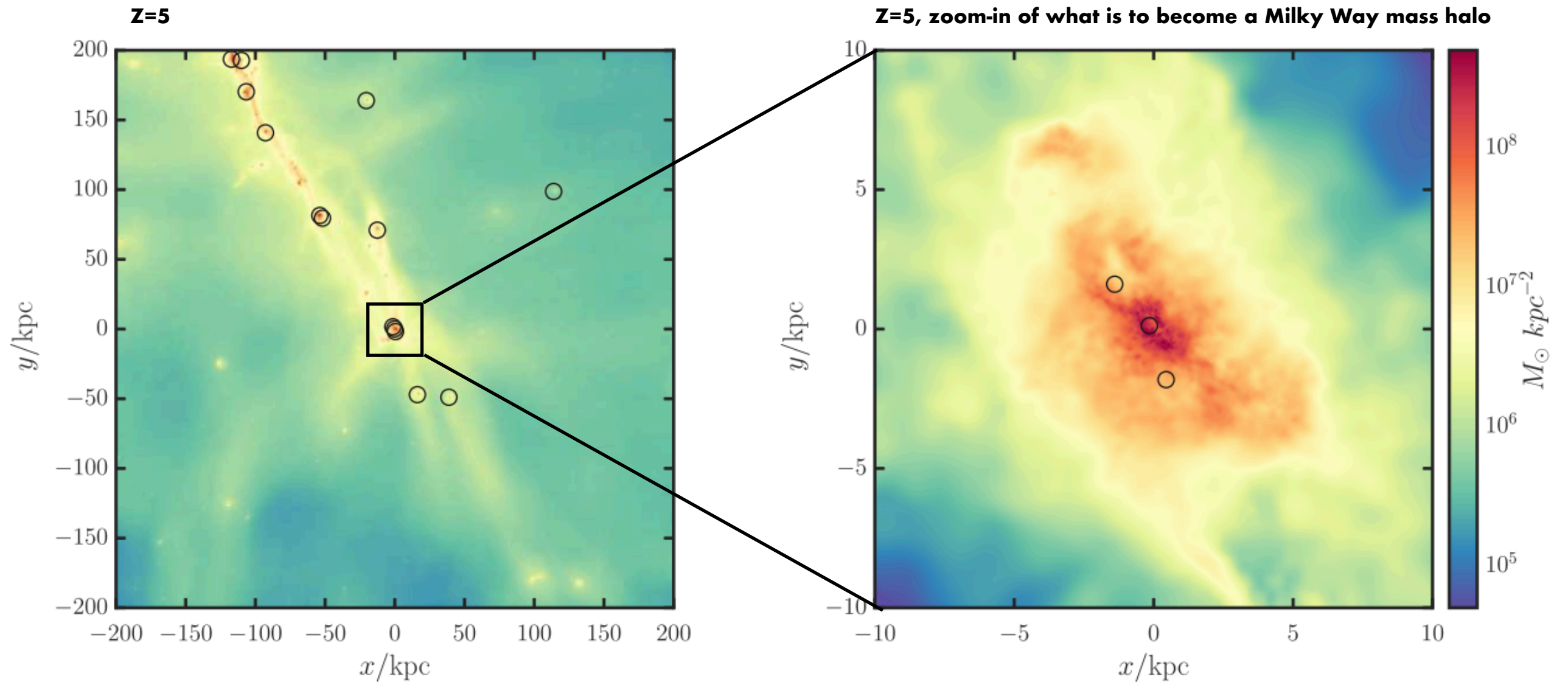


# One channel: Heavy seeds from directly collapsing black holes



From excellent Volonteri review



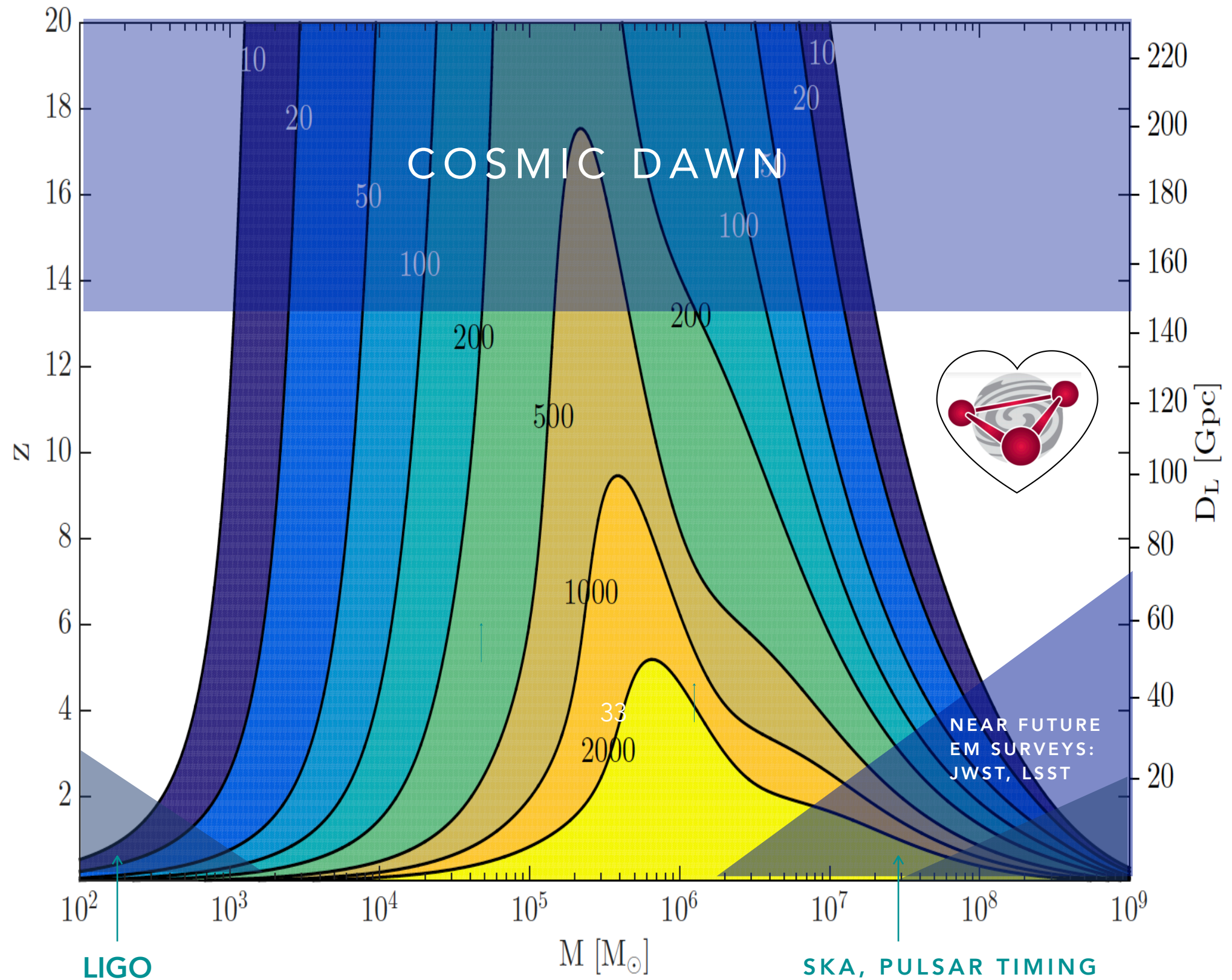


# Cosmological Hydrodynamical Simulations of Direct Collapse Black Hole Formation

Dunn, Bellovary, KHB, Christensen, Quinn 2018



Gravitational Waves will get the only\* direct view of seed formation and black hole early growth!



# It's a wonderful era to be an astronomer!

We need to get robust SMBH masses and pin down SMBH binaries

We need to know the real SMBH-galaxy correlation

We don't know how black holes are born or how they grow

We don't understand SMBH accretion and feedback

We need to understand SMBH dynamics

Spin! We aren't thinking enough about spin!